

# Review questions for Solar System and the Universe answer

- 1) distance between Earth and Sun.  
 $1.5 \times 10^{11}$  meters
- 2) distance light travels in 1 year  
1 parsec = 3.3 ly
- 3) "Sun-centered universe" copernicus
- 4) "Earth-centered universe" aristotle
- 5) spots on Sun, craters on the Moon, 4 moons orbiting Jupiter
- 6) elliptical
- 7) 13.6 BYA
- 8) 5 BYA
- 9) 1) nebula 2) protostar 3) star  
4) brown dwarf 5) Main Sequence 6) red giant  
7) white dwarf 8) giant  
main sequence 9) super giant 10) supernova  
11) neutron star 12) black hole
- 10) Mercury = Me Venus = V Earth = E  
Mars = M Jupiter = J  
Saturn = S Uranus = U  
Neptune = N

Me	V	E	M	J	S	U	N	
Me	V	E	M	J	S	U	N	
Me	V	E	M	M	J	S	U	N

11) optional: spherical, in orbit around Sun, “cleared neighborhood of debris”

12) a)  $T^2 = R^3$     b)  $F_g = \frac{G M_1 M_2}{r^2}$     c)  $r_s = \frac{2GM}{c^2}$   
time yrs  
dist AU

13) XXXXX

14) universe originally very small point that exploded outward 13.6 BYA

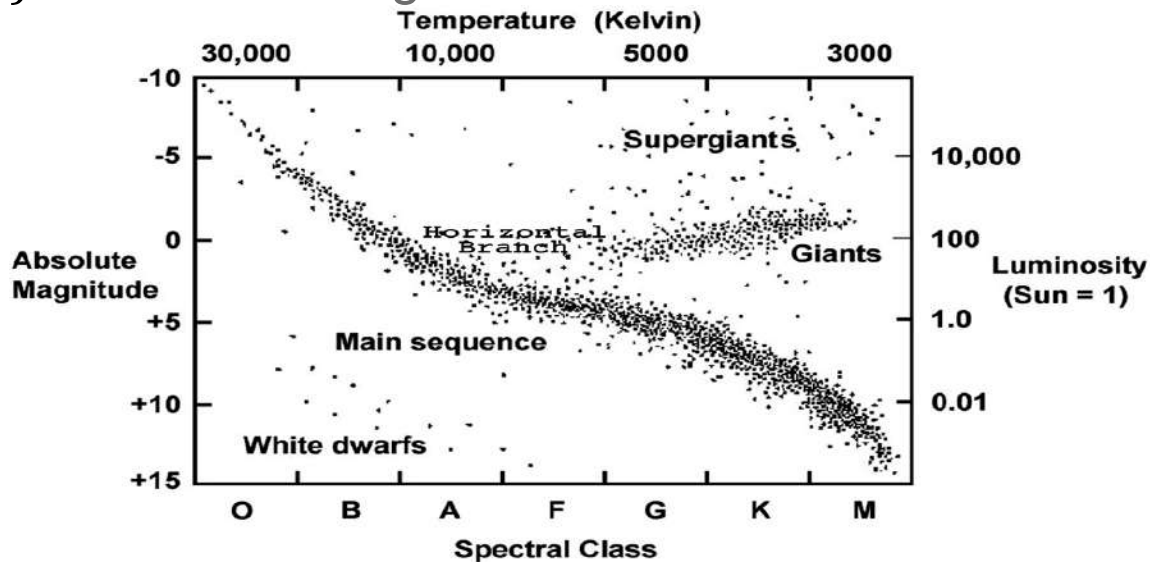
15) size [mass]

16) and temperature [color]

17) analyze small sample of data and apply to entire population

18)  $\frac{1}{2}$  angle star “moves” ; closer objects appear to move further than distance objects.

19) The H-R Diagram



20) Red , blue

21) Sun

22) Satisfy relationship of luminosity and mass

Relationship given by  $L = L_0 [M/M_0]^{3.5}$

23) Fusing hydrogen into helium in core; stable gravity balances pressure keep stars stable, when pressure increases, star expands and cools, when pressure decreases, stars collapse from gravity

24) spiral, elliptical, irregular

25) dwarf elliptical

26) spiral, supermassive black hole at center

27) supermassive star collapsed at end of life, so dense light cannot escape.

28) Yes and accelerating apart

29) As stars move away, wavelength appears longer; as stars approach, wavelength appears shorter

30)  $F_g = [G M_1 M_2] / r^2$   $F_g = \frac{[6.67 \times 10^{-11}] [5.98 \times 10^{24}] [1.99 \times 10^{30}]}{[3 \times 10^8][3 \times 10^8]}$

$F_g = 3.53 \times 10^{22} \text{ N}$

31)  $T^2 = R^3$   $T^2 = [6]^3$   $T = 14.7 \text{ years}$

third page of questions:

Question 20	true
Question 21	false
Question 22	true
Question 23	true
Question 24	true

Question 25      false  
 Question 26      C  
 Question 27      A  
 Question 28      D  
 Question 29      A  
 Question 30      C  
 Question 31      horizontal axis: absolute temperature    AT  
                          vertical axis relative luminosity RL

diagonal line      MSS      oval that contains B:    SGS      oval that contains D: WD  
                          other oval :GS

32: use     $F_{\text{grav}}$     mass of sun, mass of Neptune    distance  $r$  in meters

$$F_{\text{grav}} = \frac{[6.67 \times 10^{-11}][1.99 \times 10^{30}][1.02 \times 10^{26}]}{[4.5 \times 10^{12}]^2} = \frac{13.54 \times 10^{45}}{20.25 \times 10^{24}}$$

$$F_{\text{grav}} = 0.669 \times 10^{21} = 6.69 \times 10^{20} \text{ N}$$

33. use     $T^2 = R^3$       use distance in AU for Neptune; answer in years

$$T^2 = [30][30][30] = 27000$$

$$T = 164.3 \text{ years}$$

34 use  $r_s = 2MG/c^2$      $c = 3 \times 10^8$

$$r_s = \frac{2[1.02 \times 10^{26}][6.67 \times 10^{-11}]}{[3 \times 10^8]^2} = \frac{13.61 \times 10^{15}}{9 \times 10^{16}} = 1.51 \times 10^{-1} = 0.151 \text{ m}$$