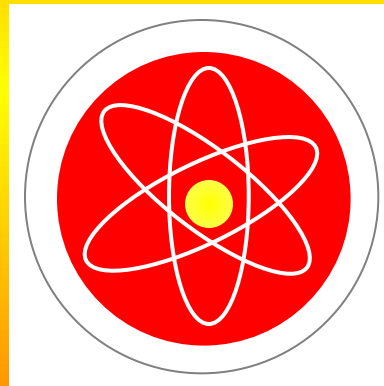


# Nuclear



# Guiding Questions

Is radiation dangerous?

Is nuclear power a good choice?



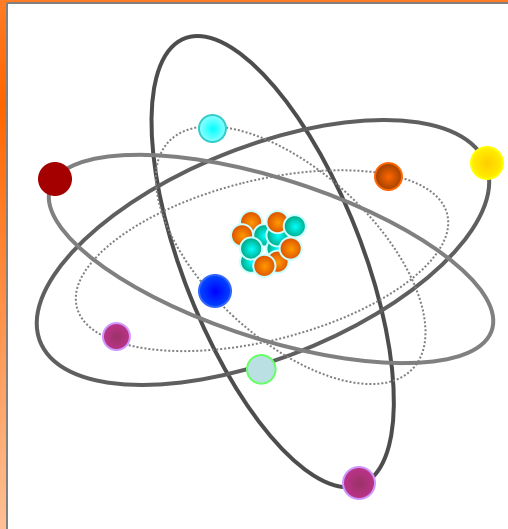
What is nuclear energy?

Are nuclear energy and nuclear bombs both dangerous?



**Bravo - 15,000 kilotons**

# Development of the Atom



# Nuclear



Review Background

Nuclear Radiation

Fission

Nuclear Power Plants

Half-Life

Decay Series

Fusion

# Key Terms

alpha decay

alpha particles

artificial transmutation

background radiation

beta decay

beta particle

chain reaction

control rods

critical mass

curie

disintegrations per second

gamma decay

Geiger counter

half-life

ionizing radiation

irradiate

isotope

moderator

natural radioactivity

nuclear equation

nuclear fission

nuclear fusion

nuclide

plasma

positrons

rad

radioisotope

rem

roentgen

tracers

transmutation

X-rays

# Radioactivity

Much of our understanding of atomic structure came from studies of radioactive elements.

## Radioactivity

The process by which atoms spontaneously emit high energy particles or rays from their nucleus.

First observed by  
Henri Becquerel in 1896

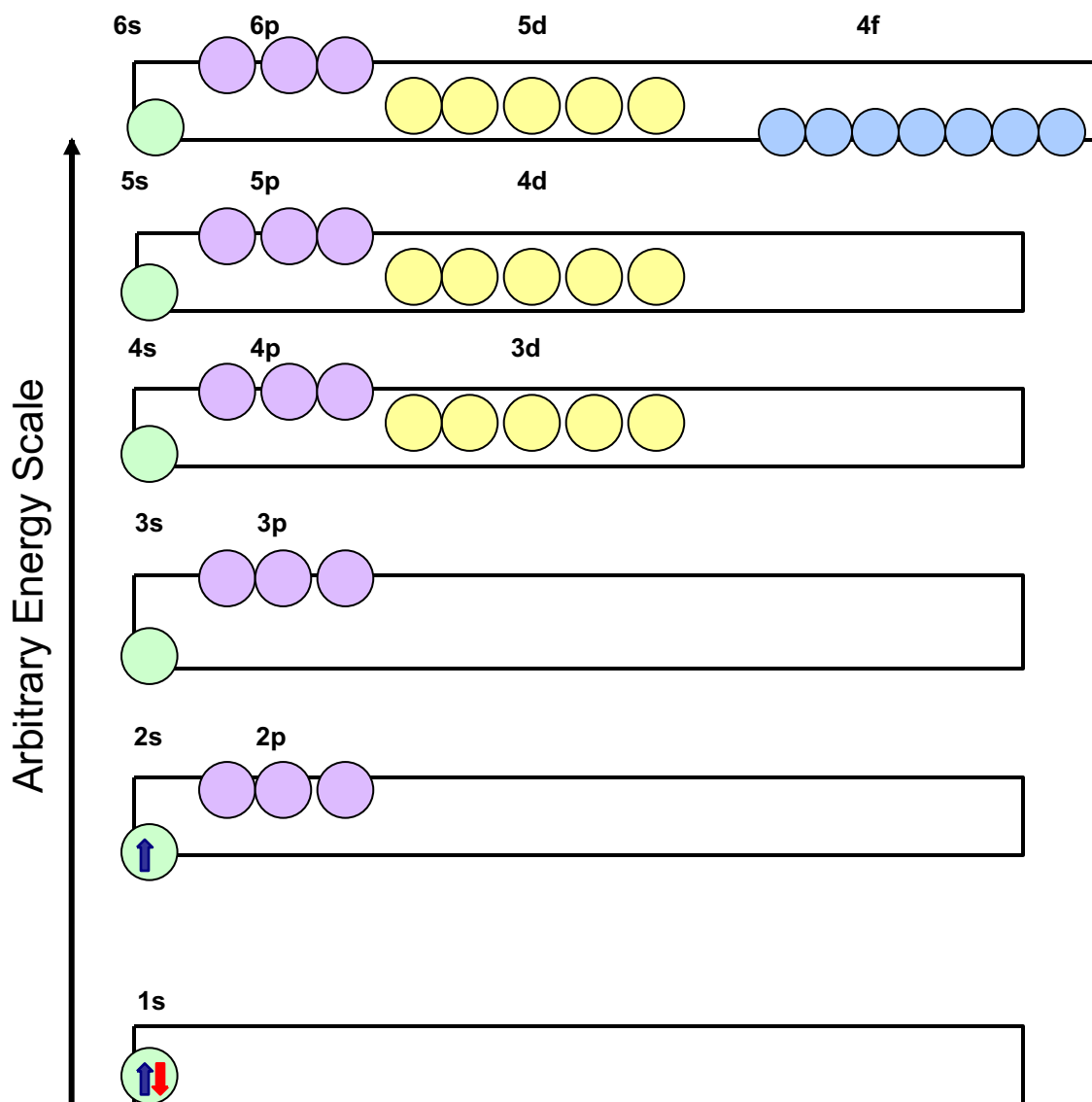


# History: On The Human Side

- 1834** *Michael Faraday* - electrolysis experiments suggested electrical nature of matter
- 1895** *Wilhelm Roentgen* - discovered X-rays when cathode rays strike anode
- 1896** *Henri Becquerel* - discovered "uranic rays" and radioactivity
- 1896** *Marie (Marya Skłodowska) and Pierre Curie* - discovered that radiation is a property of the atom, and not due to chemical reaction. (Marie named this property *radioactivity*.)
- 1897** *Joseph J. Thomson* - discovered the electron through Crookes tube experiments
- 1898** *Marie and Piere Curie* - discovered the radioactive elements polonium and radium
- 1899** *Ernest Rutherford* - discovered alpha and beta particles
- 1900** *Paul Villard* - discovered gamma rays
- 1903** *Ernest Rutherford* and *Frederick Soddy* - established laws of radioactive decay and transformation
- 1910** *Frederick Soddy* - proposed the isotope concept to explain the existence of more than one atomic weight of radioelements
- 1911** *Ernest Rutherford* - used alpha particles to explore gold foil; discovered the nucleus and the proton; proposed the nuclear theory of the atom
- 1919** *Ernest Rutherford* - announced the first artificial transmutation of atoms
- 1932** *James Chadwick* - discovered the neutron by alpha particle bombardment of Beryllium
- 1934** *Frederick Joliet* and *Irene Joliet Curie* - produced the first artificial radioisotope
- 1938** *Otto Hahn, Fritz Strassmann, Lise Meitner, and Otto Frisch* - discovered nuclear fission of uranium-235 by neutron bombardment
- 1940** *Edwin M McMillan* and *Philip Abelson* - discovered the first transuranium element, neptunium, by neutron irradiation of uranium in a cyclotron
- 1941** *Glenn T. Seaborg, Edwin M. McMillan, Joseph W. Kennedy* and *Arthur C. Wahl* - announced discovery of plutonium from beta particle emission of neptunium
- 1942** *Enrico Fermi* - produced the first nuclear fission chain-reaction
- 1944** *Glenn T. Seaborg* - proposed a new format for the periodic table to show that a new actinide series of 14 elements would fall below and be analogous to the 14 lanthanide-series elements.
- 1964** *Murray Gell-Mann* hypothesized that quarks are the fundamental particles that make up all known subatomic particles except leptons.

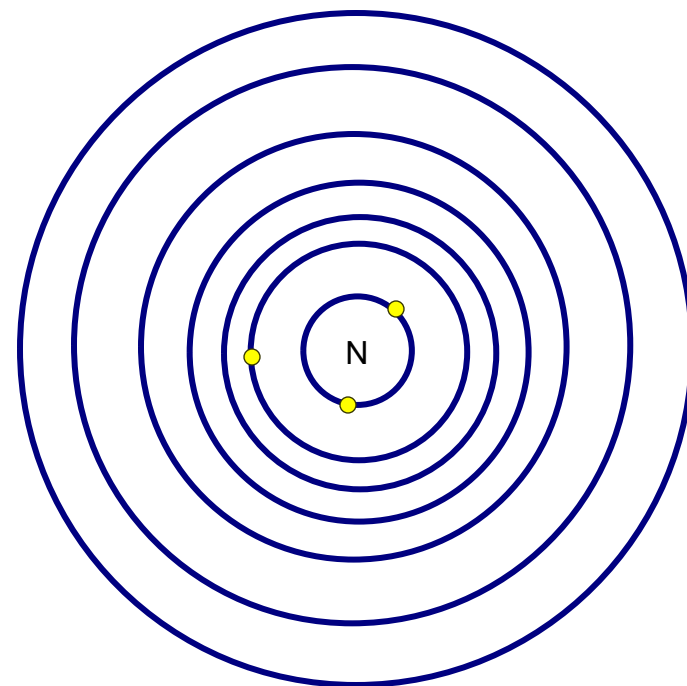


# Energy Level Diagram



Lithium

Bohr Model



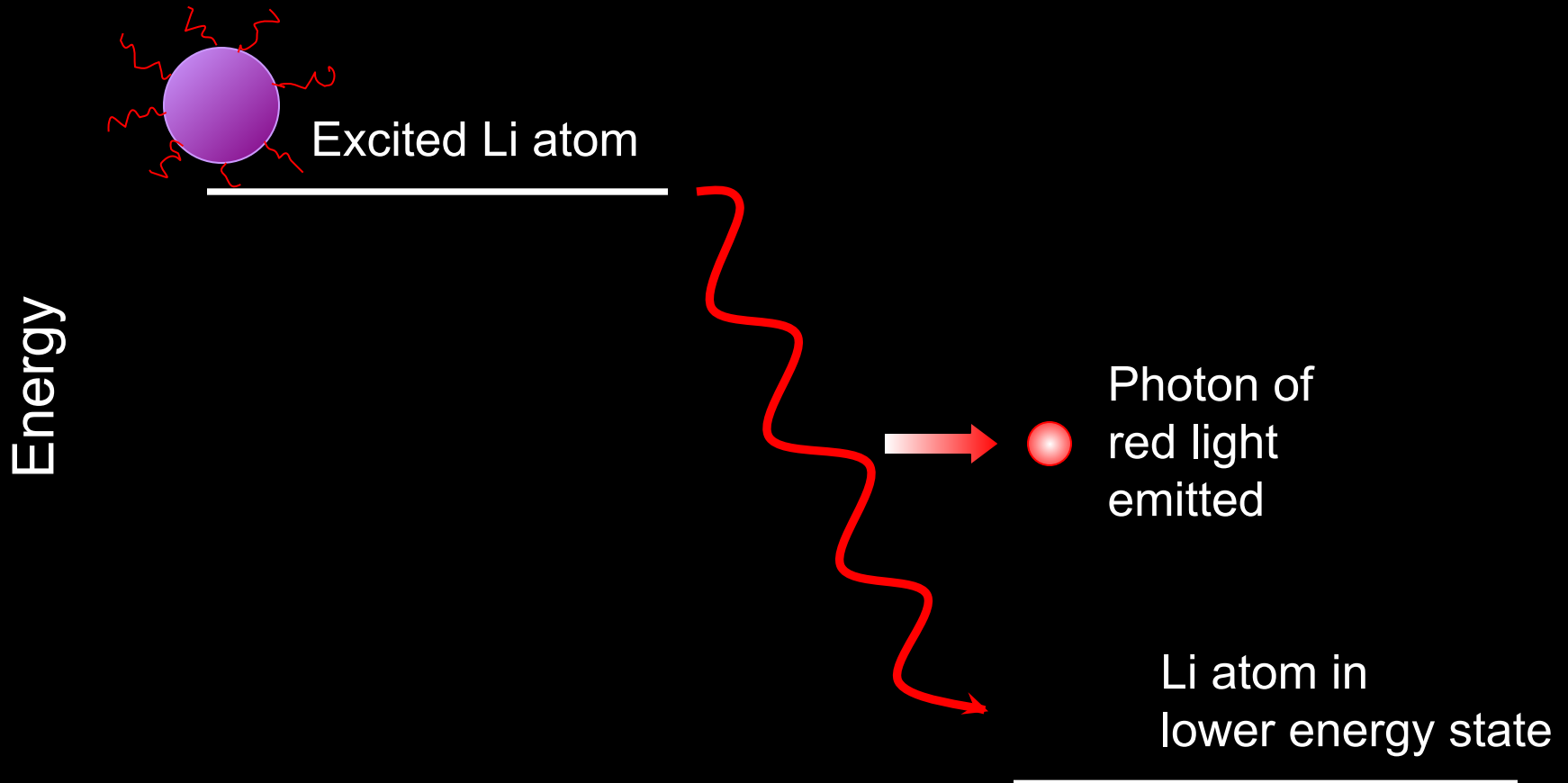
Electron Configuration



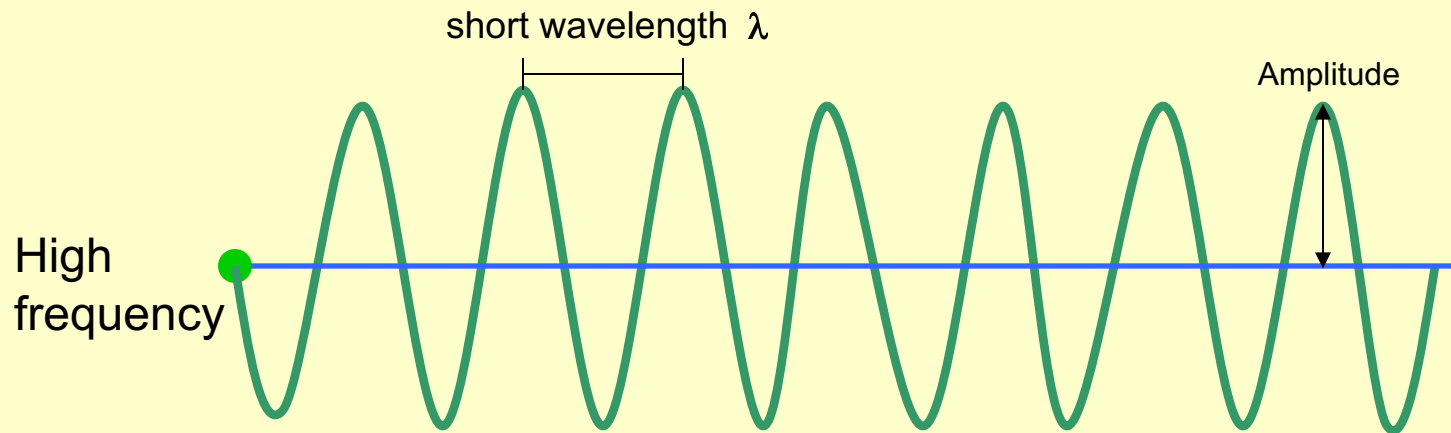
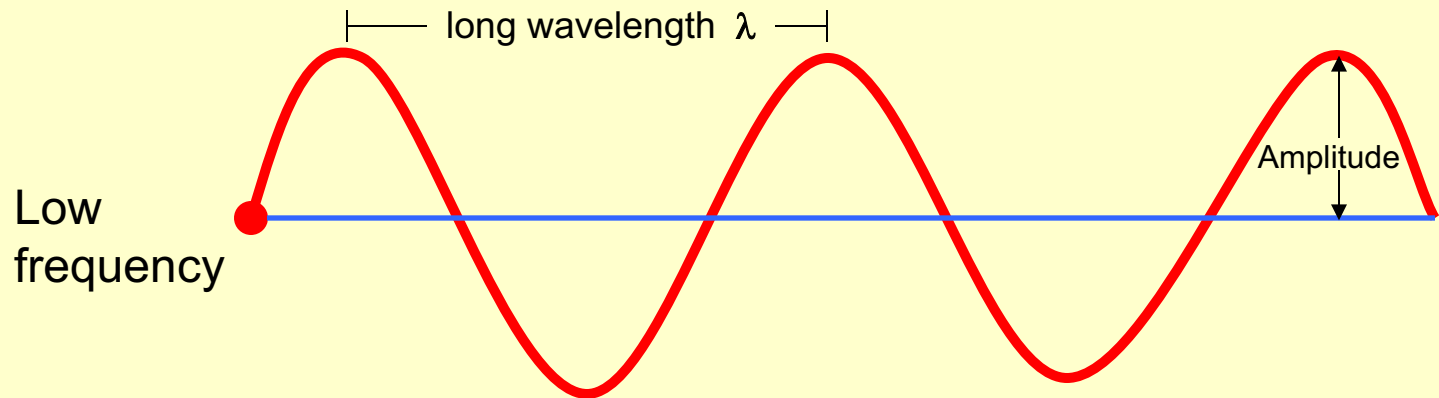
H He Li C N Al Ar F Fe La

CLICK ON ELEMENT TO FILL IN CHARTS

# An Excited Lithium Atom

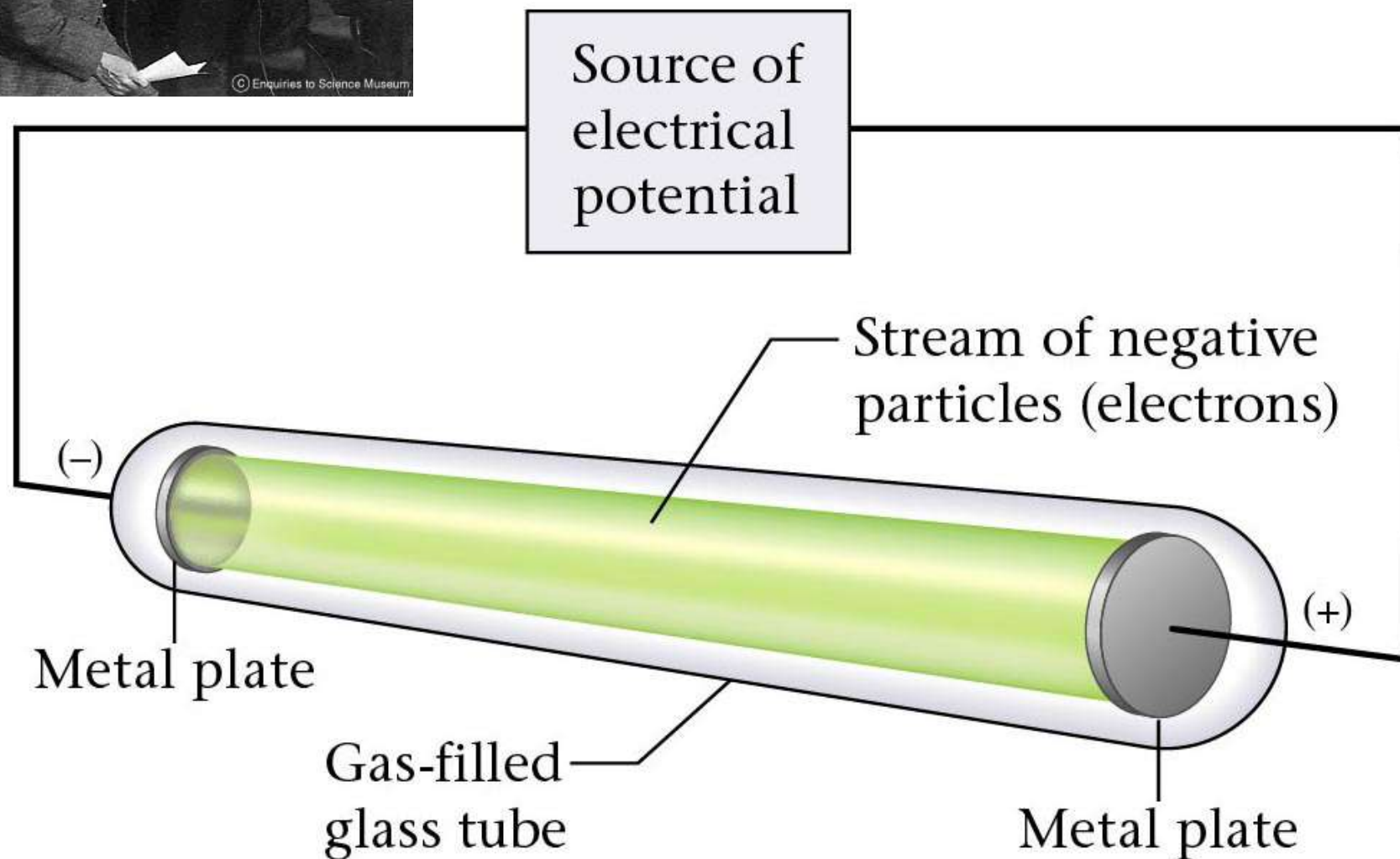


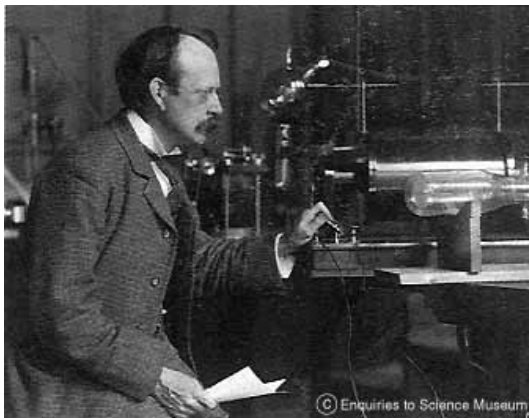
# Waves



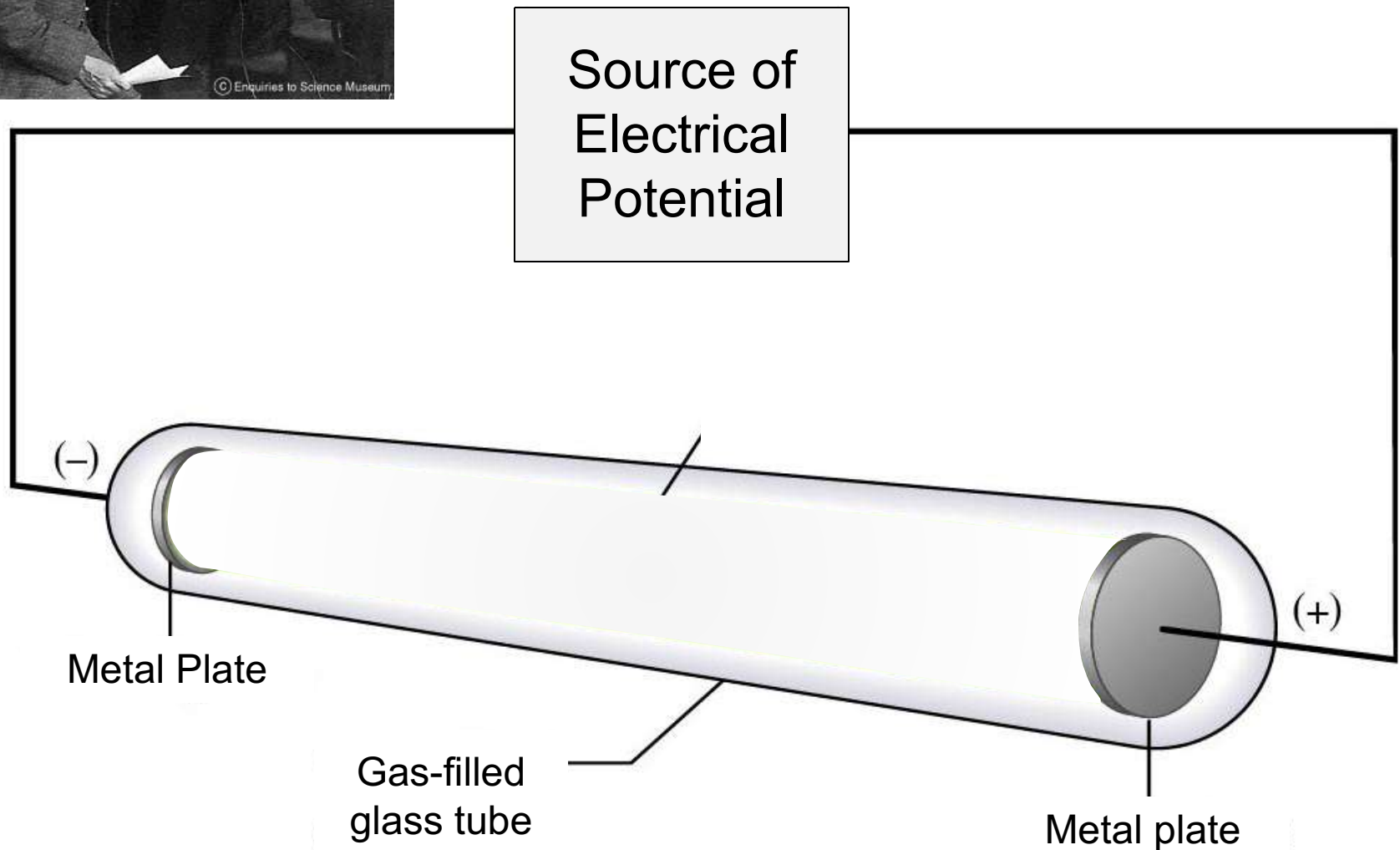


# A Cathode Ray Tube

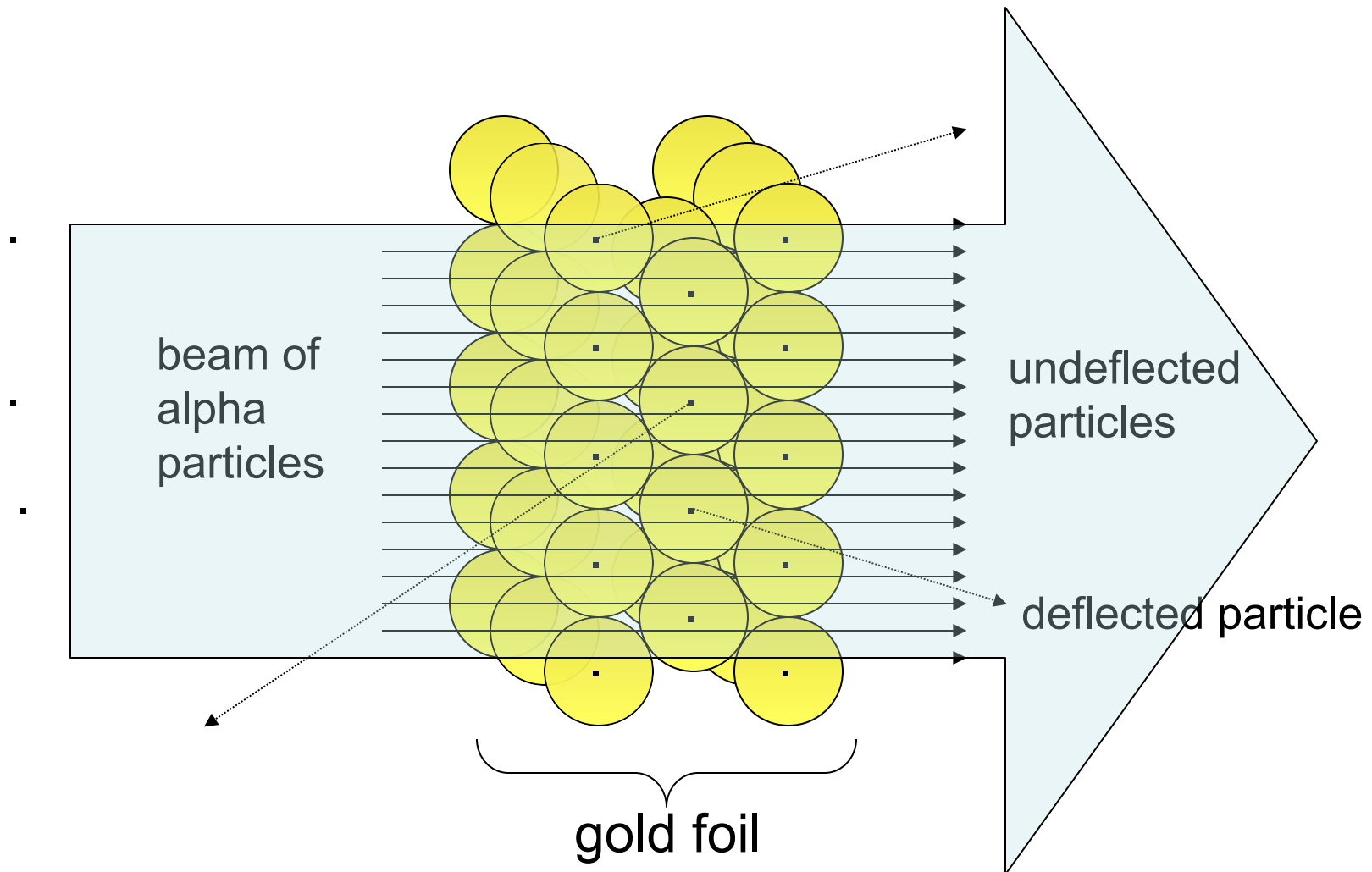




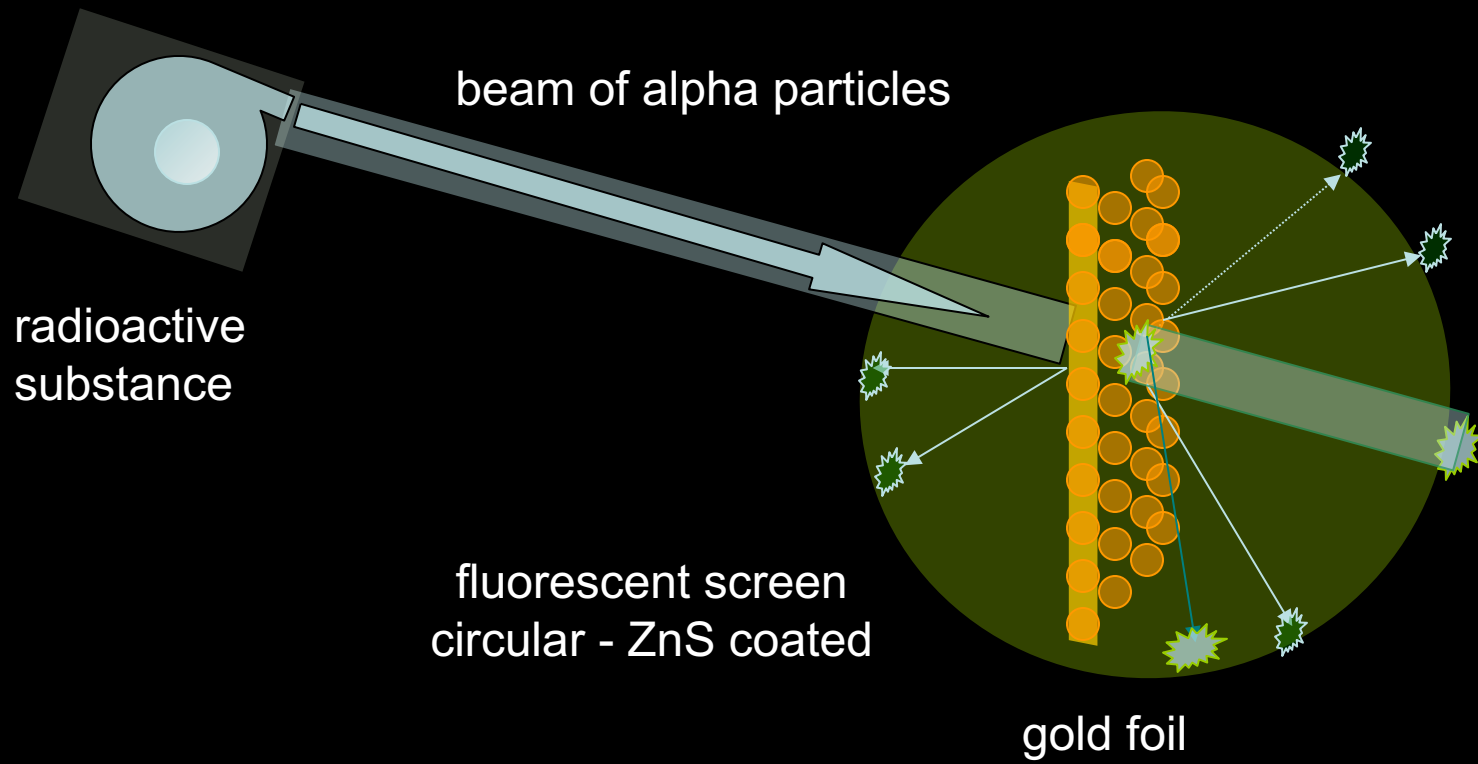
# A Cathode Ray Tube



# Interpreting the Observed Deflections



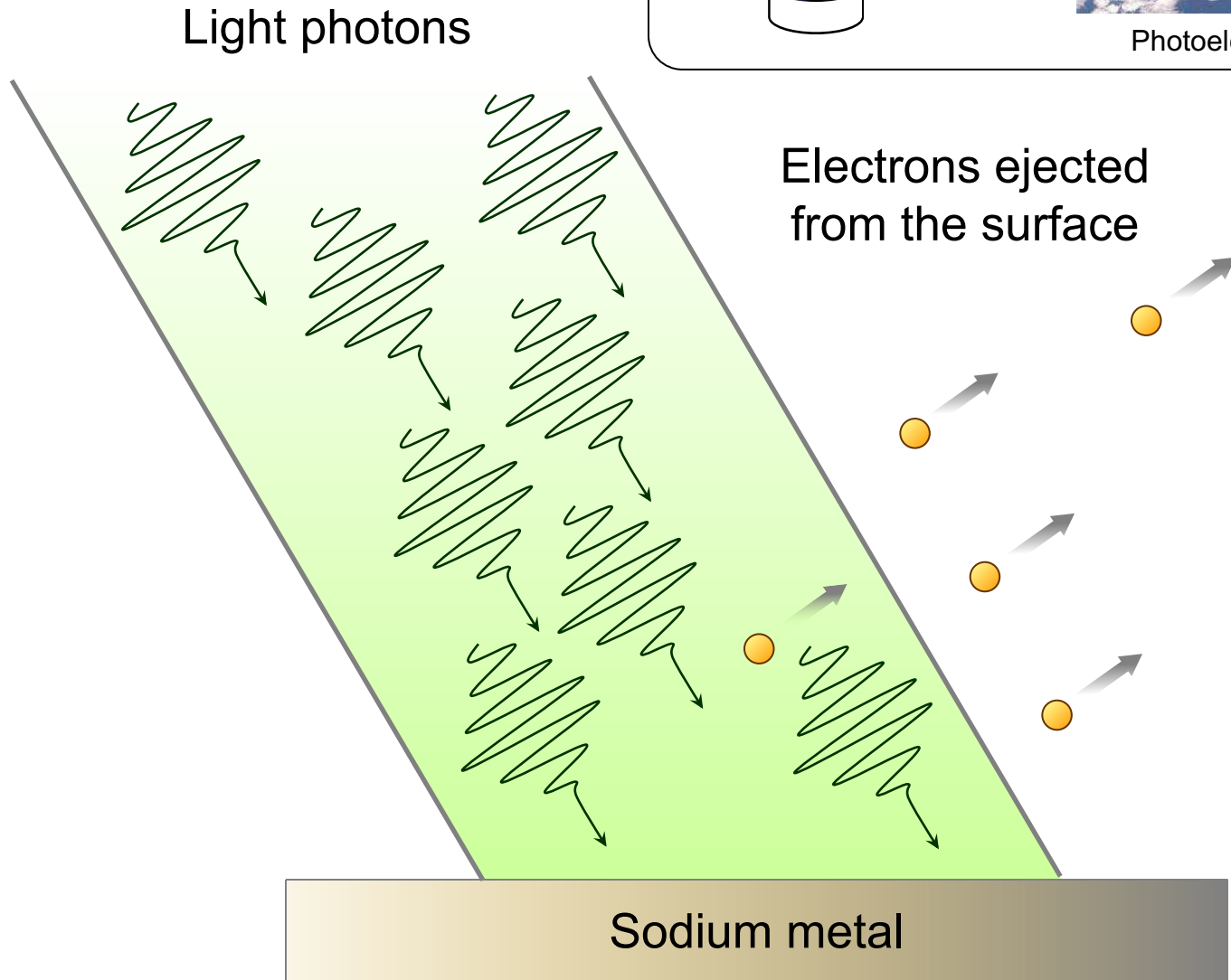
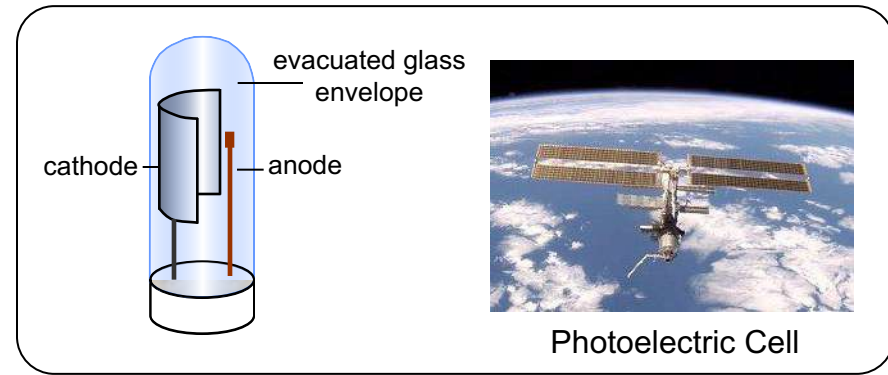
# Rutherford's Apparatus



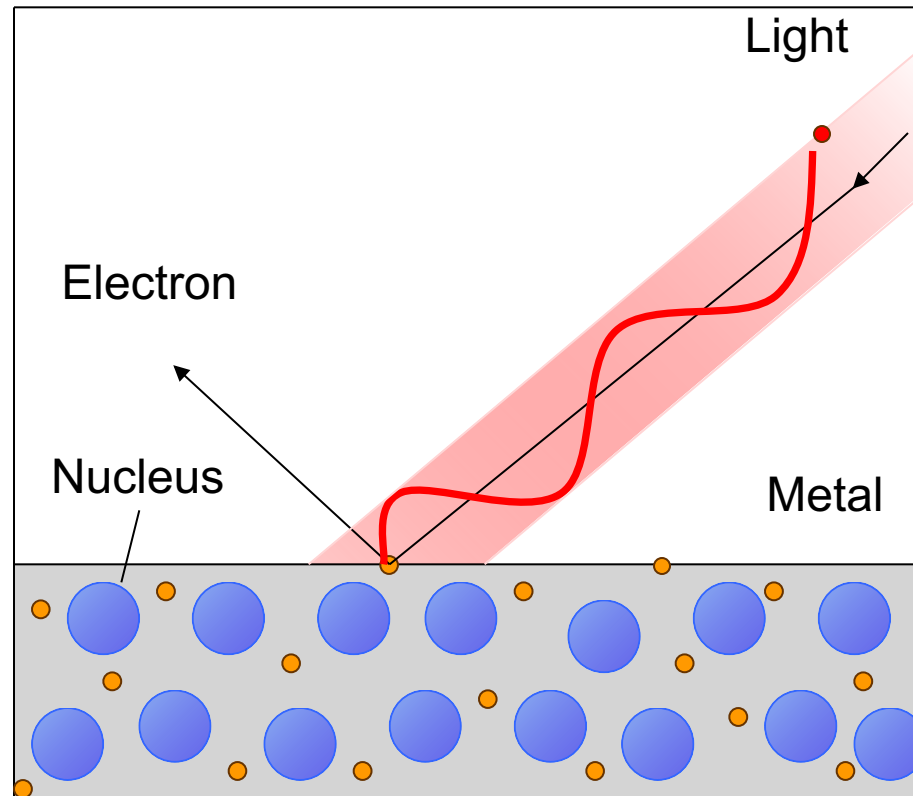




# Photoelectric Effect

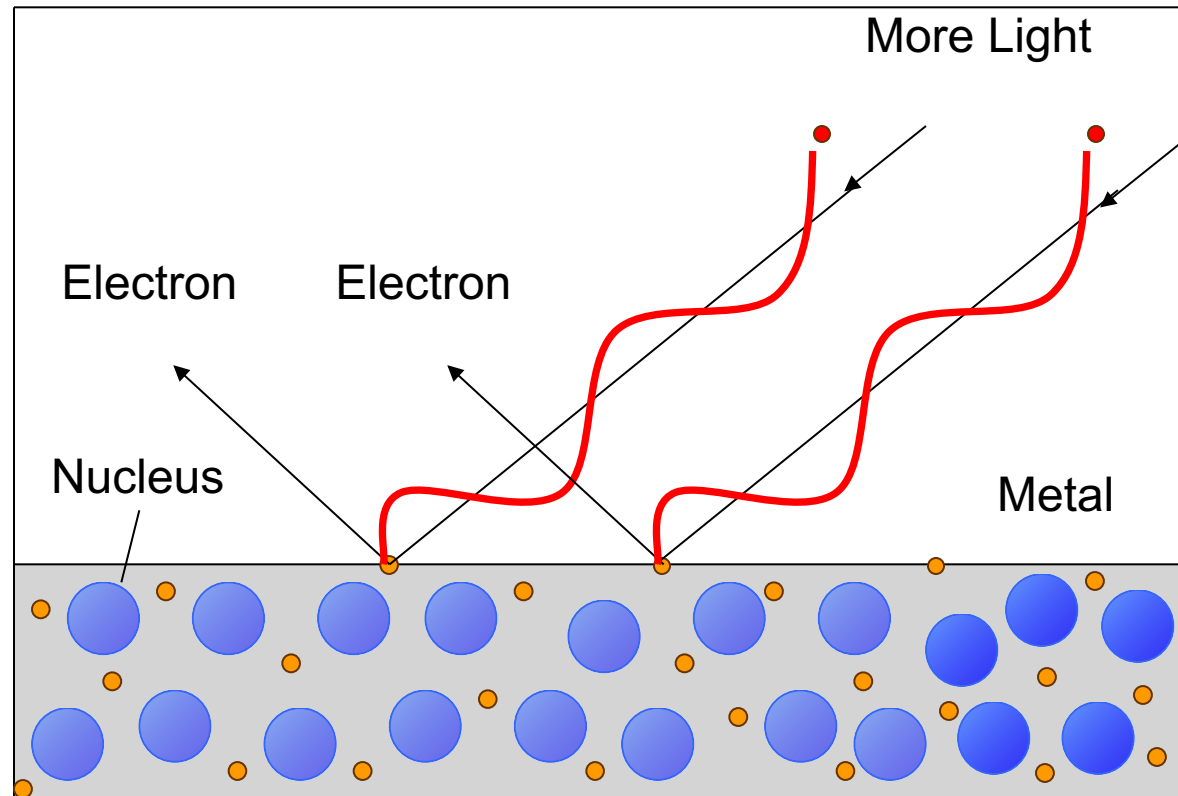


# Photoelectric Effect



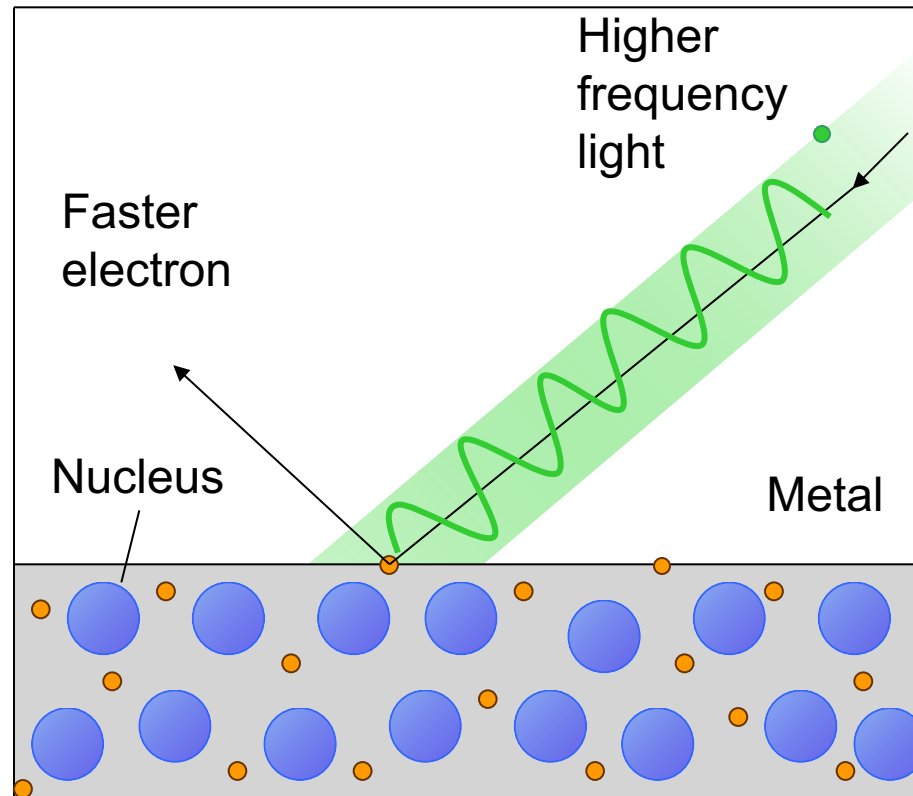
When light strikes a metal surface, electrons are ejected.

# Photoelectric Effect



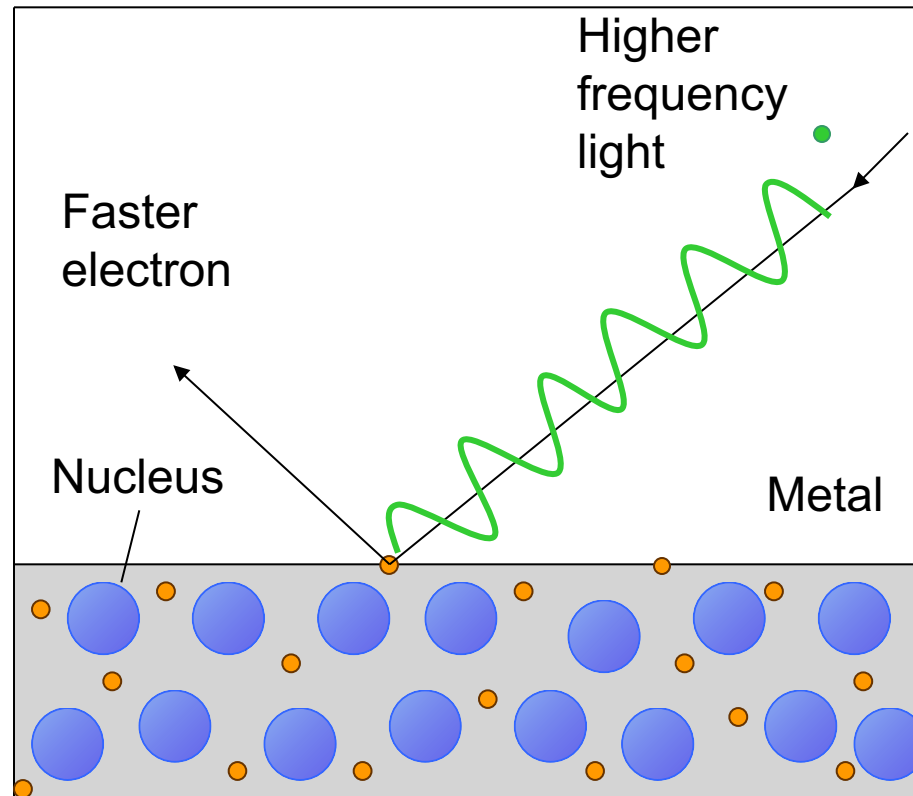
If the threshold frequency has been reached, increasing the intensity only increases the number of the electrons ejected.

# Photoelectric Effect



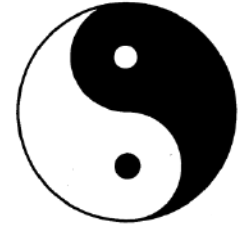
If the frequency is increased, the ejected electrons will travel faster.

# Photoelectric Effect



If the frequency is increased, the ejected electrons will travel faster.

# Strong vs. Weak Force

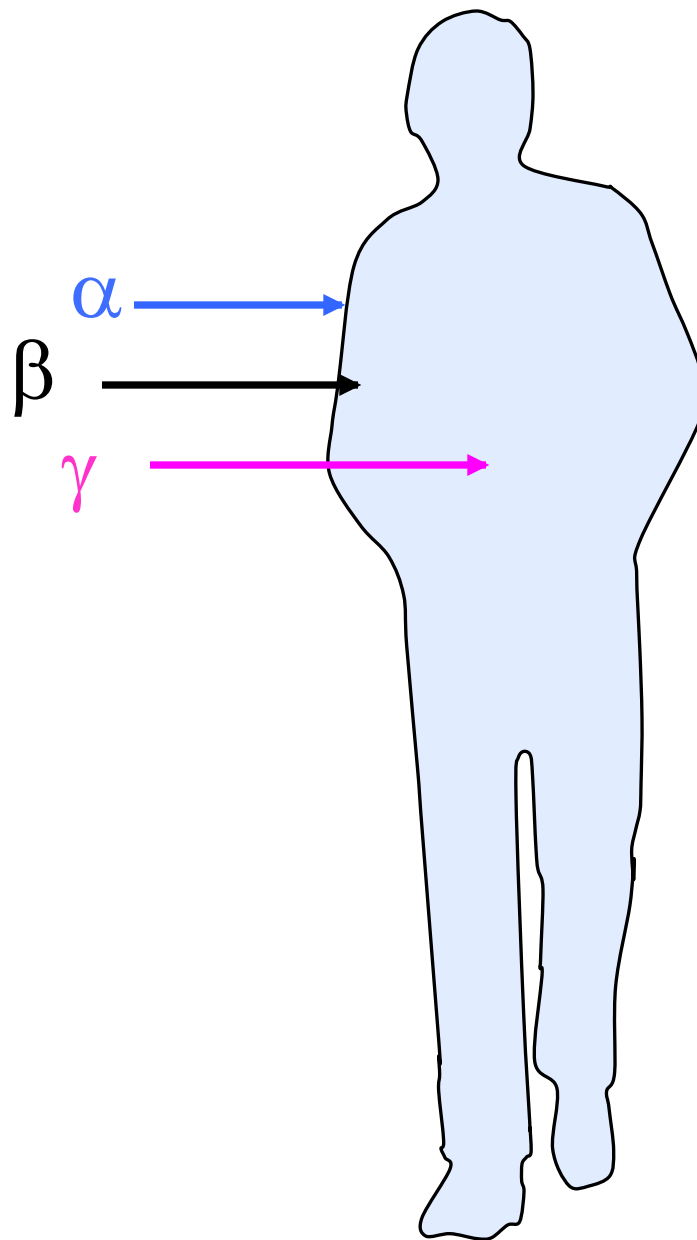
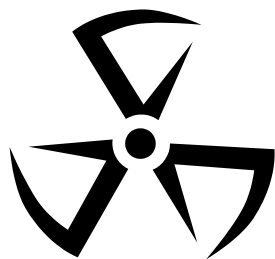


**Weak force:** electrostatic attractions between protons and electrons in atoms  
e.g. covalent bonding, ionic bonding, hydrogen bonding

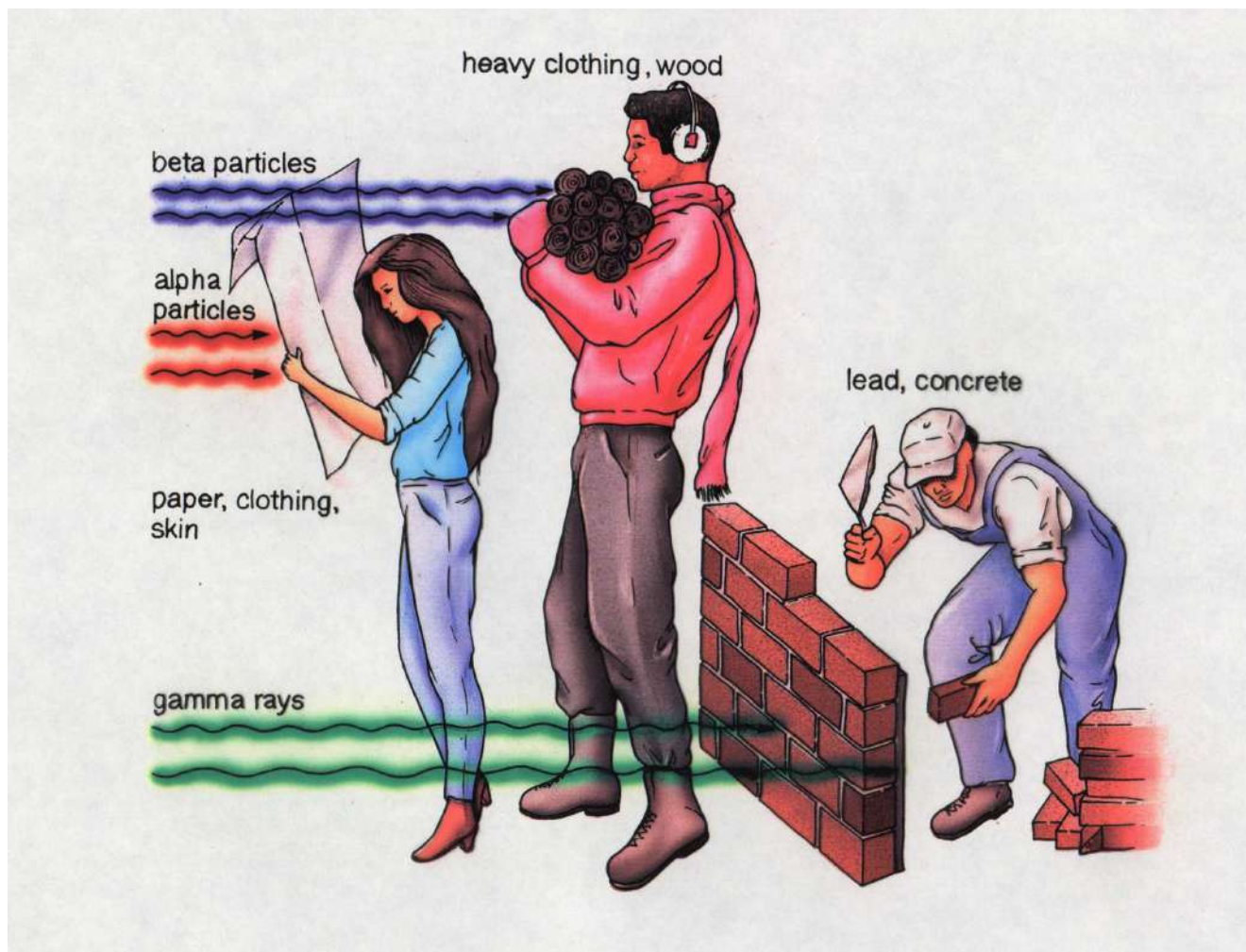
**Strong force:** force that holds the nucleus together.

*i.e.* The nucleus contains protons that naturally repel each other. The *strong force* holds the nucleus together. When the nucleus is split, the energy released is the energy of the strong force.

# Absorption of Radiation



# Absorption of Radiation



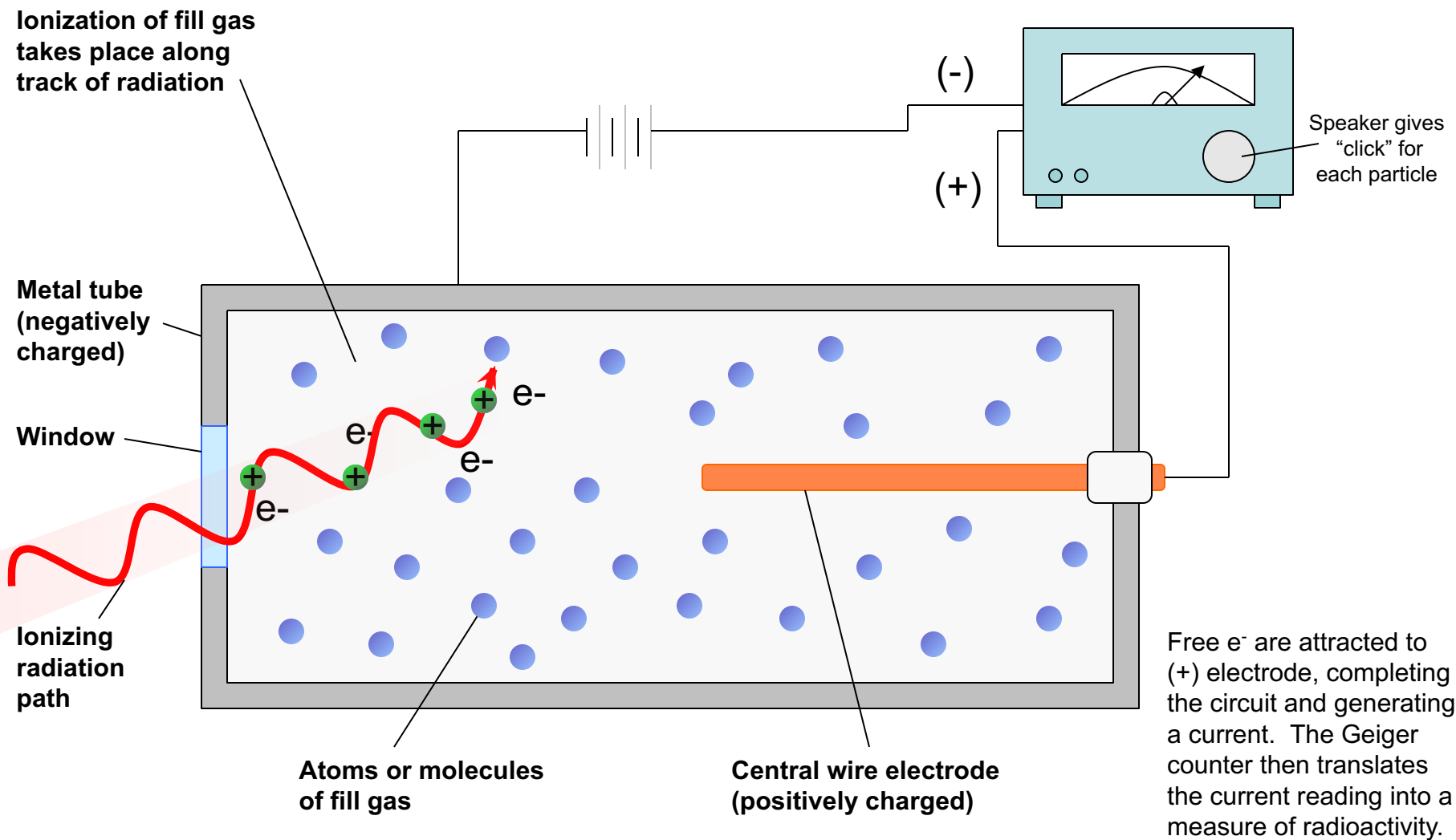


### Typical Radiation Exposure per Person per Year in the United States

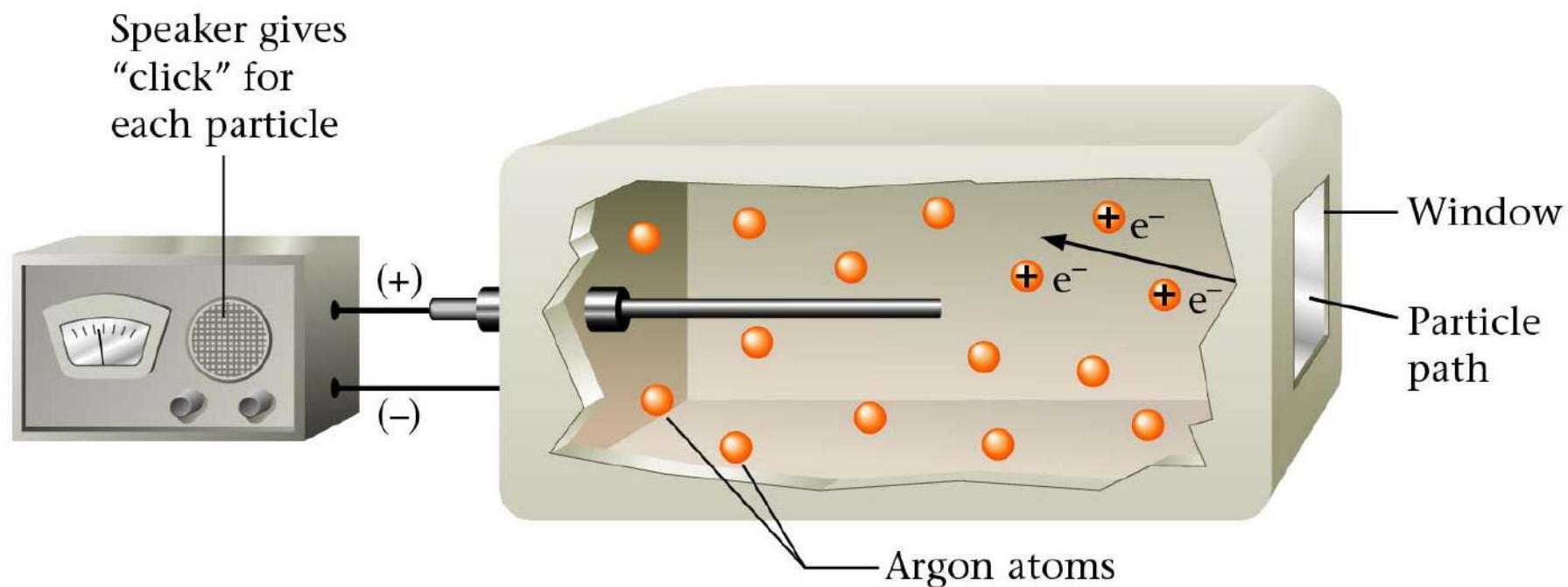
Source	Radiation	Source	Radiation
atmosphere at sea level*	26 mrem	dental X-ray	1 mrem
ground	30 mrem	chest X-ray	6 mrem
foods	20 mrem	X-ray of hip	65 mrem
air travel above 1,800 m	4 mrem	CAT scan	110 mrem
construction site	7 mrem	nuclear power plant nearby	0.02 mrem
X-ray of arm or leg	1 mrem	TV and computer use	2 mrem

\*Add 3 mrem for every 300 m of elevation

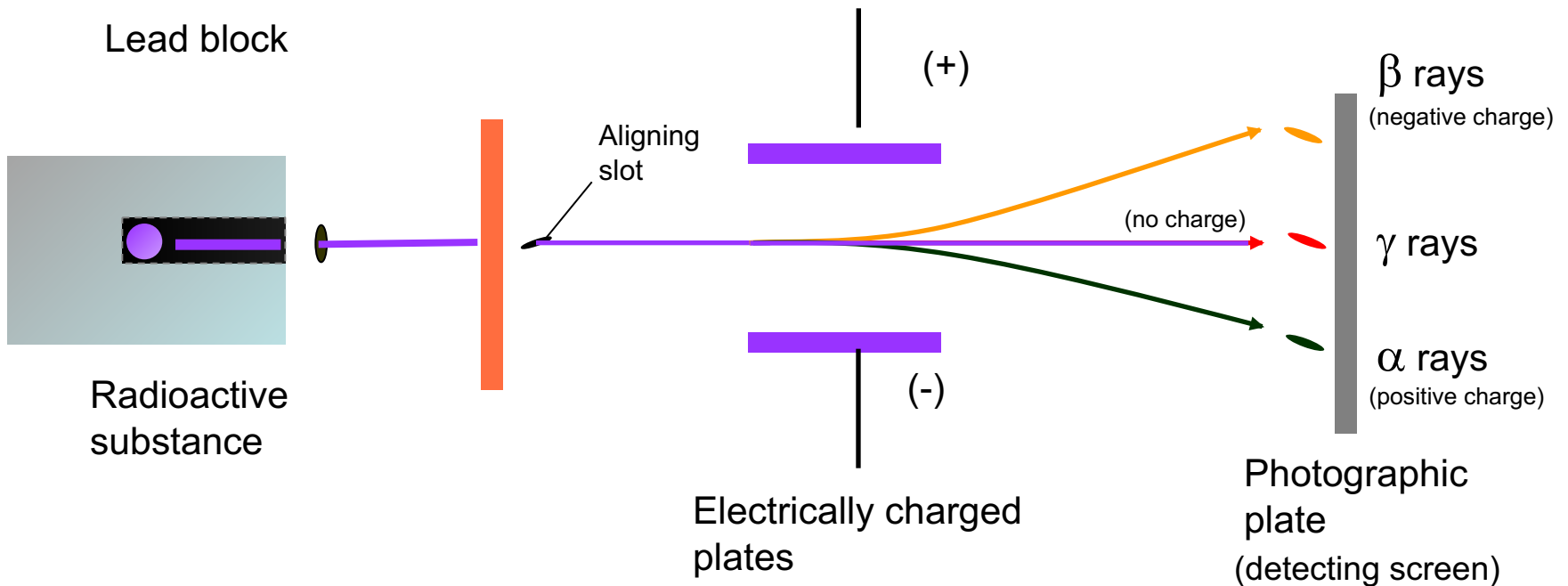
# Geiger Counter



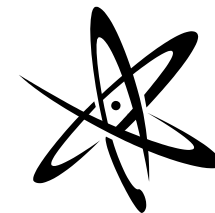
# Geiger-Muller Counter



# Alpha, Beta, Gamma Rays



# Types of Radiation



Type	Symbol	Charge	Mass (amu)
Alpha particle	${}^4_2\text{He}$	2+	4.015062
Beta particle	${}^0_{-1}\beta$	1-	0.0005486
Positron	${}^0_{+1}\beta$	1+	0.0005486
Gamma ray	$\gamma$	0	0

## Characteristics of Some Ionizing Radiations

<b><i>Property</i></b>	<b><i>Alpha radiation</i></b>	<b><i>Beta radiation</i></b>	<b><i>Gamma radiation</i></b>
<b>Composition</b>	Alpha particle (helium nucleus)	Beta particle (electron)	High-energy electromagnetic radiation
<b>Symbol</b>	$\alpha$ , He-4	$\beta$ , e	$\gamma$
<b>Charge</b>	2+	1-	0
<b>Mass (amu)</b>	4	$1/1837$	0
<b>Common source</b>	Radium-226	Carbon-14	Cobalt-60
<b>Approximate energy</b>	5 MeV*	0.05 to 1 MeV	1 MeV
<b>Penetrating power</b>	Low (0.05 mm body tissue)	Moderate (4 mm body tissue)	Very high (penetrates body easily)
<b>Shielding</b>	Paper, clothing	Metal foil	Lead, concrete (incomplete shields)

\*(1 MeV =  $1.60 \times 10^{-13}$  J)

# Nuclear reactions

Nuclear equations show how atoms decay.

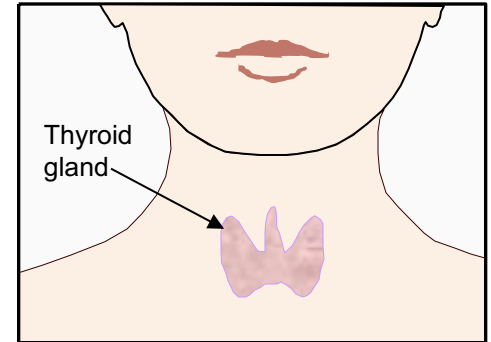
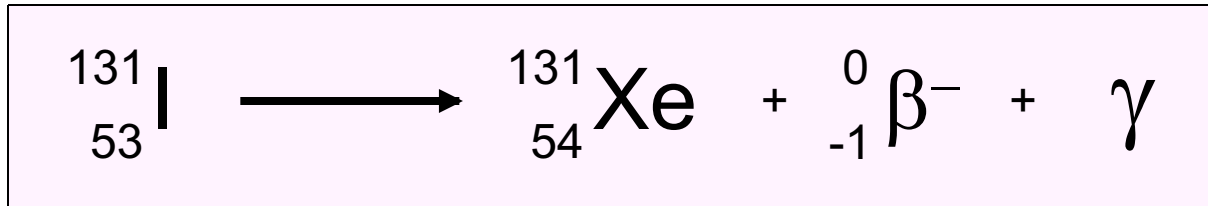
Similar to chemical equations.

- *must still balance mass and charge.*

Differ from chemical equations because

- *we can change the elements.*  
...transmutation
- *the type of isotope is important.*

A patient is given radioactive iodine to test thyroid function.  
What happens to the iodine?



Is this equation balanced?

You must see if the mass and charge are the same on both sides.

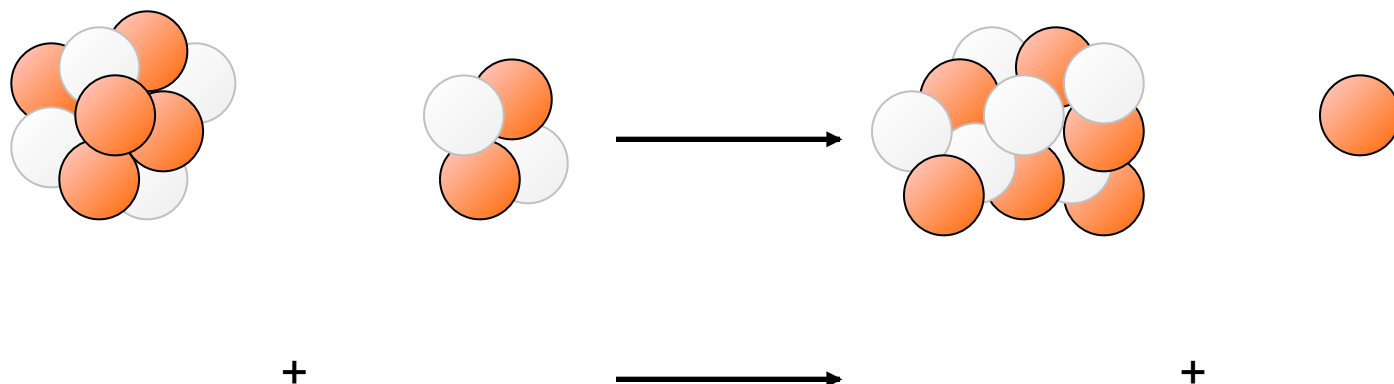
Mass		Charge	
53 protons	54 protons	+53, protons	+54, protons
78 neutrons	77 neutrons		-1 charge from $\beta^{-}$
<hr/> 131 total mass	<hr/> 131 total mass	<hr/> +53 total charge	<hr/> +53 total charge

Yes – it's balanced



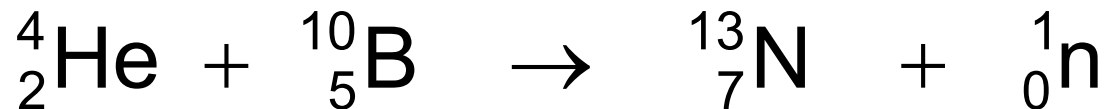
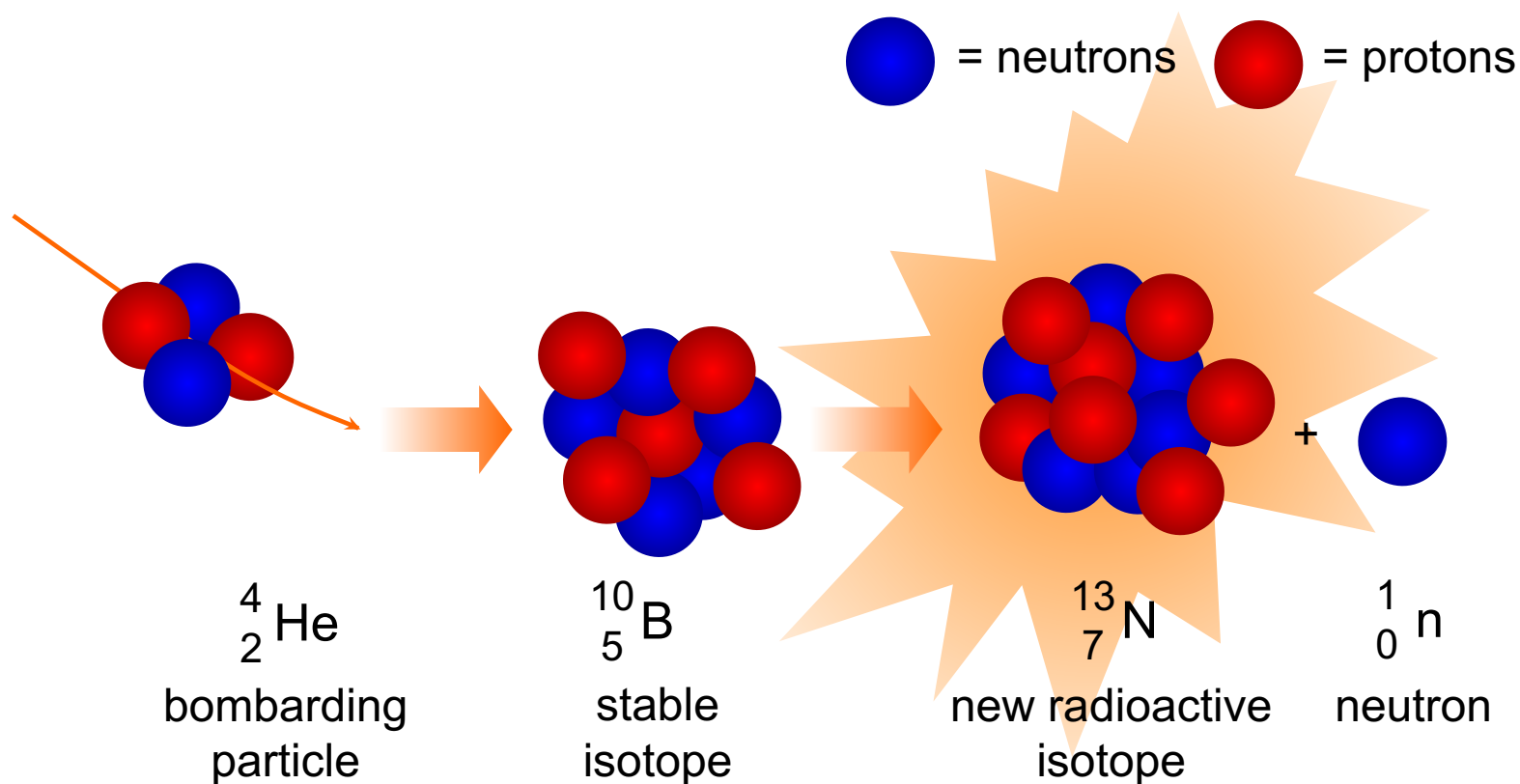


# Discovery of the Neutron

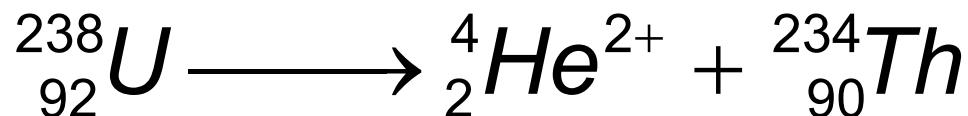
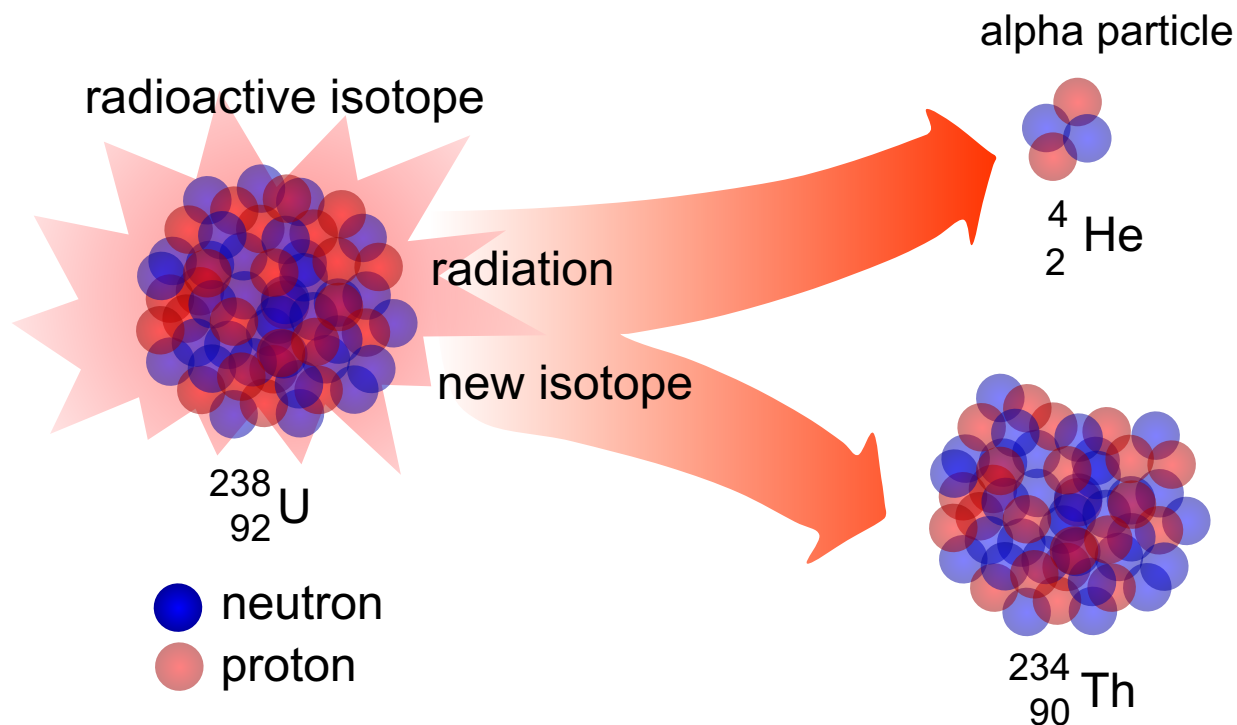


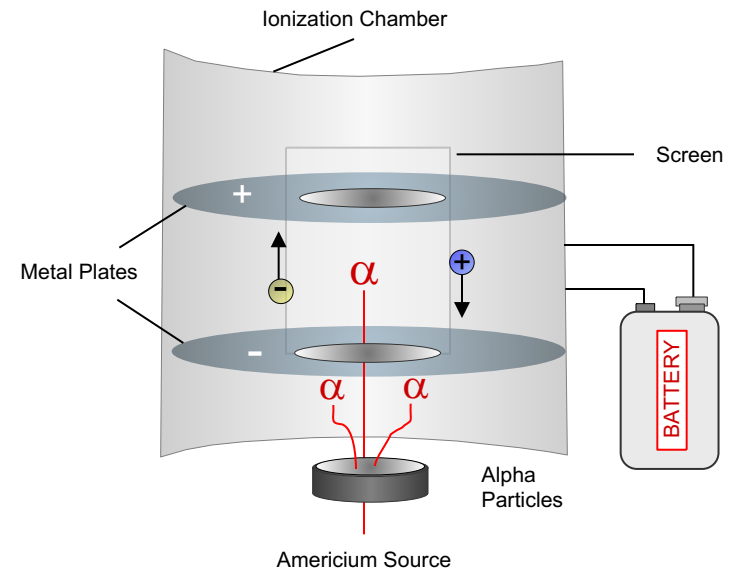
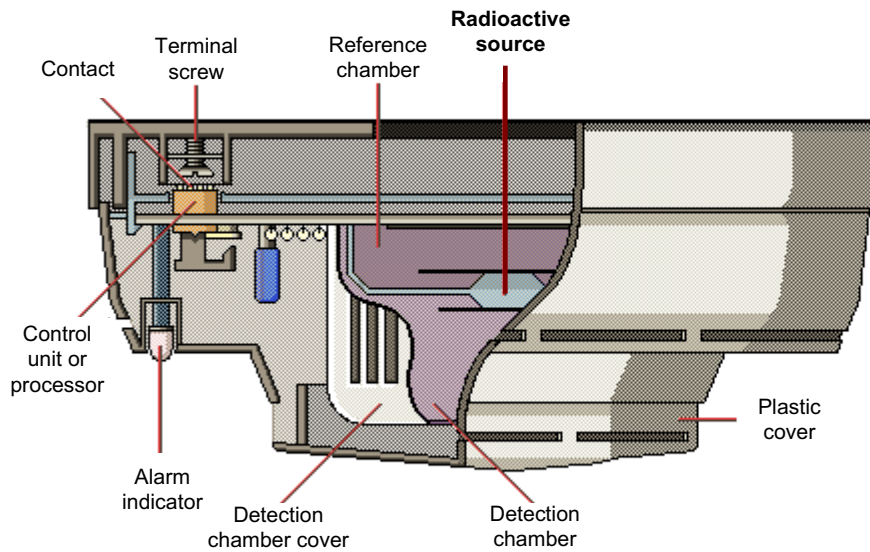
James Chadwick bombarded beryllium-9 with alpha particles, carbon-12 atoms were formed, and neutrons were emitted.

# New Radioactive Isotope



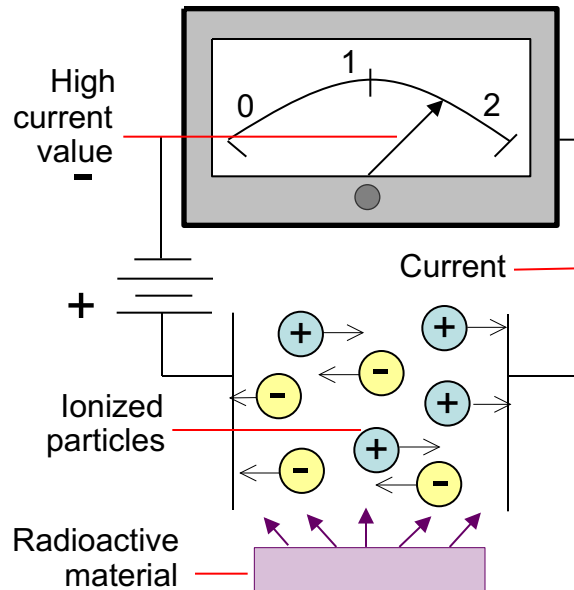
# Alpha Decay



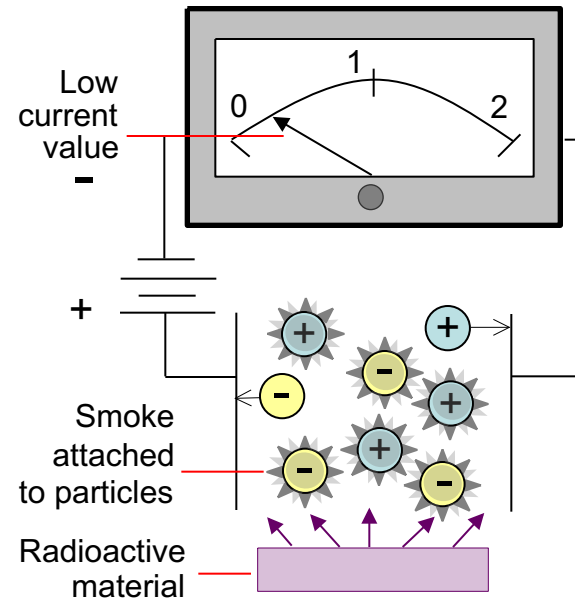


## Measuring Circuit in Detection Chamber

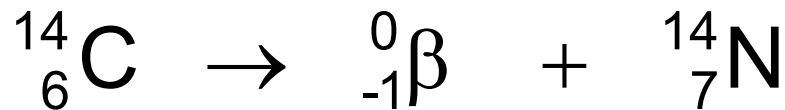
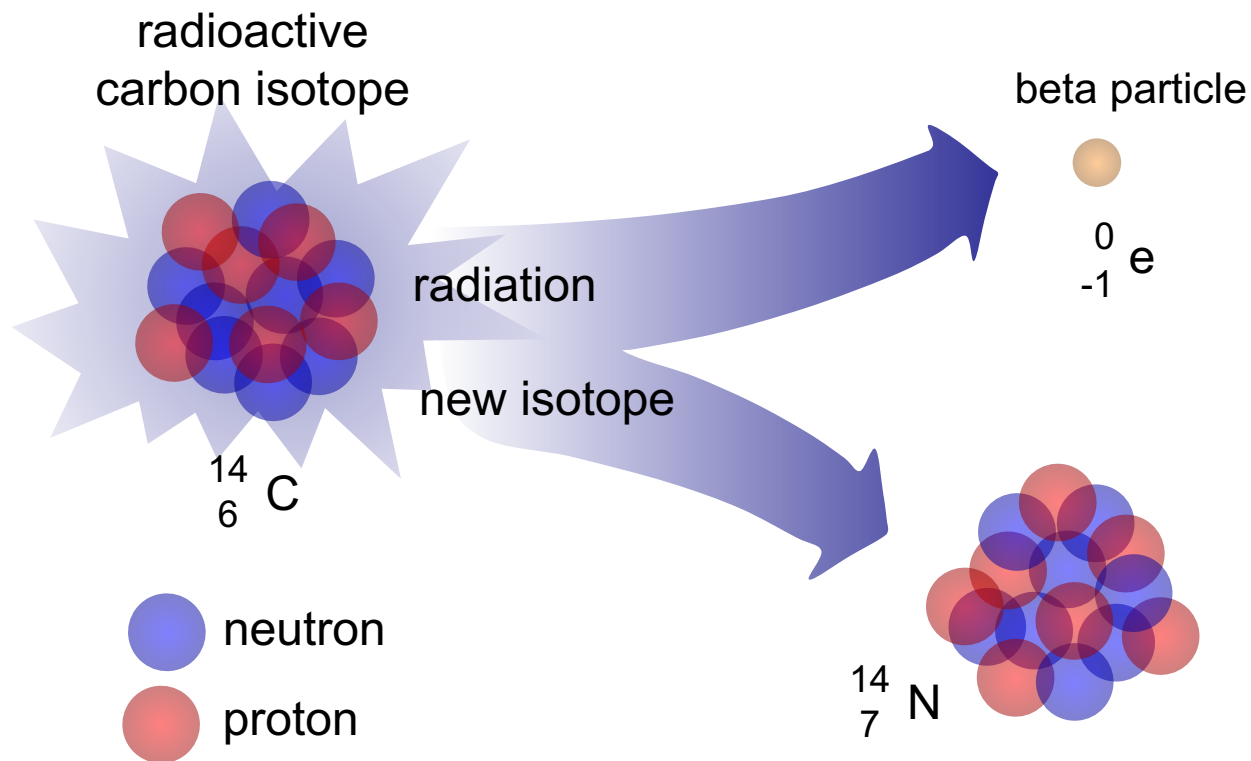
### Clean air



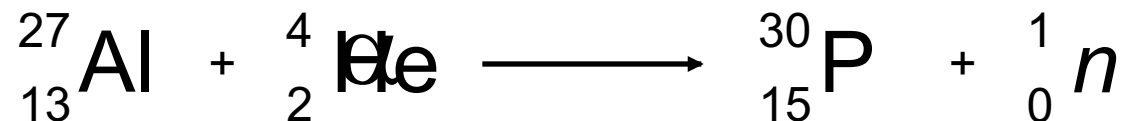
### Smoke



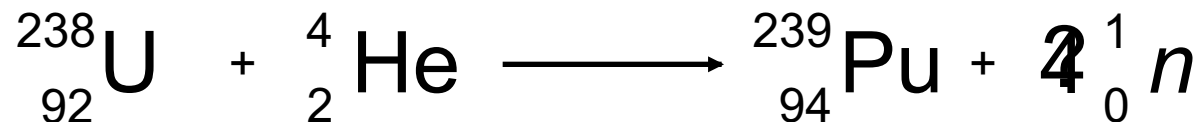
# Beta Decay



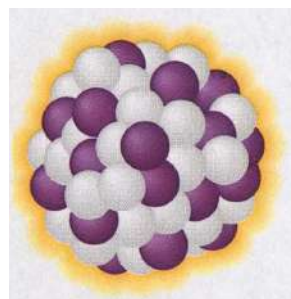
Bombardment of aluminum-27 by alpha particles produces phosphorous-30 and one other particle. Write the nuclear equation and identify the other particle.



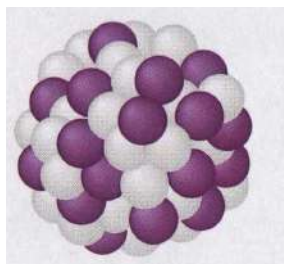
Plutonium-239 can be produced by bombarding uranium-238 with alpha particles. How many neutrons will be produced as a by product of each reaction. Write the nuclear equation for this reaction.



# Unstable Isotopes



Excited  
nucleus



Stable  
nucleus

+

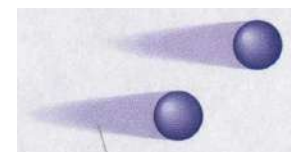


Energy

and



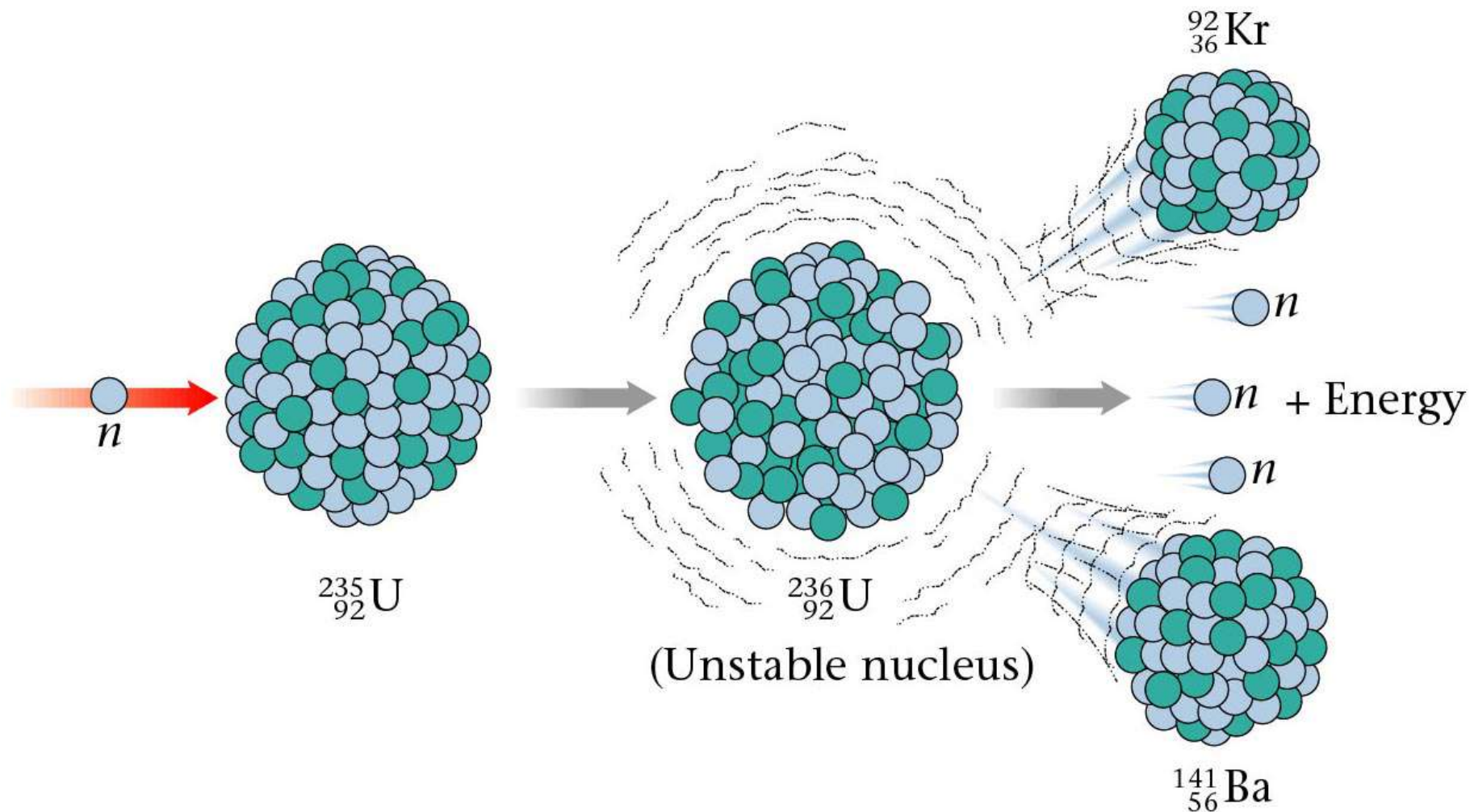
or



Particles

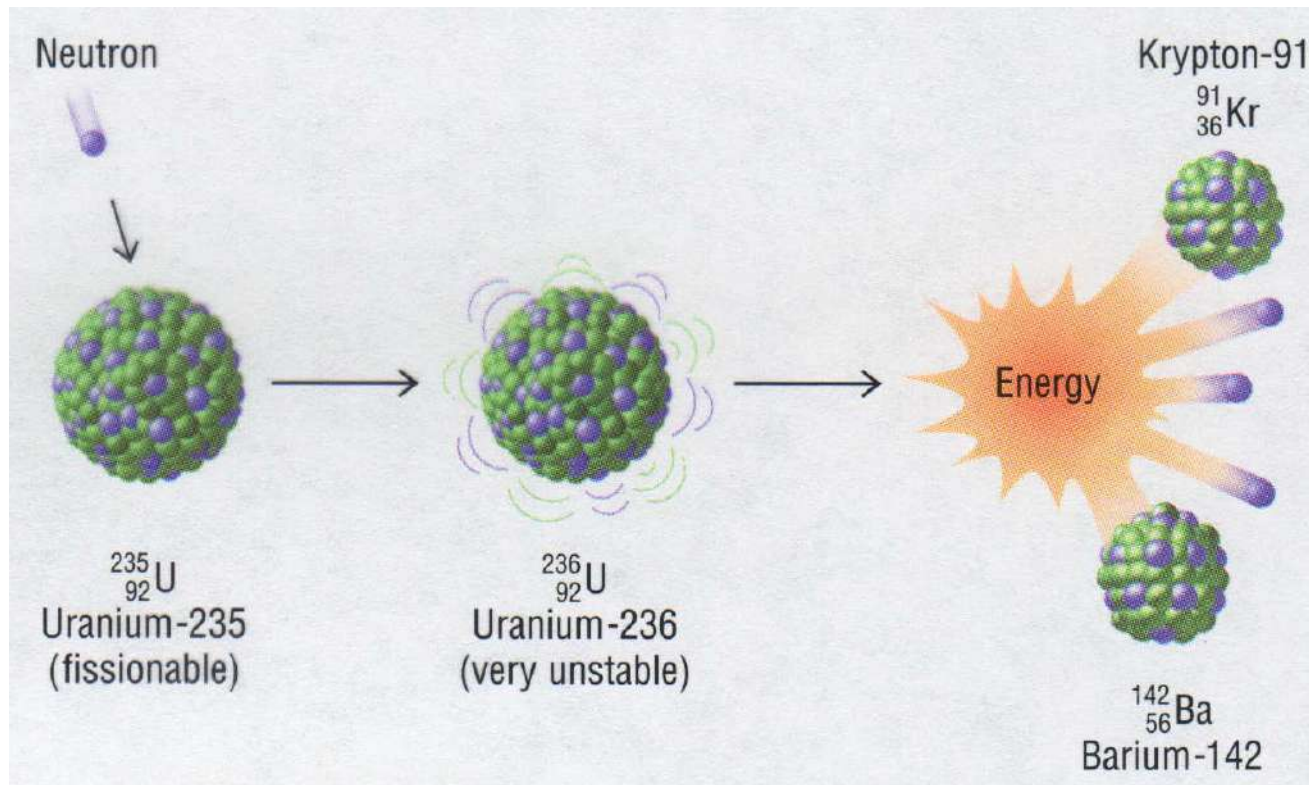
Radiation

# Unstable Nucleus

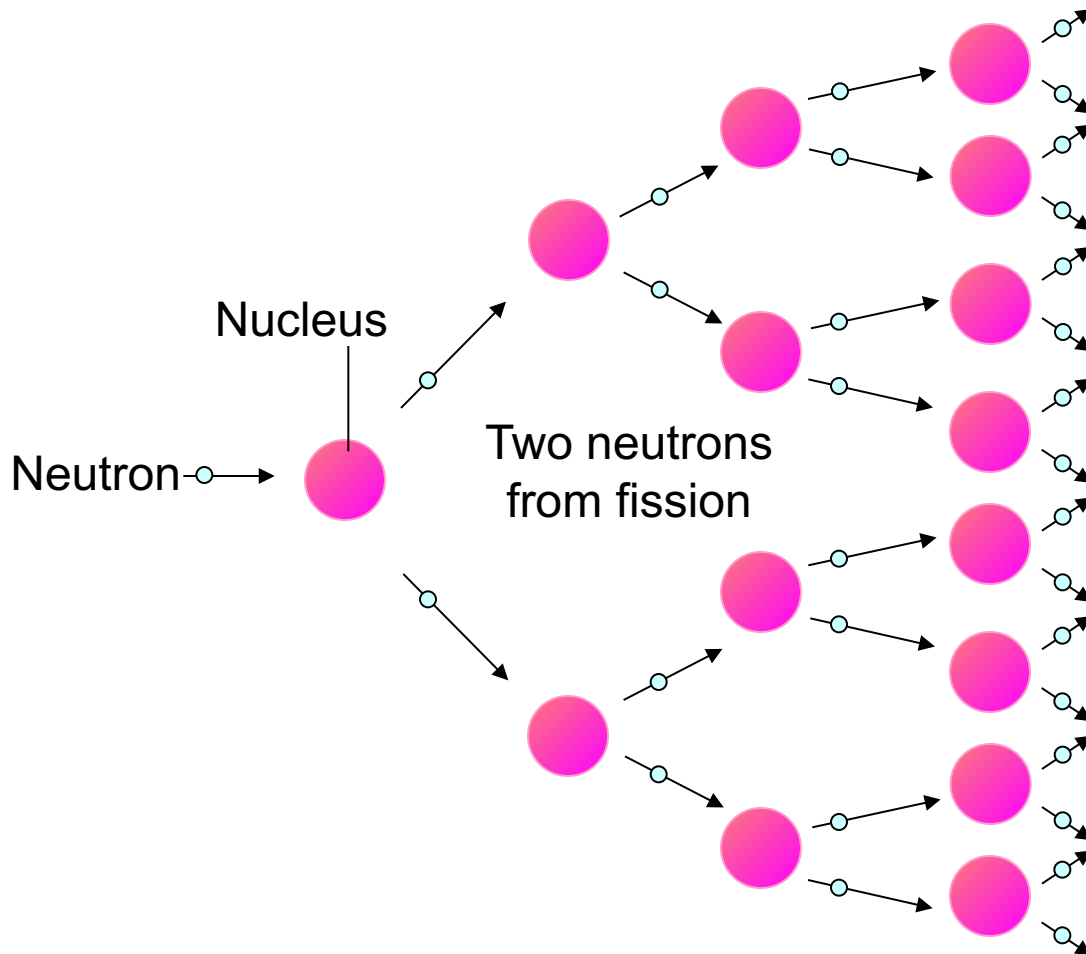




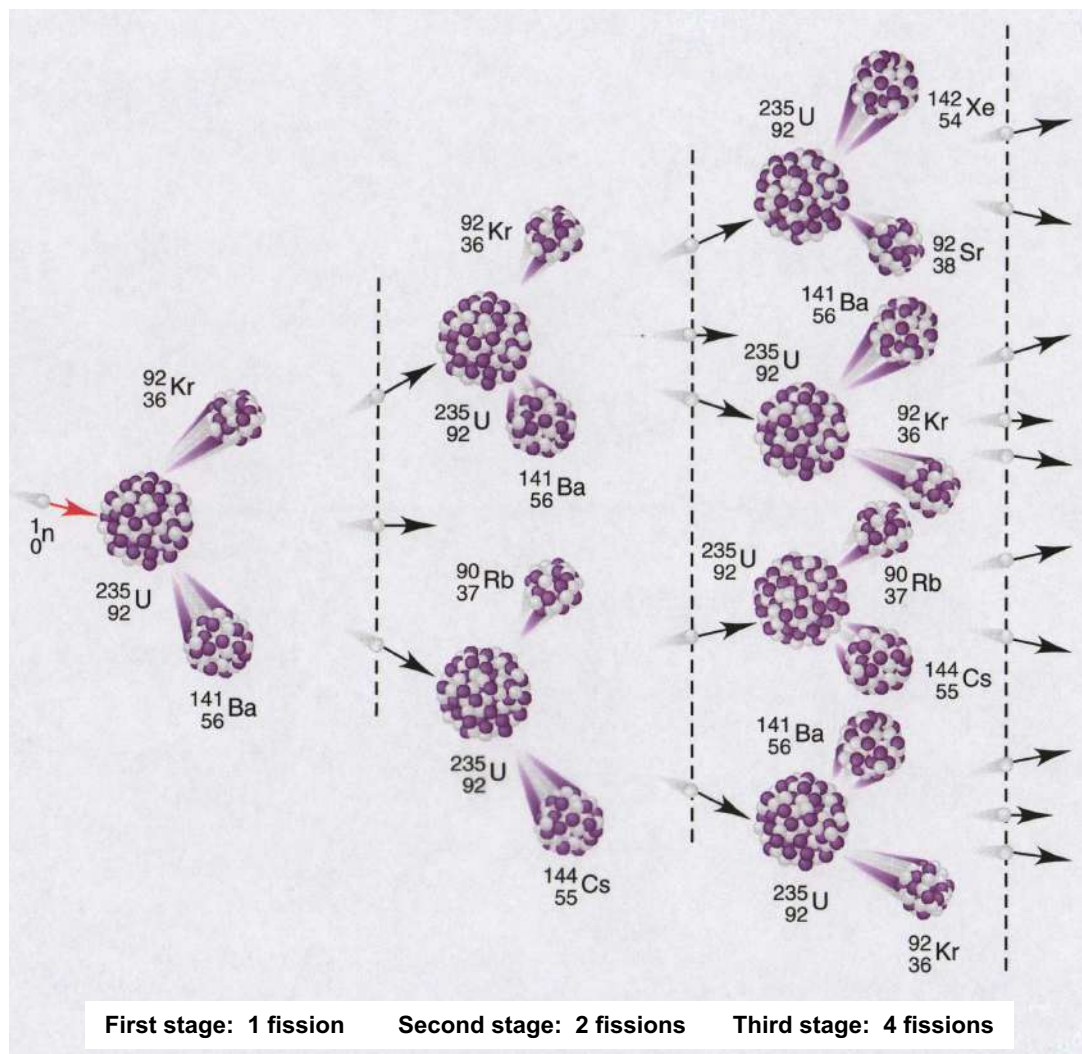
# Fissionable U-235



# Fission Process

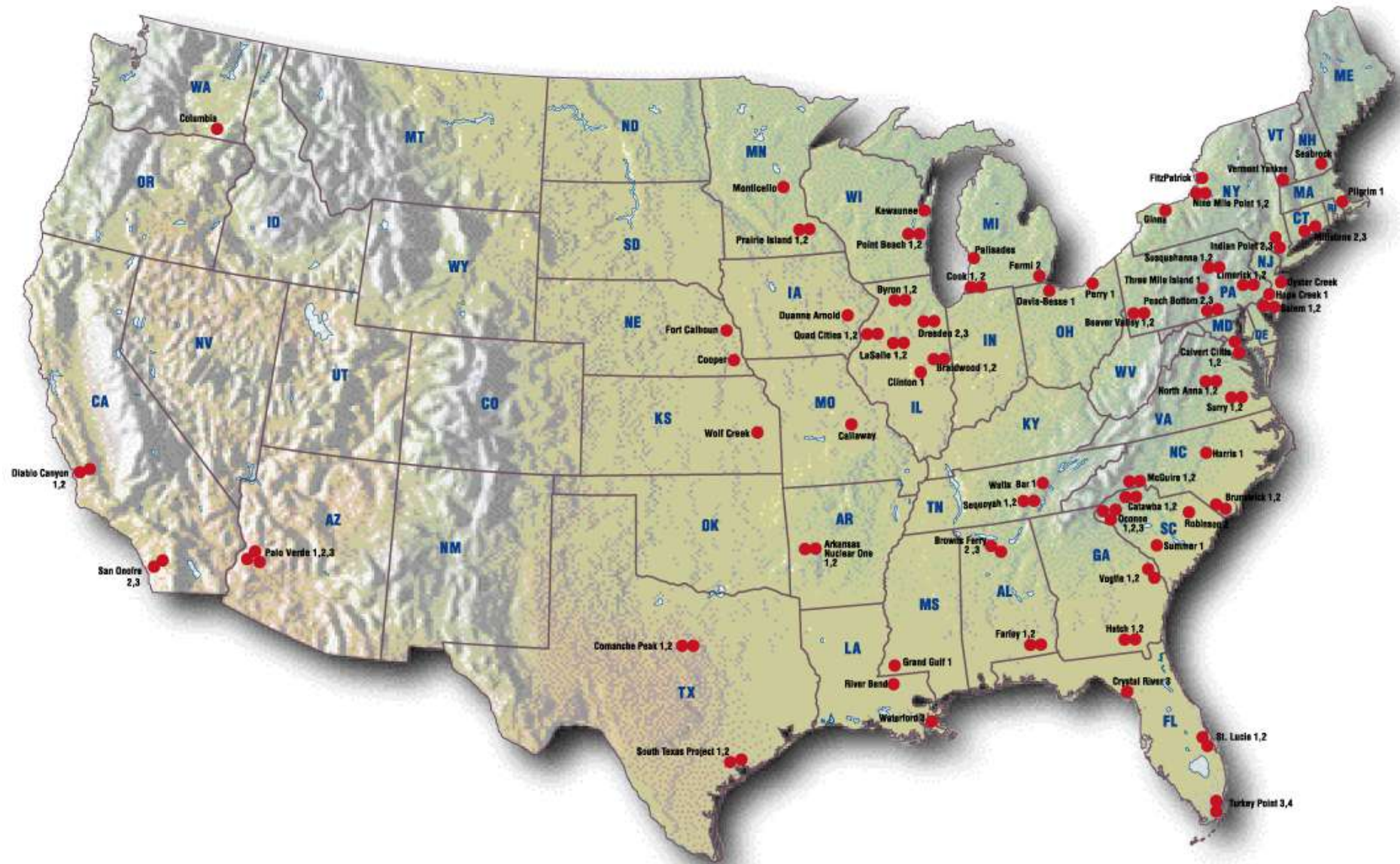


# Stages of Fission

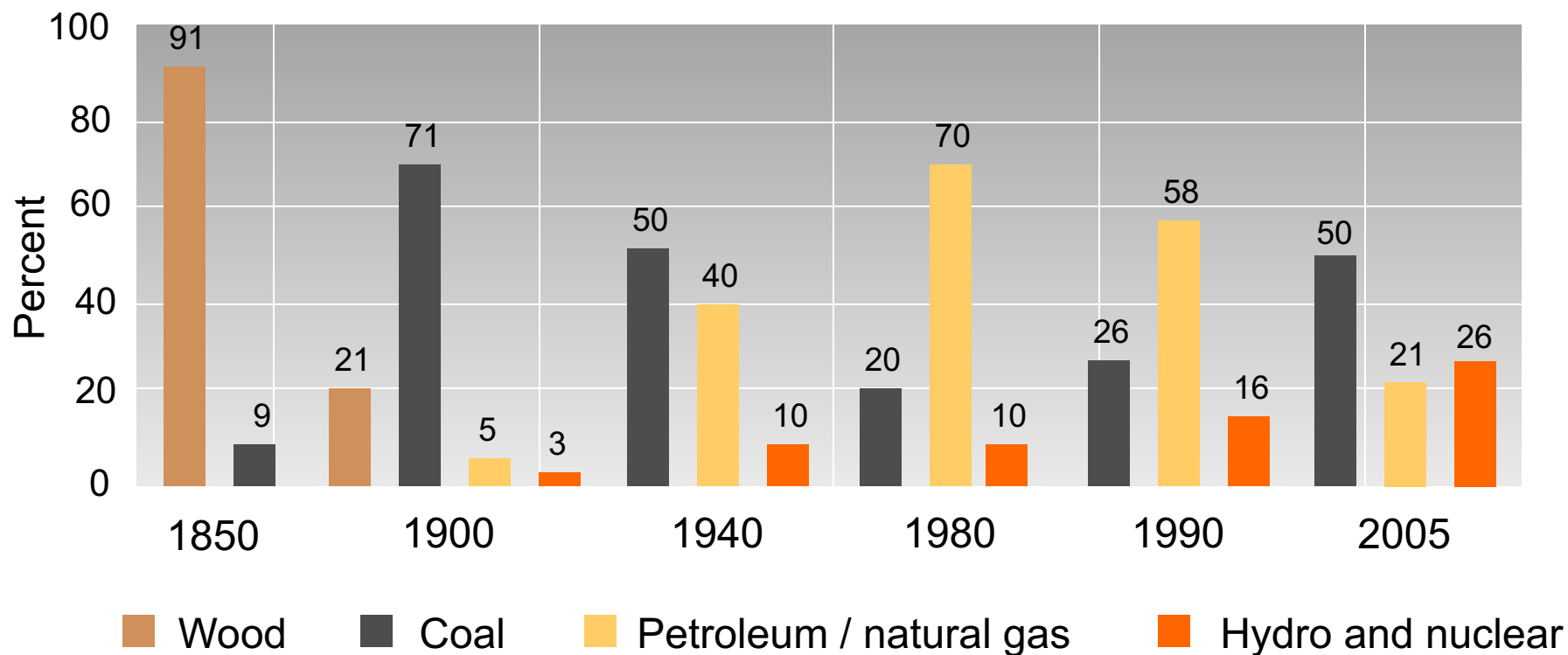




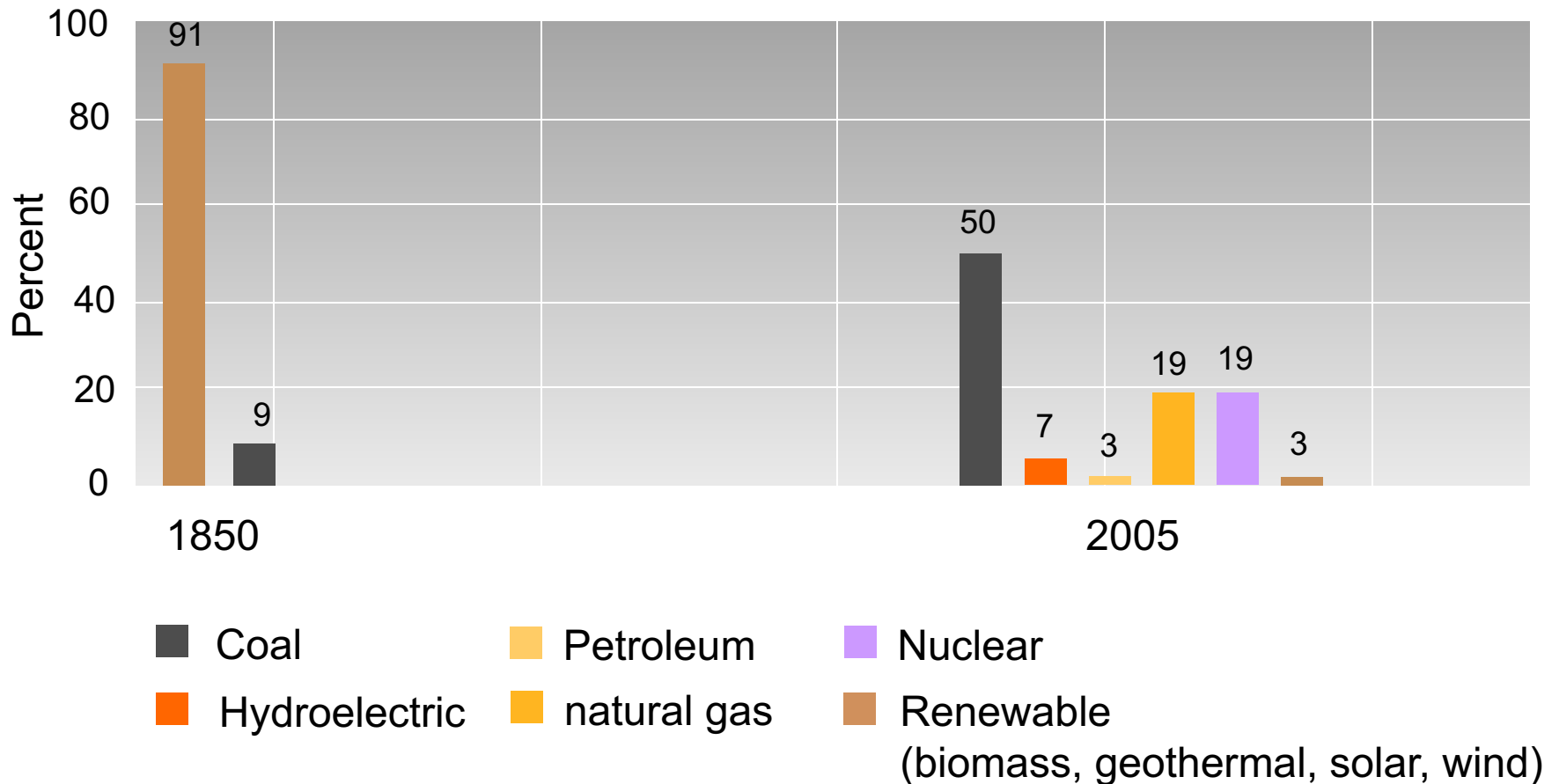
# Nuclear Power Plants



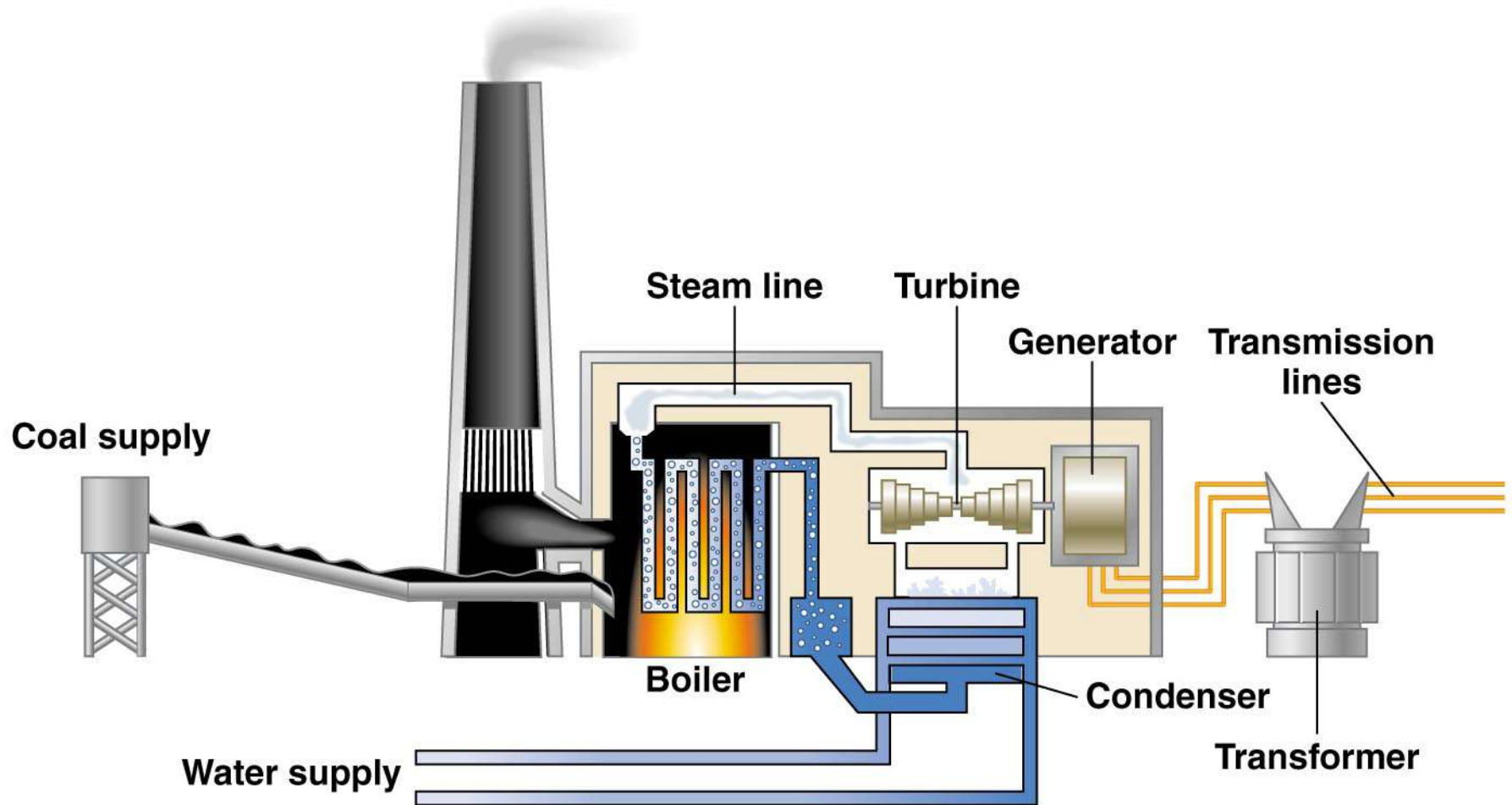
# Energy Sources in the United States



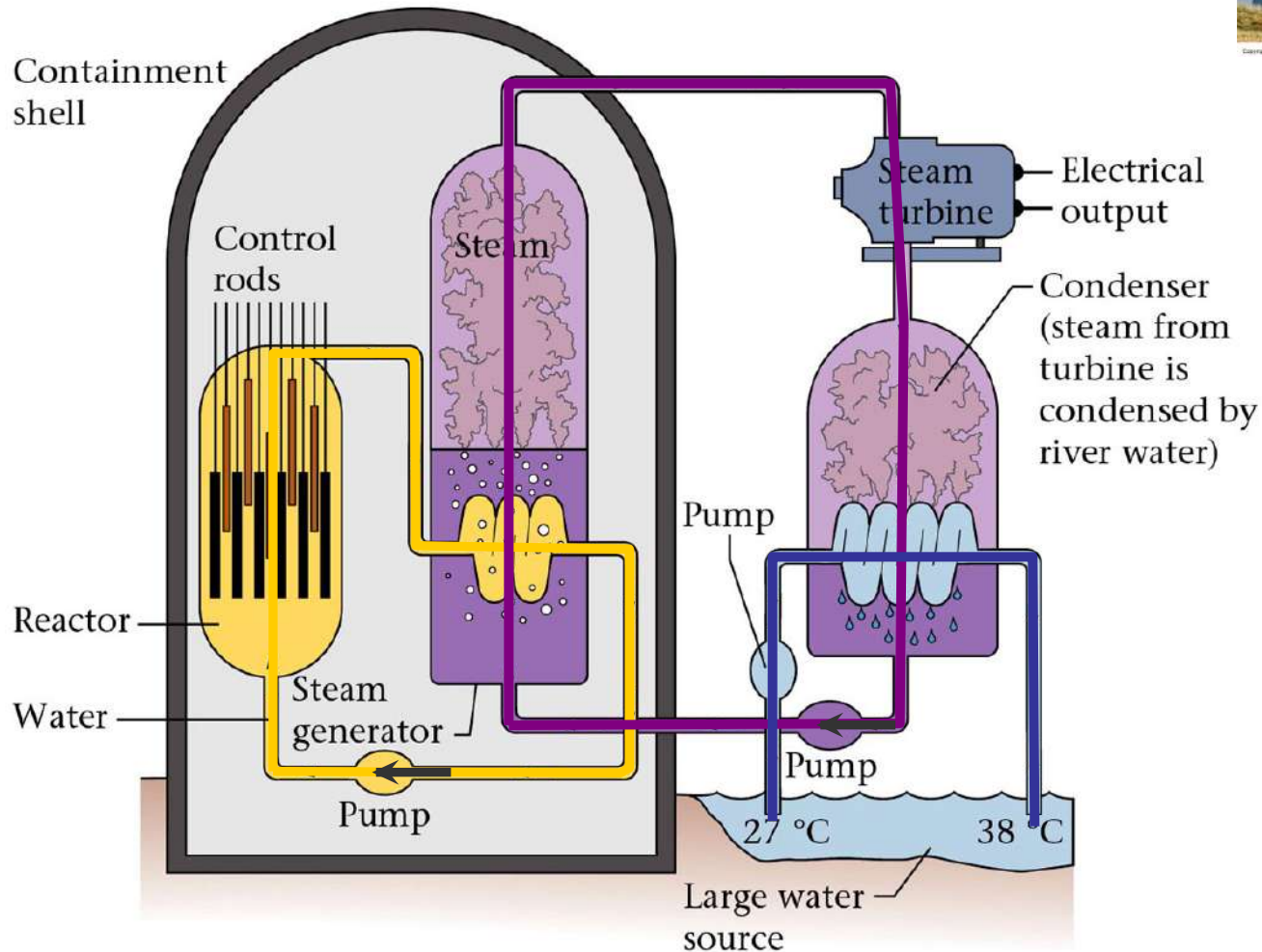
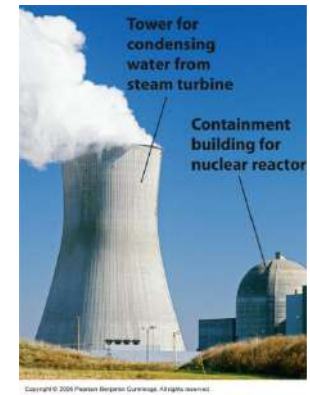
# Energy Sources in the United States



# Coal Burning Power Plant

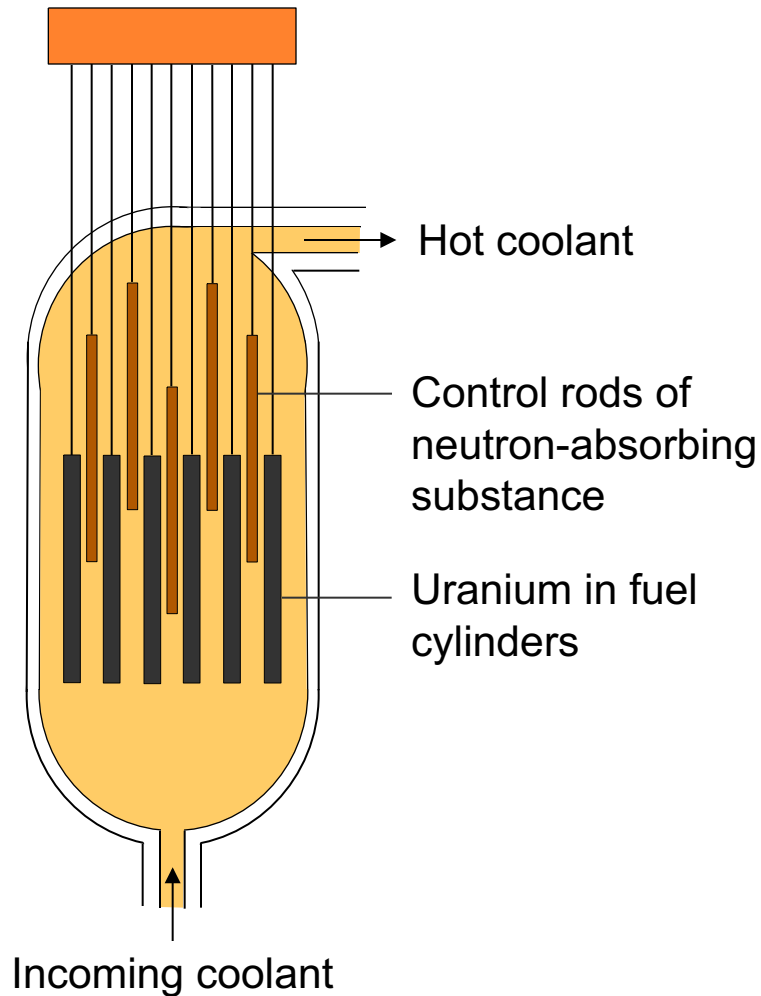


# Nuclear Power Plant

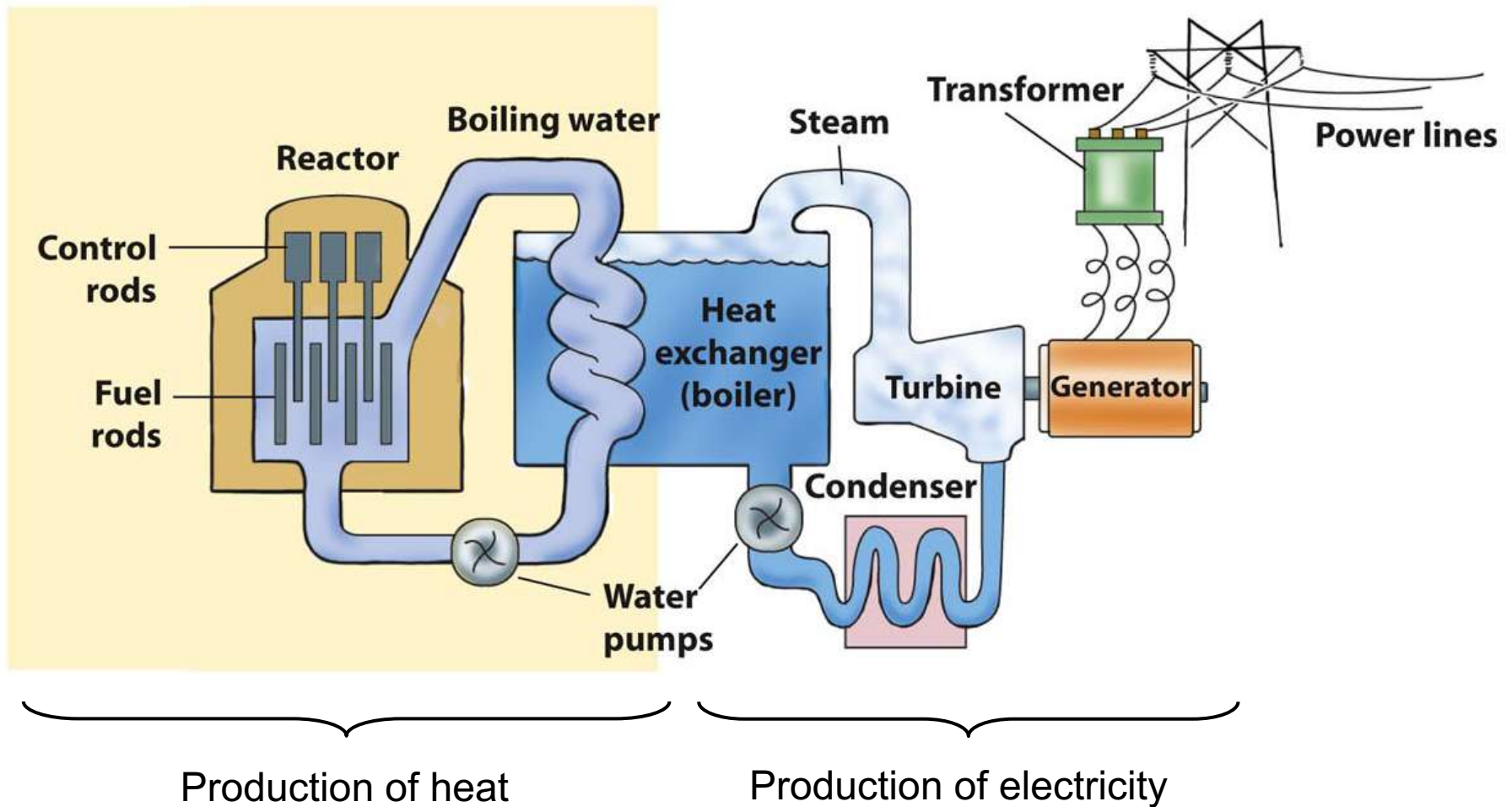




# Reactor Core



# Nuclear Power Plant



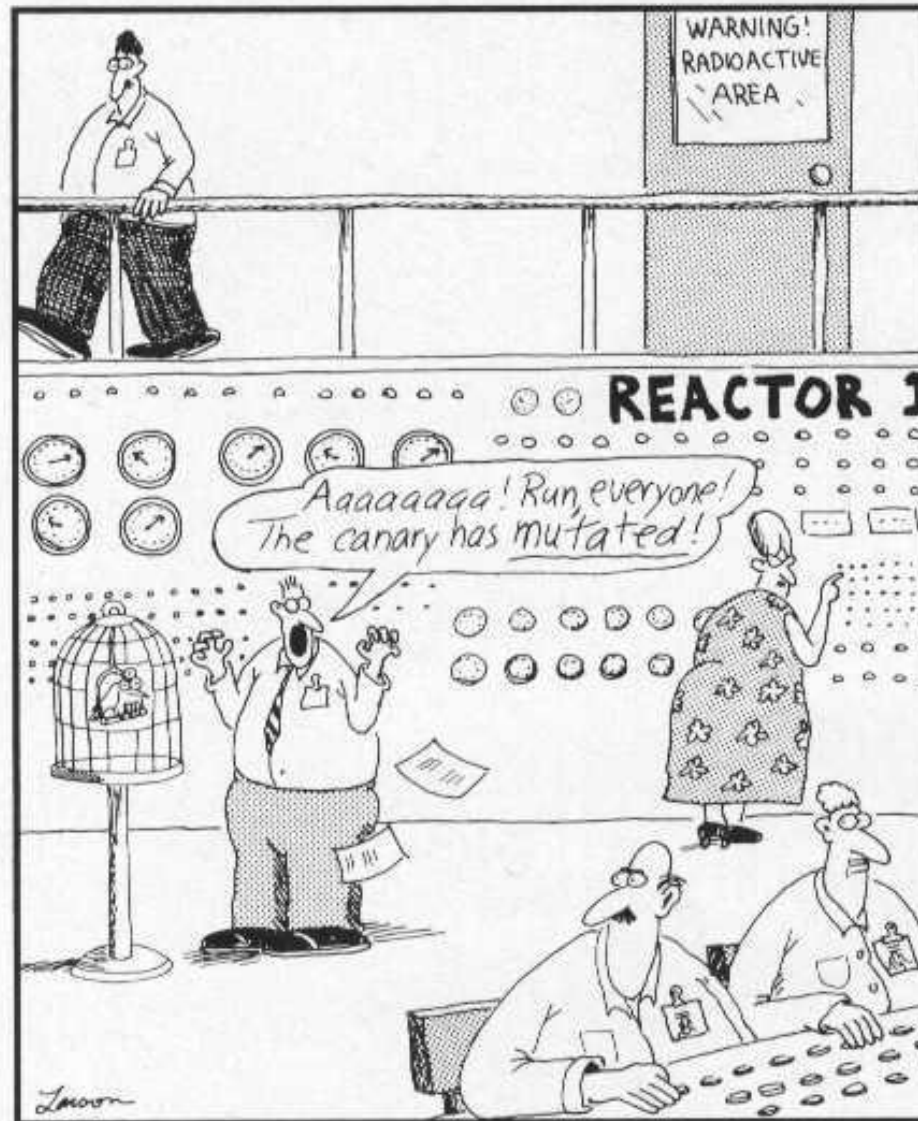


# Chant of the Radioactive Workers



We're not afraid of the alpha ray.  
A sheet of paper will keep it away!  
A beta ray needs much more care,  
Place sheets of metal here and there.  
And as for the powerful gamma ray  
(Pay careful heed to what we say)  
Unless you wish to spend weeks in bed  
Take cover behind thick slabs of lead!  
Fast neutrons pass through everything.  
Wax slabs remove their nasty sting.  
These slow them down, and even a moron  
Knows they can be absorbed by boron.  
Remember, remember all that we've said,  
Because it's no use remembering when you're dead.



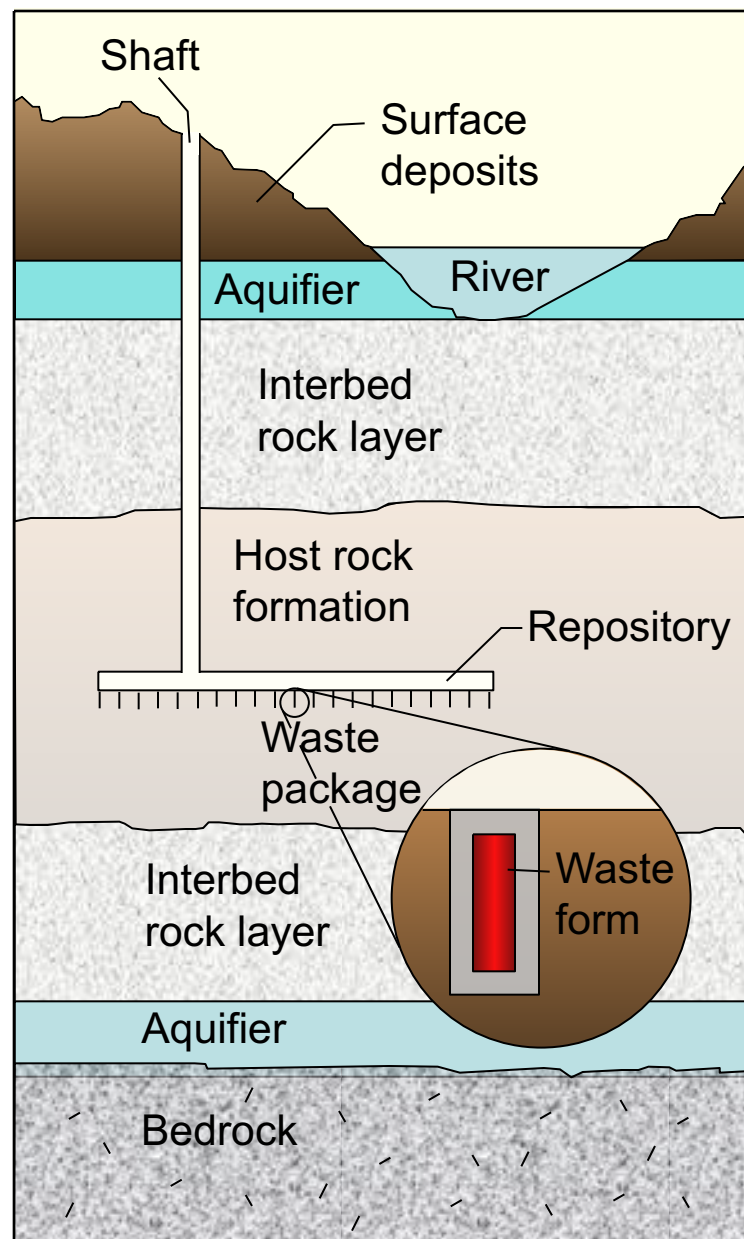


Inside a nuclear power plant.

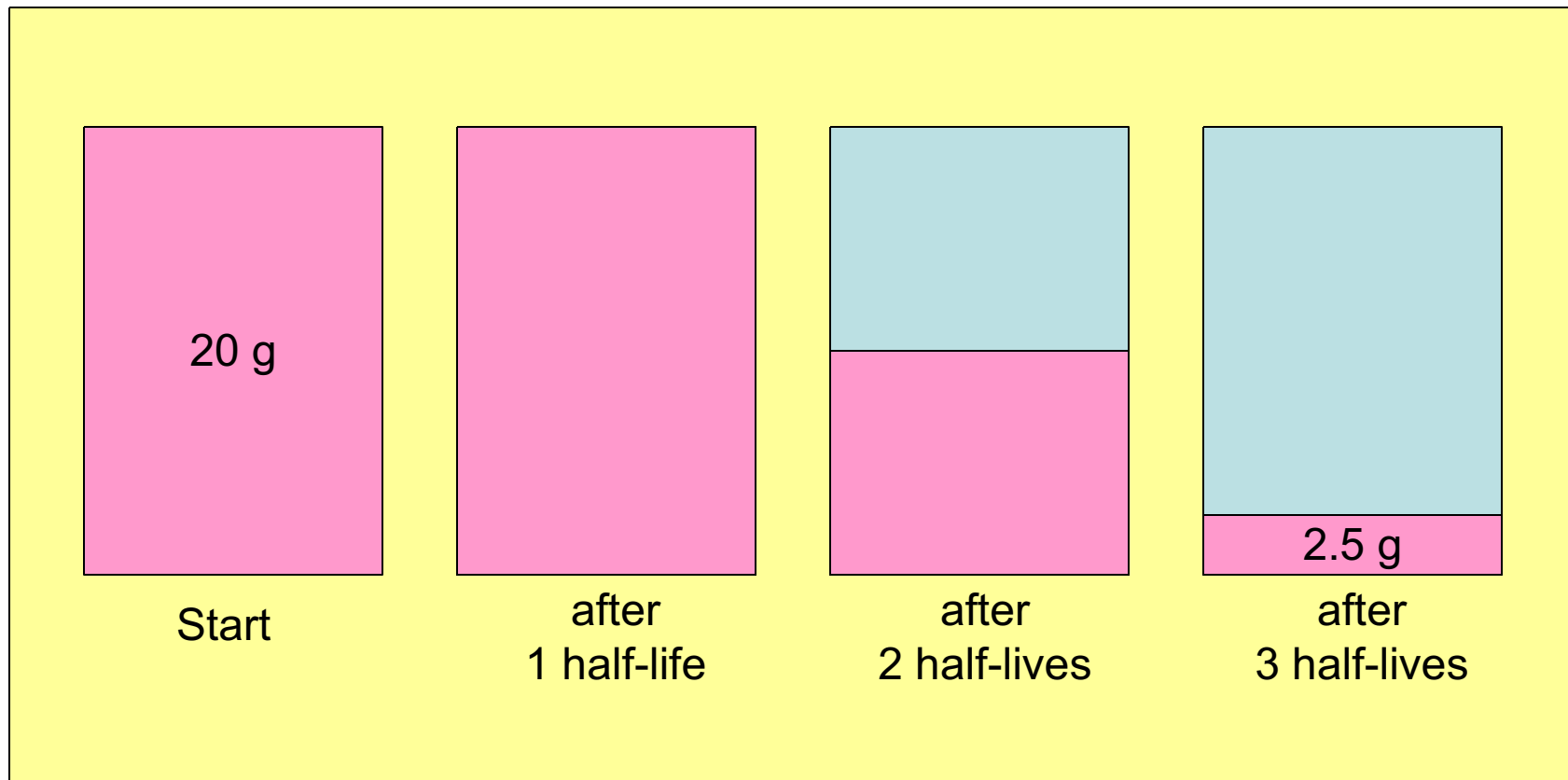
# Nuclear Waste Disposal



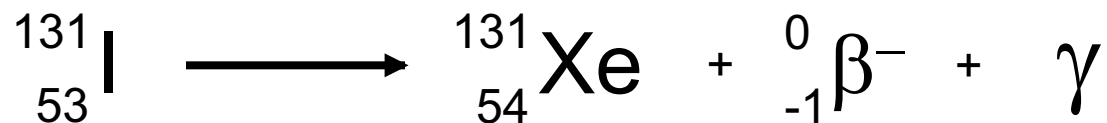
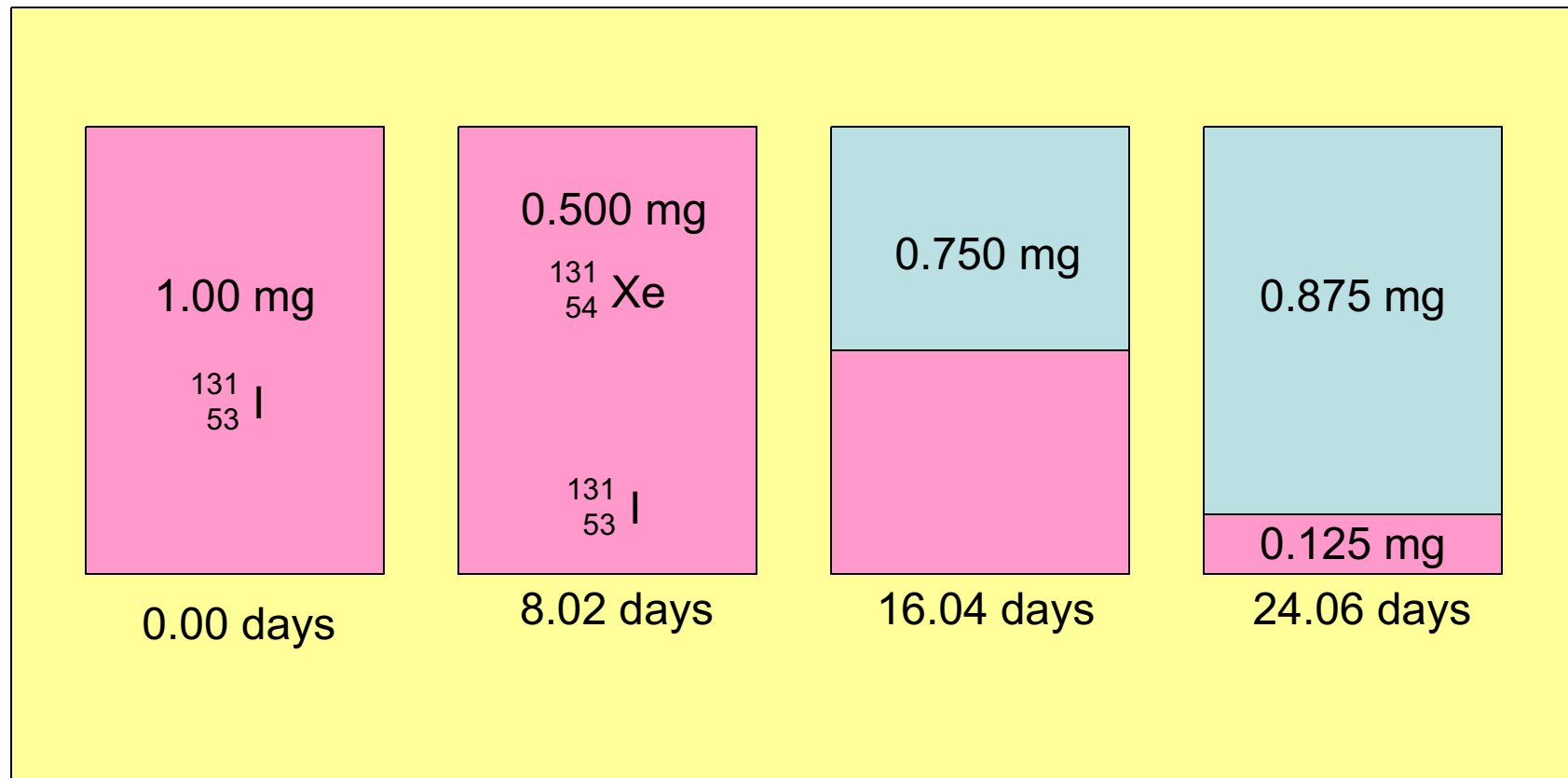
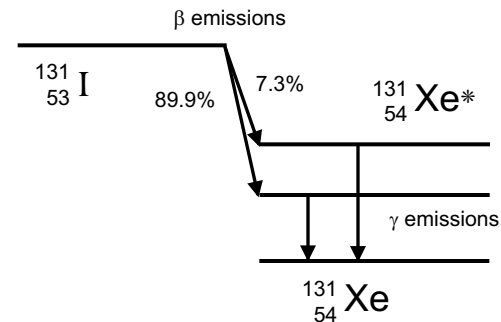
Copyright © 2009 Pearson Benjamin Cummings. All rights reserved.



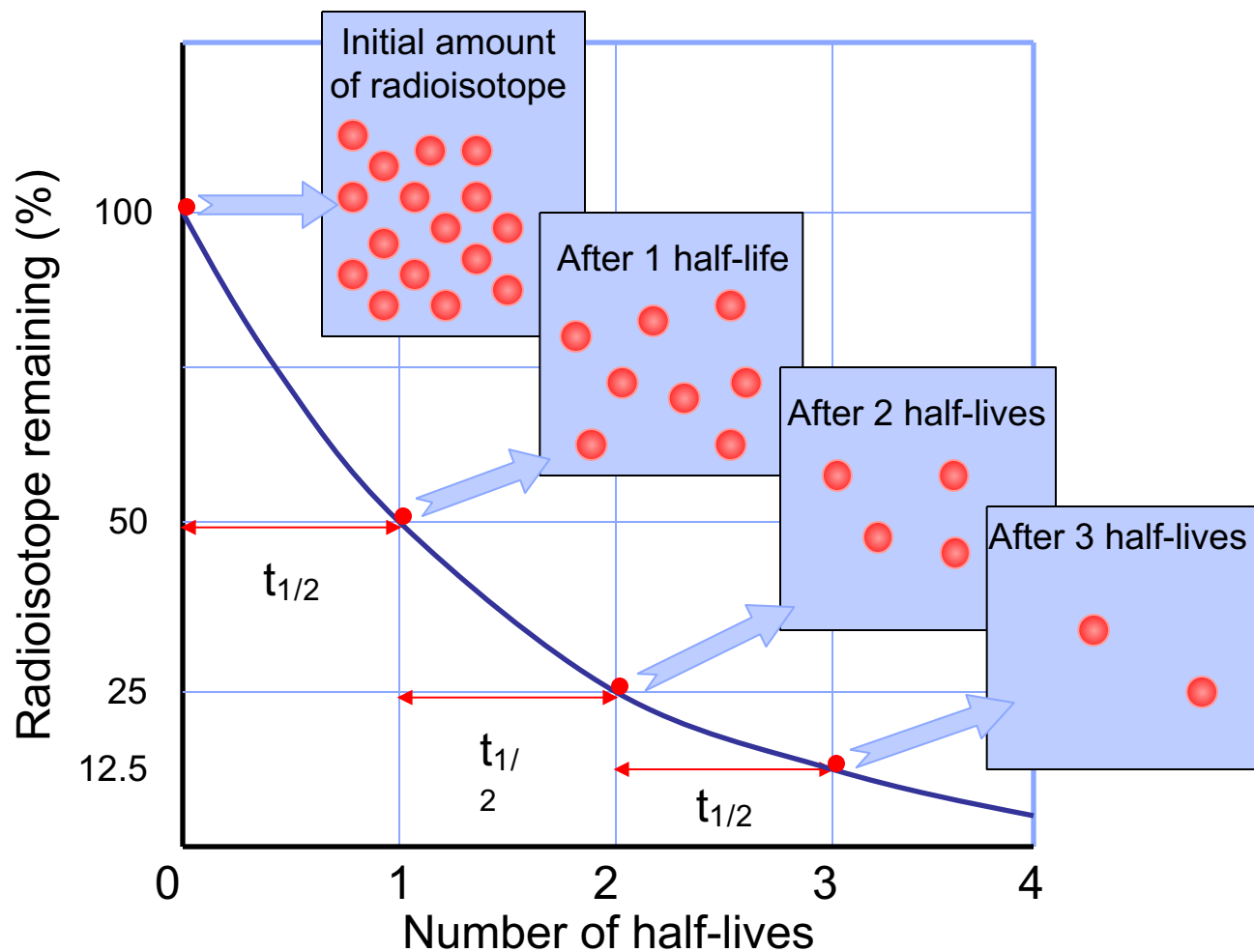
# Half-Life



# Half-Life

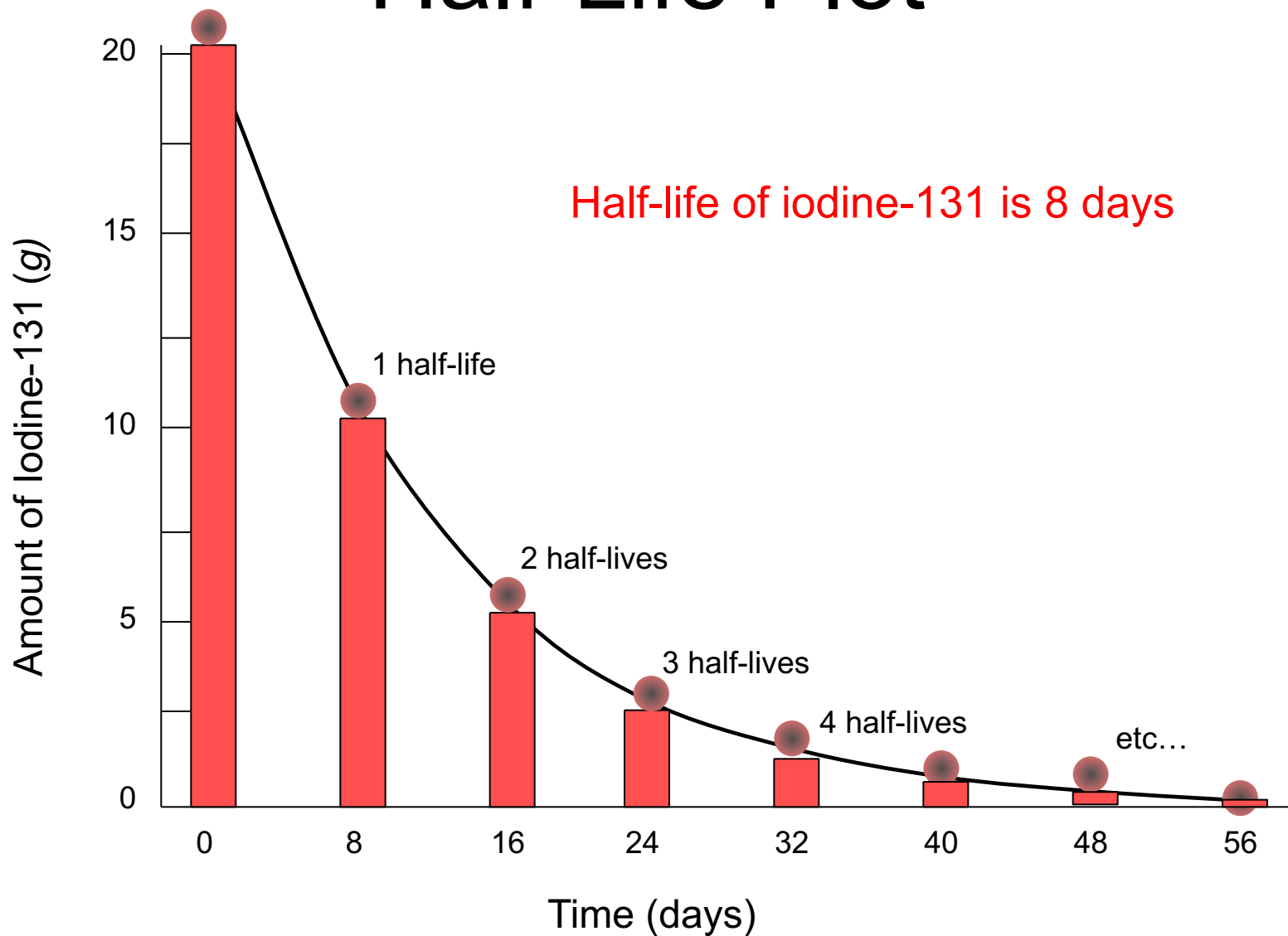


# Half-life of Radiation





# Half-Life Plot

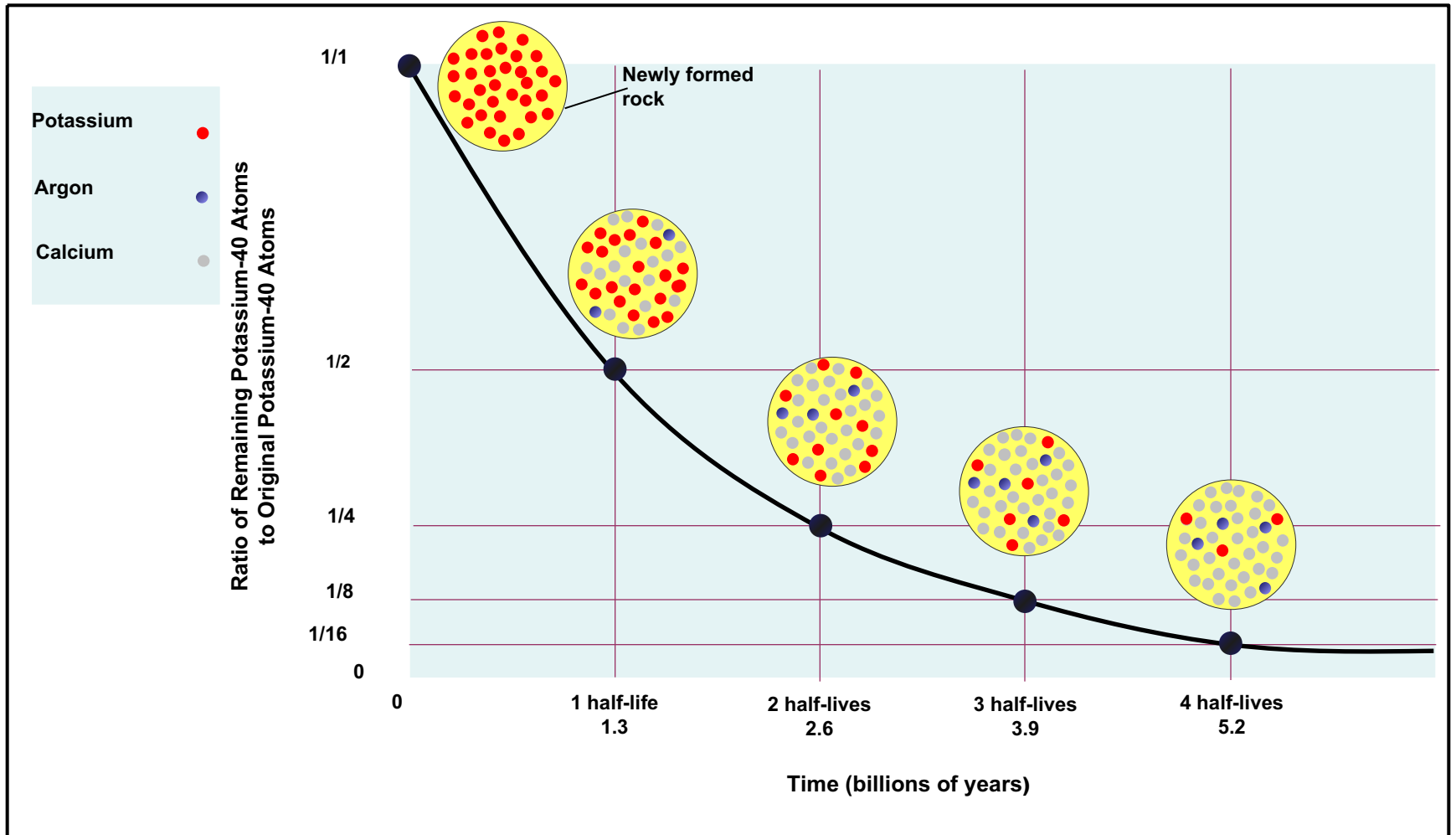


# Half-Life of Isotopes

Half-Life and Radiation of Some Naturally Occurring Radioisotopes		
Isotope	Half-Life	Radiation emitted
Carbon-14	$5.73 \times 10^3$ years	$\beta$
Potassium-40	$1.25 \times 10^9$ years	$\beta, \gamma$
Radon-222	3.8 days	$\alpha$
Radium-226	$1.6 \times 10^3$ years	$\alpha, \gamma$
Thorium-230	$7.54 \times 10^4$ years	$\alpha, \gamma$
Thorium-234	24.1 days	$\beta, \gamma$
Uranium-235	$7.0 \times 10^8$ years	$\alpha, \gamma$
Uranium-238	$4.46 \times 10^9$ years	$\alpha$

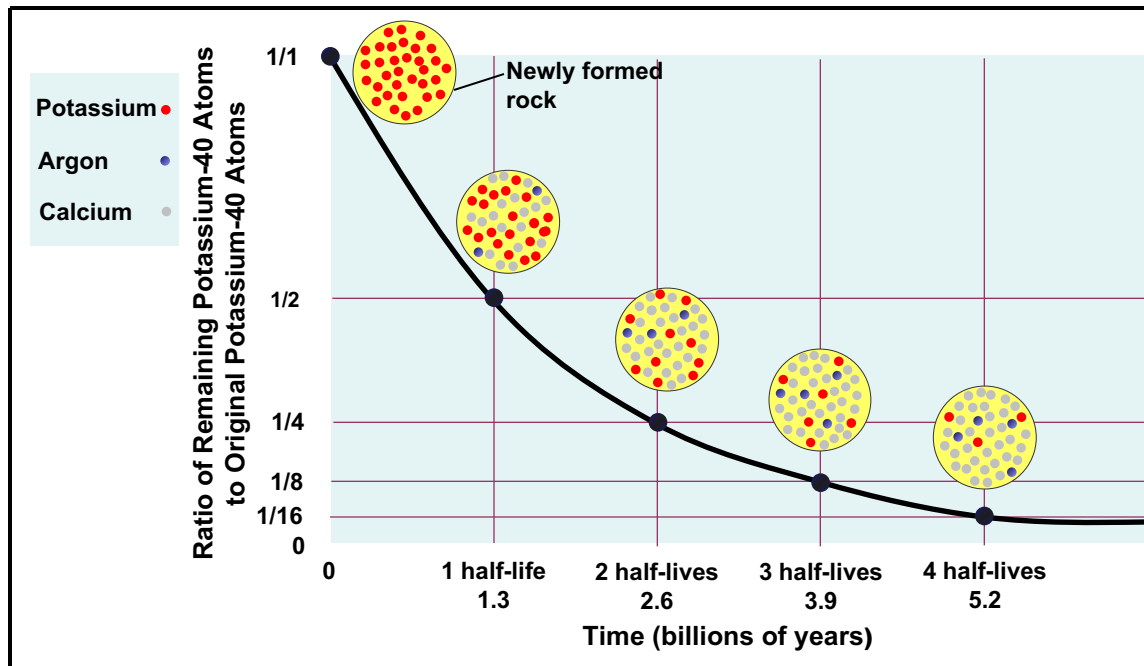


**Half-life ( $t_{1/2}$ )**  $1/2$   $1/4$   $1/8$   $1/16$



# Half-life ( $t_{1/2}$ )

- Time required for half the atoms of a radioactive nuclide to decay.
- Shorter half-life = less stable.



# How Much Remains?

After **one** half-life,  $\frac{1}{2}$  of the original atoms remain.

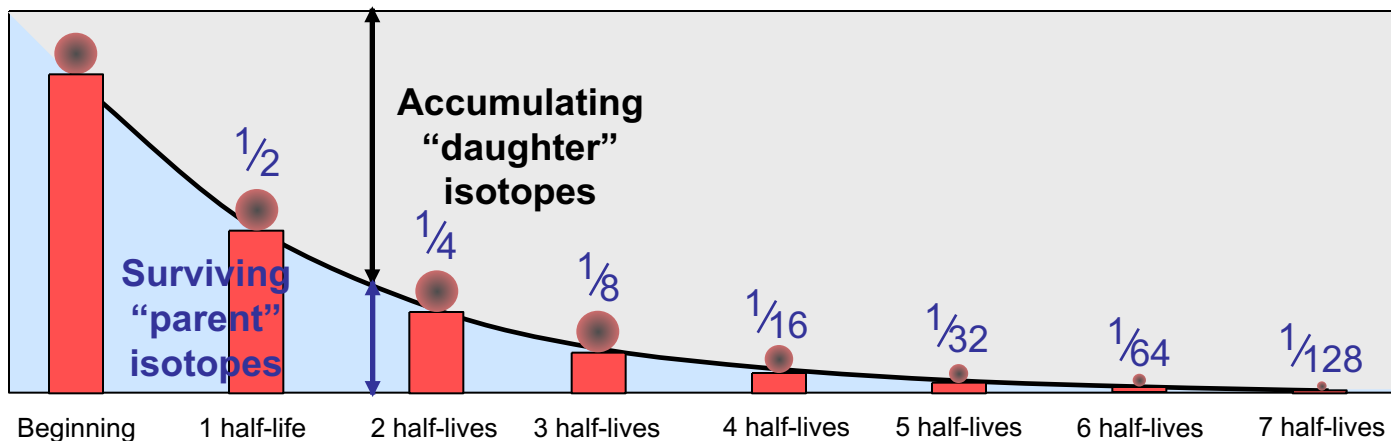
After **two** half-lives,  $\frac{1}{2} \times \frac{1}{2} = 1/(2^2) = \frac{1}{4}$  of the original atoms remain.

After **three** half-lives,  $\frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} = 1/(2^3) = \frac{1}{8}$  of the original atoms remain.

After **four** half-lives,  $\frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} = 1/(2^4) = \frac{1}{16}$  of the original atoms remain.

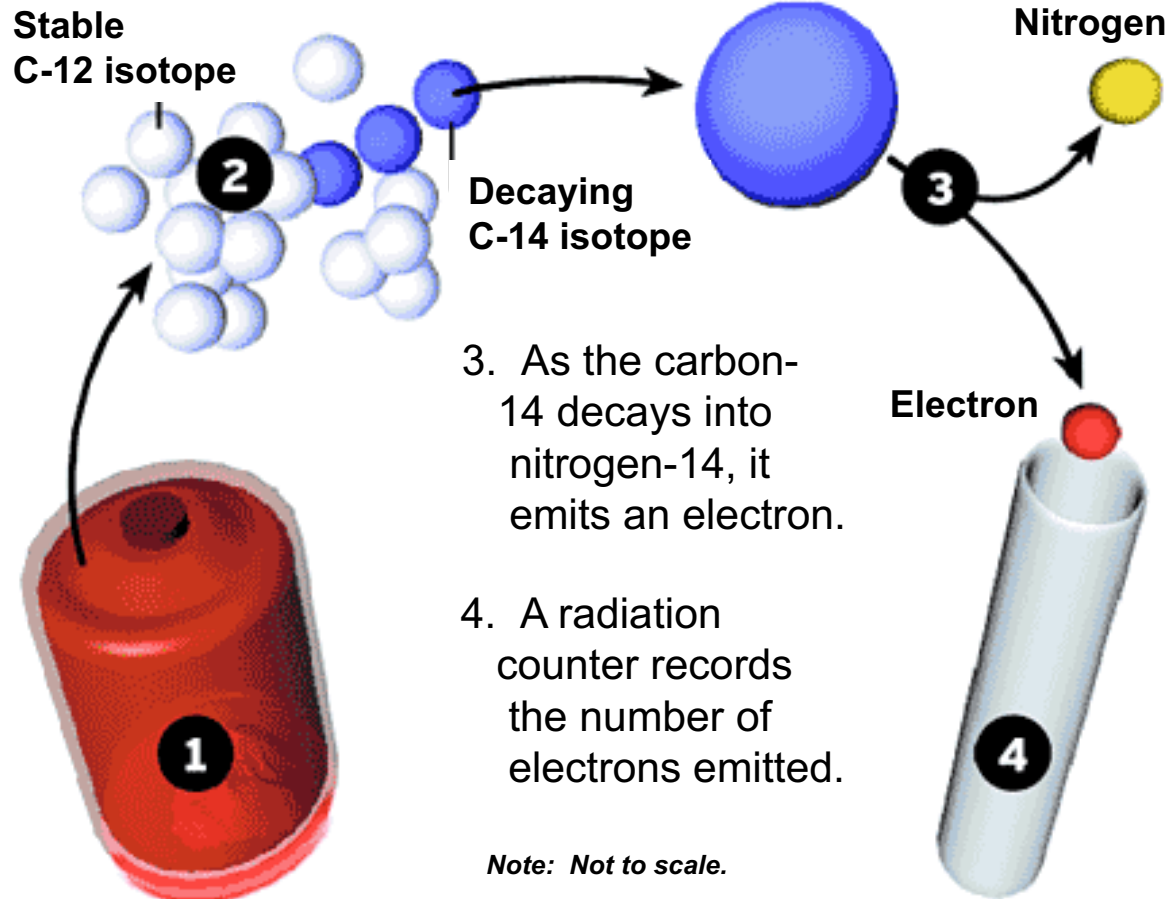
After **five** half-lives,  $\frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} = 1/(2^5) = \frac{1}{32}$  of the original atoms remain.

After **six** half-lives,  $\frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} = 1/(2^6) = \frac{1}{64}$  of the original atoms remain.



1. A small piece of fossil is burned in a special furnace.

2. The burning creates carbon dioxide gas comprised of carbon-12 isotopes and carbon-14 isotopes.



The iodine-131 nuclide has a half-life of 8 days. If you originally have a 625-g sample, after 2 months you will have approximately?

- a. 40 g
- b. 20 g
- c. 10 g
- ★ d. 5 g
- e. less than 1 g

$$N = N_o \left(\frac{1}{2}\right)^n$$

N = amount remaining

N<sub>o</sub> = original amount

n = # of half-lives

$$N = (625 \text{ g}) \left(\frac{1}{2}\right)^{7.5}$$

$$N = 3.45 \text{ g}$$

Data Table: Half-life Decay

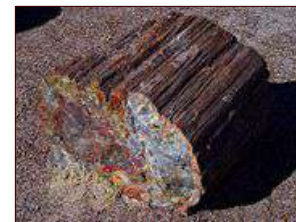
~ Amount	Time	# Half-Life
625 g	0 d	0
312 g	8 d	1
156 g	16 d	2
78 g	24 d	3
39 g	32 d	4
20 g	40 d	5
10 g	48 d	6
5 g	56 d	7
2.5 g	64 d	8
1.25 g	72 d	9

Assume 30 days = 1 month

$$\frac{60 \text{ days}}{8 \text{ days}} = 7.5 \text{ half-lives}$$

Given that the half-life of carbon-14 is 5730 years, consider a sample of fossilized wood that, when alive, would have contained 24 g of carbon-14. It now contains 1.5 g of carbon-14.

How old is the sample?



$$\ln \frac{N}{N_0} = -k t$$

$$t_{1/2} = \frac{0.693}{k}$$

$$5730 \text{ y} = \frac{0.693}{k}$$

$$k = 1.209 \times 10^{-4}$$

Data Table: Half-life Decay

Amount	Time	# Half-Life
24 g	0 y	0
12 g	5,730 y	1
6 g	11,460 y	2
3 g	17,190 y	3
1.5 g	22,920 y	4

$$\ln \frac{1.5 \text{ g}}{24 \text{ g}} = - (1.209 \times 10^{-4}) t$$

$$t = 22,933 \text{ years}$$



# Half-Life Practice Calculations

- The half-life of carbon-14 is 5730 years. If a sample originally contained 3.36 g of C-14, how much is present after 22,920 years? 0.21 g C-14
- Gold-191 has a half-life of 12.4 hours. After one day and 13.2 hours, 10.6 g of gold-191 remains in a sample. How much gold-191 was originally present in the sample? 84.8 g Au-191
- There are 3.29 g of iodine-126 remaining in a sample originally containing 26.3 g of iodine-126. The half-life of iodine-126 is 13 days. How old is the sample? 39 days old
- A sample that originally contained 2.5 g of rubidium-87 now contains 1.25 g. The half-life of rubidium-87 is  $6 \times 10^{10}$  years. How old is the sample? Is this possible? Why or why not?  $6 \times 10^{10}$  years  
(60,000,000,000 billions years old)  
What is the age of Earth???

Demo: Try to cut a string in half seven times (if it begins your arm's length).

The half-life of carbon-14 is 5730 years. If a sample originally contained 3.36 g of C-14, how much is present after 22,920 years?

$$t_{1/2} = 5730 \text{ years}$$

$$n = \frac{22,920 \text{ years}}{5,730 \text{ years}}$$

$$n = 4 \text{ half-lives}$$

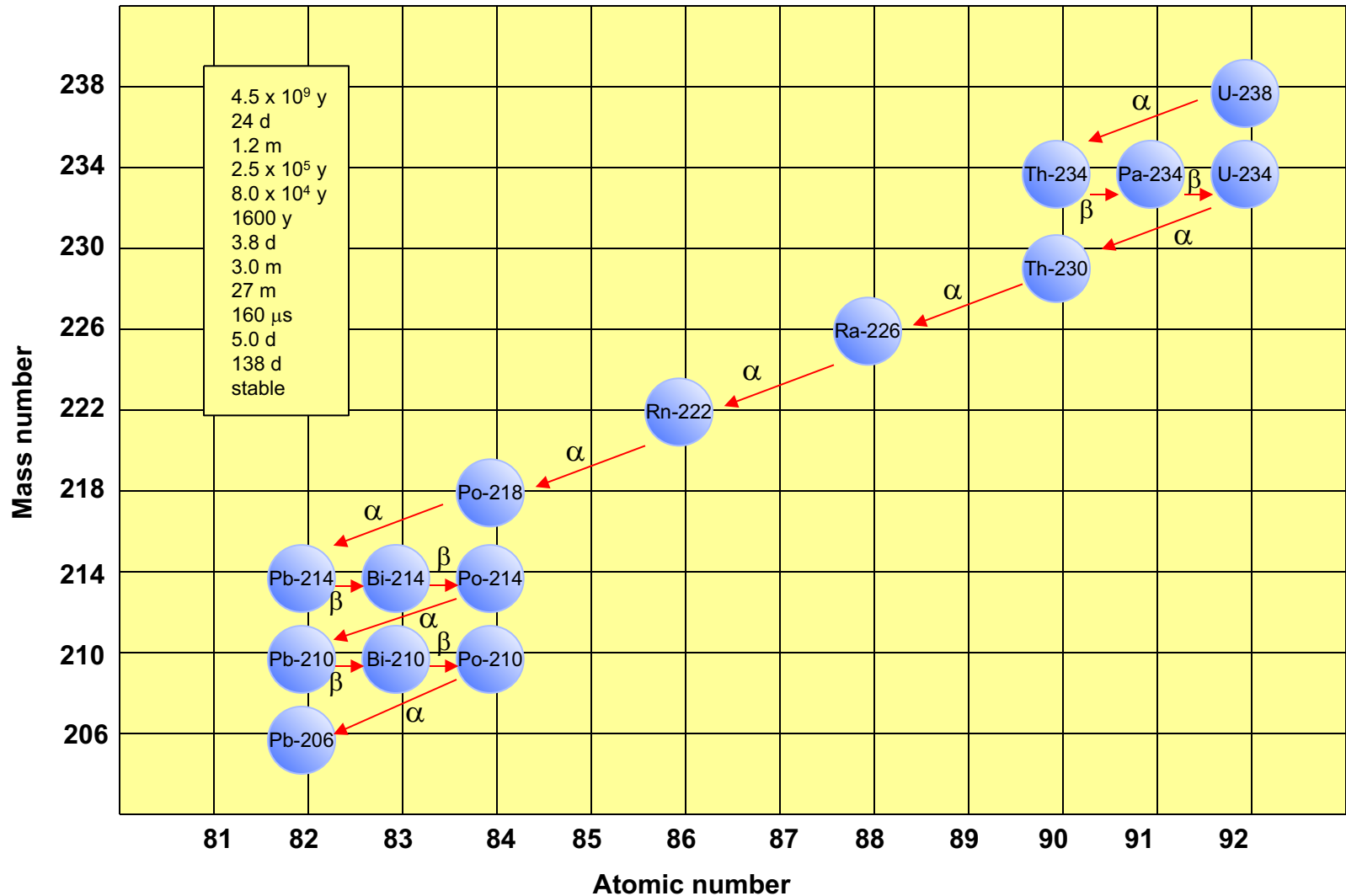
$$(\# \text{ of half-lives})(\text{half-life}) = \text{age of sample}$$

$$(4 \text{ half-lives})(5730 \text{ years}) = \text{age of sample}$$

$$22,920 \text{ years}$$

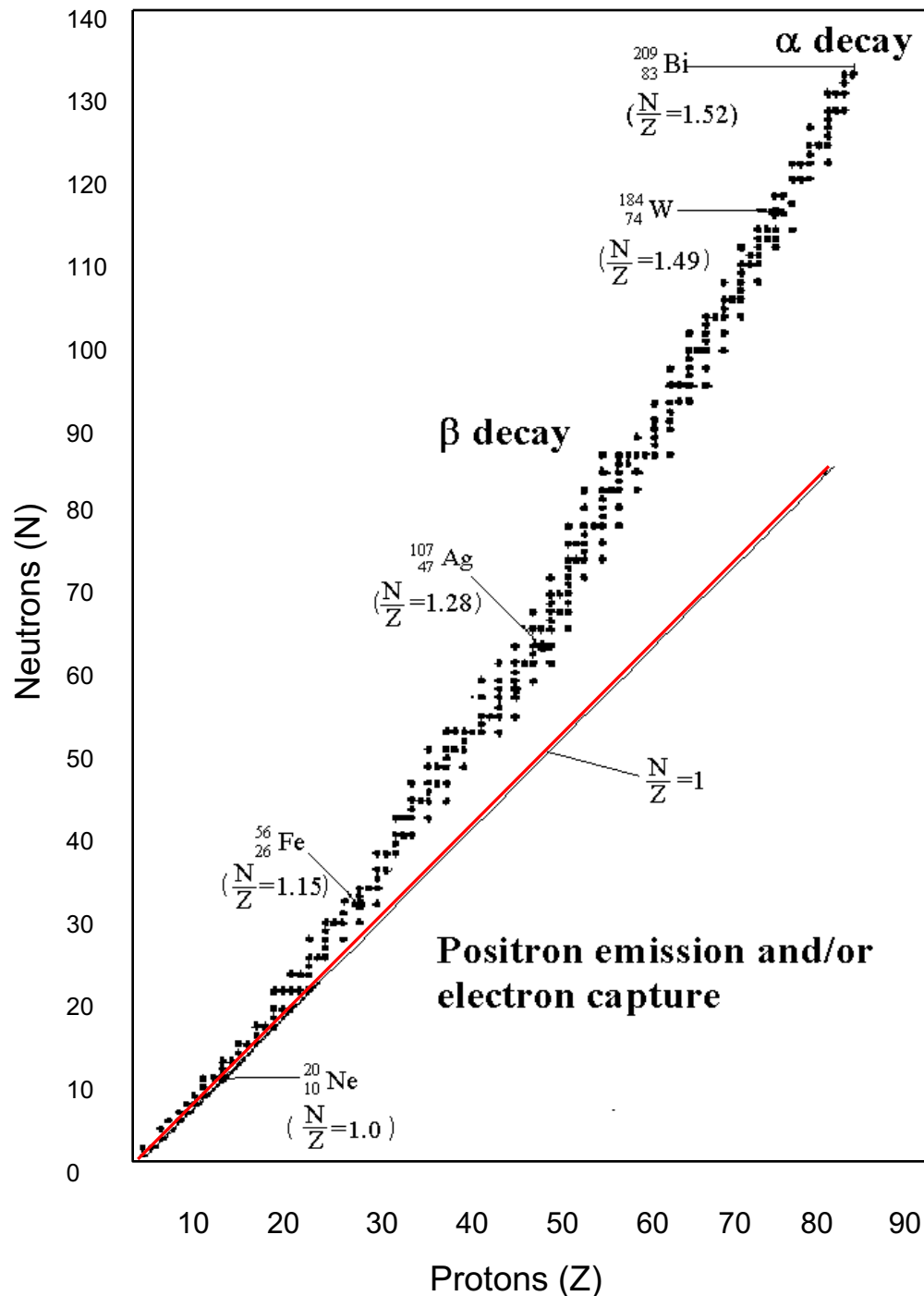


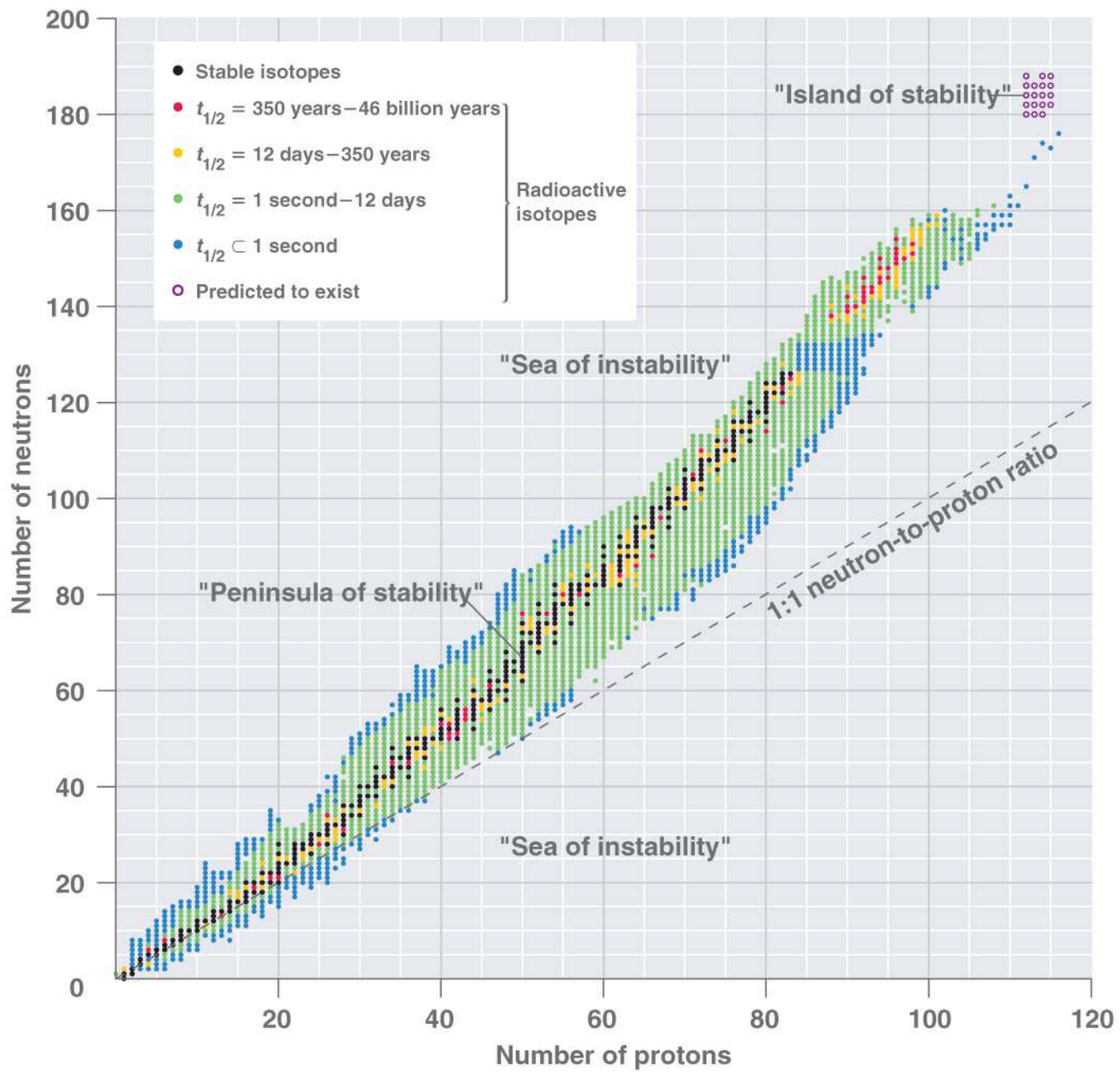
# Uranium Radioactive Decay



# Nuclear Stability

Decay will occur in such a way as to return a nucleus to the band (line) of stability.

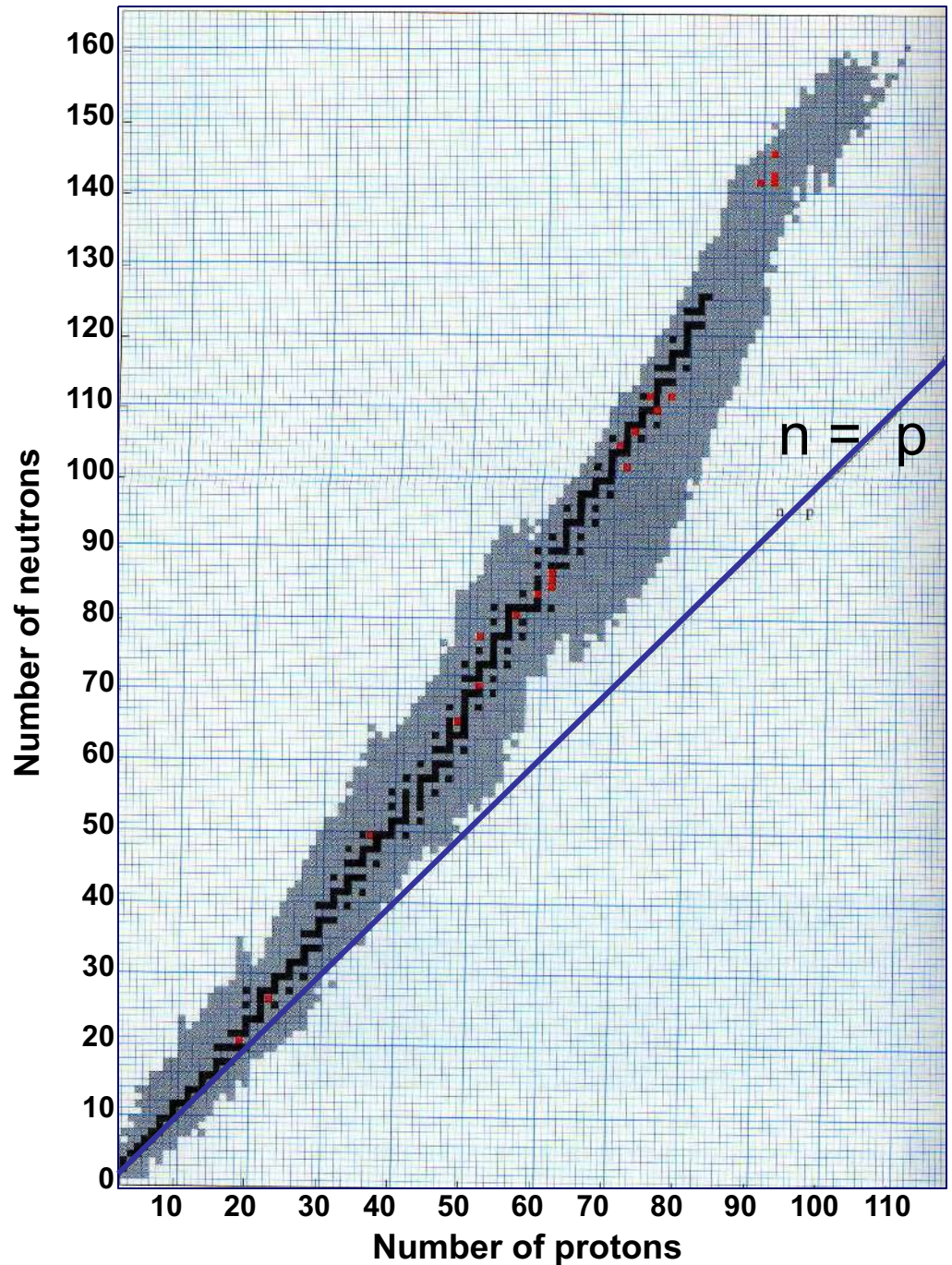


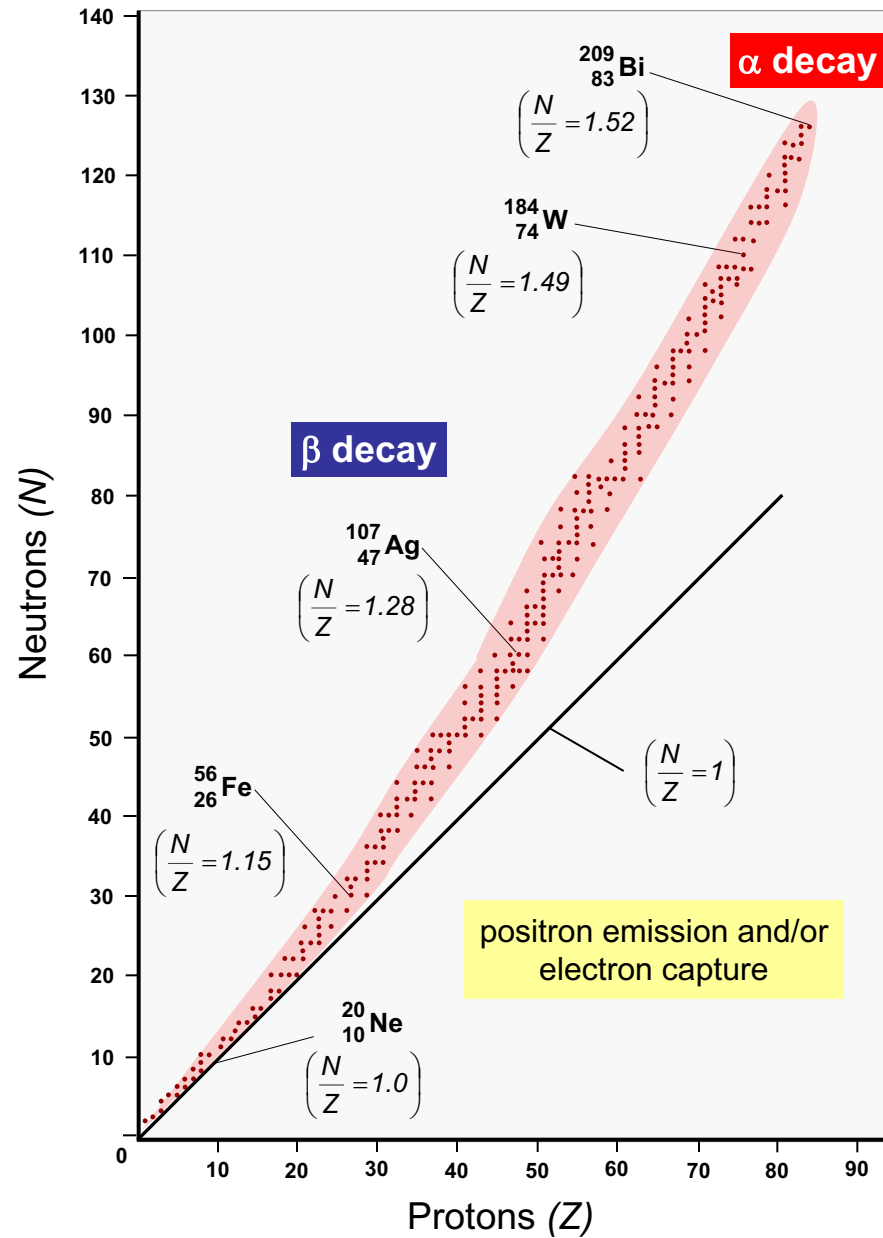




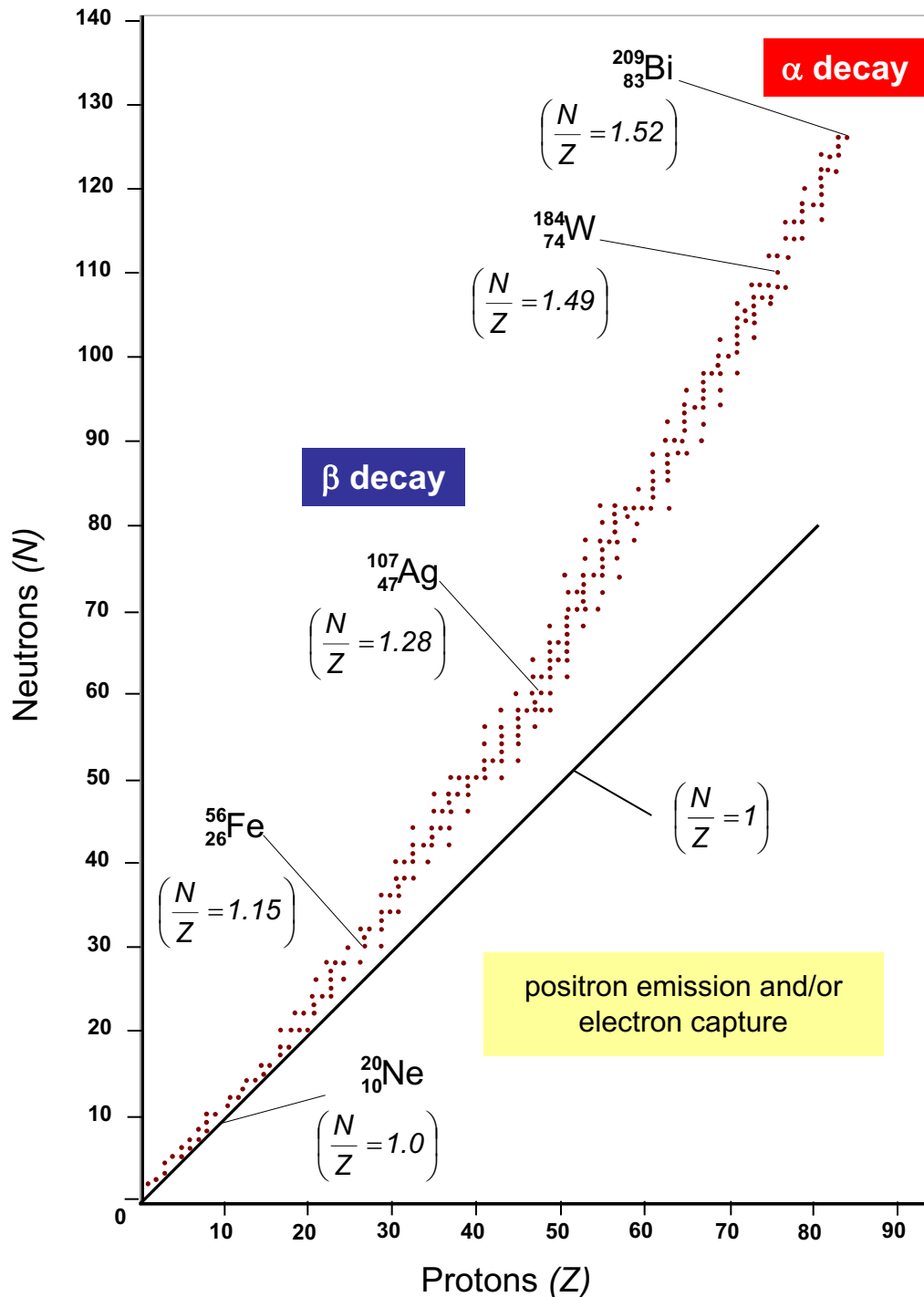
# Band of Stability

- Stable nuclides
- Naturally occurring radioactive nuclides
- Other known nuclides





# Nuclear Stability



Decay will occur in such a way as to return a nucleus to the band (line) of stability.

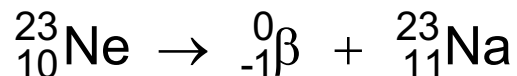


# Half-Lives of Some Isotopes of Carbon

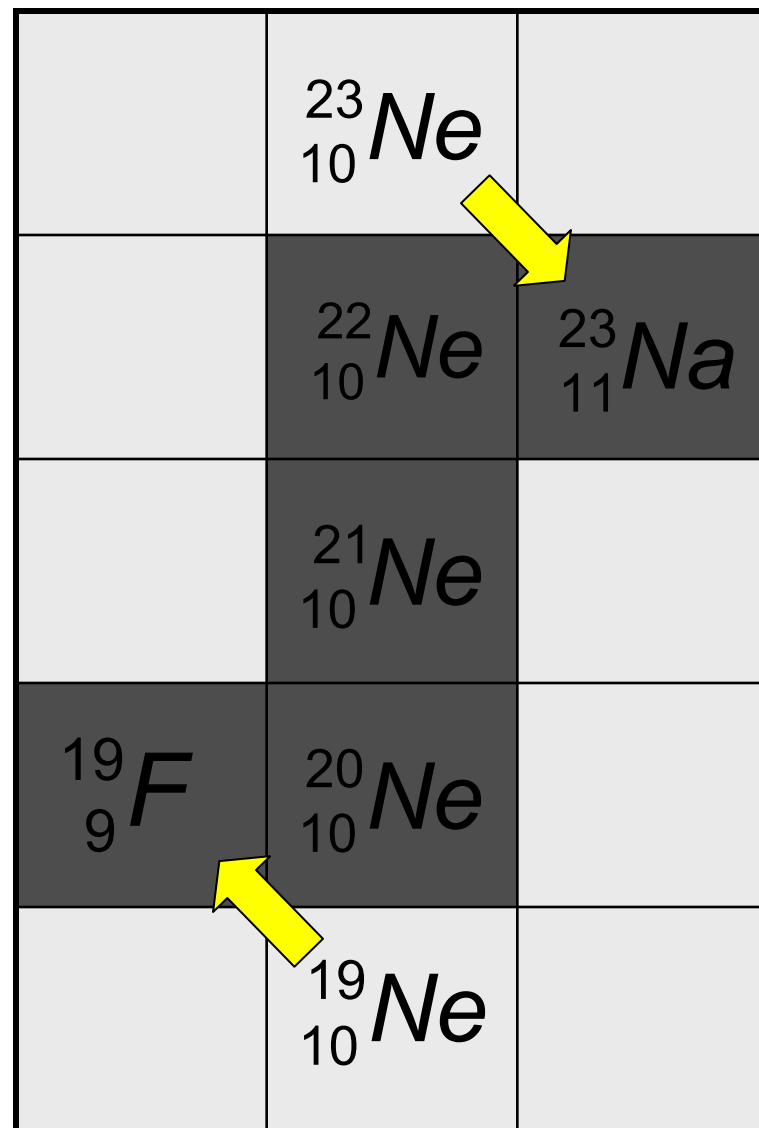
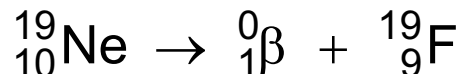
Nuclide	Half-Life
Carbon-9	0.127 s
Carbon-10	19.3 s
Carbon-11	10.3 m
Carbon-12	Stable
Carbon-13	Stable
Carbon-14	5715 y
Carbon-15	2.45 s
Carbon-16	0.75 s

# Enlargement of part of band of stability around Neon

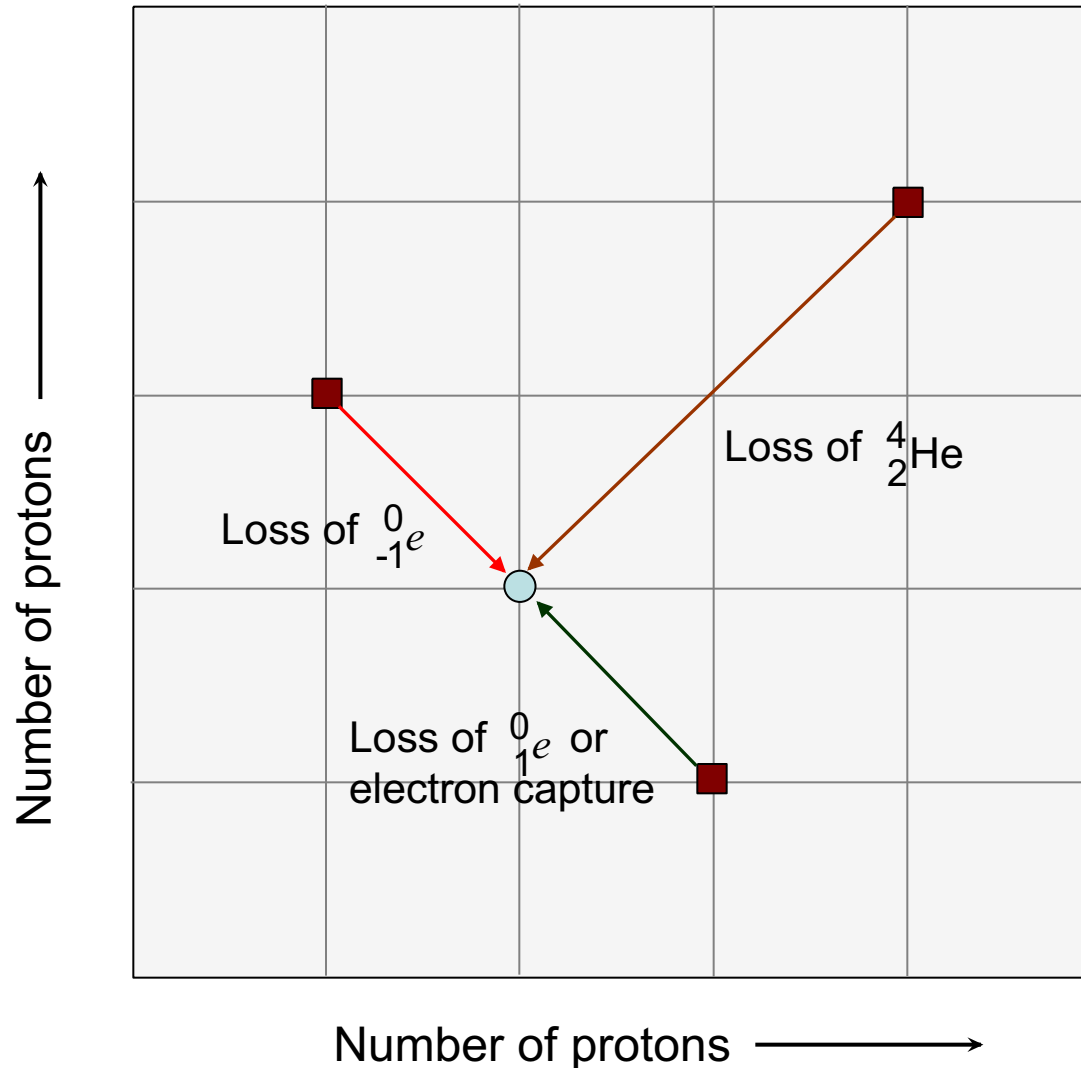
moves into band of stability by beta decay.



moves into band of stability by positron emission. Electron capture would also move into the band of stability.

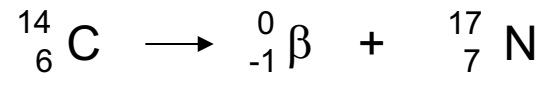
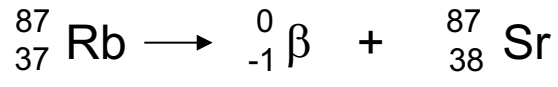
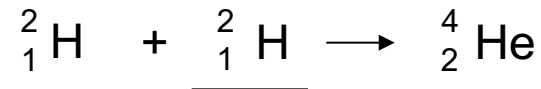
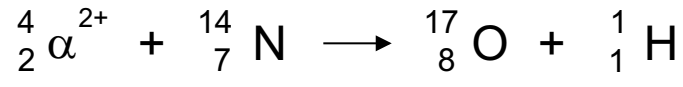
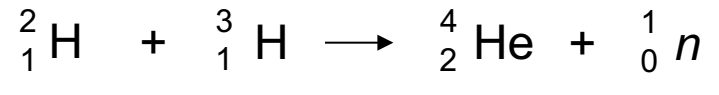
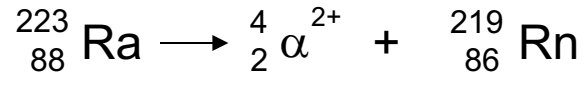


# Effects of Radioactive Emissions on Proton and Neutrons

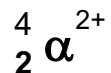


# Nuclear Decay

“absorption”, “bombardment” vs. “production”, “emission”



Alpha



Beta



Positron



Gamma



neutron



proton





# Units Used in Measurement of Radioactivity



## Units

## Measurements



Curie (C)

radioactive decay



Becquerel (Bq)

radioactive decay



Roentgens (R)

exposure to ionizing radiation

Rad (rad)

energy absorption caused by ionizing radiation

Rem (rem)

biological effect of the absorbed dose in humans

# Effects of Instantaneous Whole-Body Radiation Doses on People



Alexander Litvinenko

Dose, Sv (rem)	Effect
$\geq 10$ (1000)	Death within 24 h from destruction of the neurological system.
7.5 (750)	Death within 4-30 d from gastrointestinal bleeding.
1.5 – 7.5 (150 – 750)	Intensive hospital care required for survival. At the higher end of range, death through infection resulting from destruction of white-blood cell-forming organs usually takes place 4 – 8 weeks after accident. Those surviving this period usually recover.
$< 0.5$ (50)	Only proven effect is decrease in white blood cell count.

The intensity of radiation is proportional to  $1/d^2$ , where  $d$  is the distance from the source.



# Alpha, Beta, Positron Emission

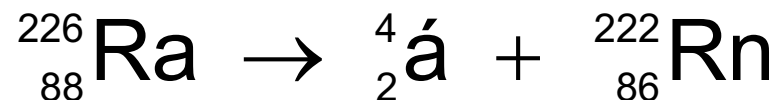
Examples of Nuclear Decay Processes		
$\alpha$ emission (alpha)	$\beta^-$ emission (beta)	$\beta^+$ emission (positron)
${}_{92}^{238}\text{U} \rightarrow {}_2^4\text{He} + {}_{90}^{234}\text{Th}$	${}_{12}^{27}\text{Mg} \rightarrow {}_{-1}^0\text{e} + {}_{13}^{27}\text{Al}$	${}_{8}^{14}\text{O} \rightarrow {}_{+1}^0\text{e} + {}_{7}^{14}\text{N}$
${}_{90}^{230}\text{Th} \rightarrow {}_2^4\text{He} + {}_{88}^{226}\text{Ra}$	${}_{16}^{35}\text{S} \rightarrow {}_{-1}^0\text{e} + {}_{17}^{35}\text{Cl}$	${}_{17}^{32}\text{Cl} \rightarrow {}_{+1}^0\text{e} + {}_{16}^{32}\text{S}$
${}_{88}^{226}\text{Ra} \rightarrow {}_2^4\text{He} + {}_{86}^{222}\text{Rn}$	${}_{19}^{40}\text{K} \rightarrow {}_{-1}^0\text{e} + {}_{20}^{40}\text{Ca}$	${}_{8}^{14}\text{O} \rightarrow {}_{+1}^0\text{e} + {}_{7}^{14}\text{N}$

Although beta emission involves electrons, those electrons come from the nucleus. Within the nucleus, a neutron decays into a proton and an electron. The electron is emitted, leaving behind a proton to replace the neutron, thus transforming the element. (A neutrino is also produced and emitted in the process.)

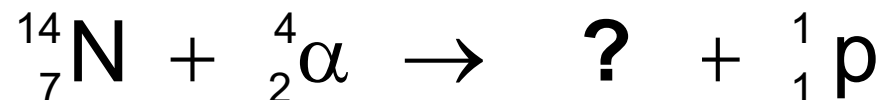


# Nuclear Reactions

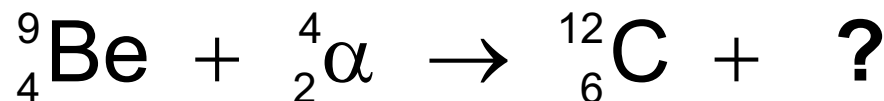
First recognized natural transmutation of an element (Rutherford and Soddy, 1902)



First artificial transmutation of an element (Rutherford, 1919)



Discovery of the neutron (Chadwick, 1932)



Discovery of nuclear fission (Otto Hahn and Fritz Strassman, 1939)



# Preparation of Transuranium Elements

<i>Atomic Number</i>	<i>Name</i>	<i>Symbol</i>	<i>Year Discovered</i>	<i>Reaction</i>
93	Neptunium	Np	1940	${}_{92}^{238}\text{U} + {}_0^1\text{n} \rightarrow {}_{93}^{239}\text{Np} + {}_{-1}^0\text{e}$
94	Plutonium	Pu	1940	${}_{92}^{238}\text{U} + {}_1^2\text{H} \rightarrow {}_{93}^{238}\text{Np} + 2{}_0^1\text{n}$ ${}_{93}^{238}\text{Np} \rightarrow {}_{94}^{238}\text{Pu} + {}_{-1}^0\text{e}$
95	Americium	Am	1944	${}_{94}^{239}\text{Pu} + {}_0^1\text{n} \rightarrow {}_{95}^{240}\text{Am} + {}_{-1}^0\text{e}$
96	Curium	Cm	1945	${}_{94}^{239}\text{Pu} + {}_2^4\text{He} \rightarrow {}_{96}^{242}\text{Cm} + {}_0^1\text{n}$
97	Berkelium	Bk	1949	${}_{95}^{241}\text{Am} + {}_2^4\text{He} \rightarrow {}_{97}^{243}\text{Bk} + 2{}_0^1\text{n}$
98	Californium	Cf	1950	${}_{96}^{242}\text{Cm} + {}_2^4\text{He} \rightarrow {}_{98}^{245}\text{Cf} + {}_0^1\text{n}$

# Preparation of Transuranium Elements

<i>Atomic Number</i>	<i>Name</i>	<i>Symbol</i>	<i>Year Discovered</i>	<i>Reaction</i>
93	Neptunium	Np	1940	${}_{92}^{238}\text{U} + {}_0^1\text{n} \rightarrow {}_{93}^{239}\text{Np} + {}_{-1}^0\text{e}$
94	Plutonium	Pu	1940	${}_{92}^{238}\text{U} + {}_1^2\text{H} \rightarrow {}_{93}^{238}\text{Np} + 2{}_0^1\text{n}$ ${}_{93}^{238}\text{Np} \rightarrow {}_{94}^{238}\text{Pu} + {}_{-1}^0\text{e}$
95	Americium	Am	1944	${}_{94}^{239}\text{Pu} + {}_0^1\text{n} \rightarrow {}_{95}^{240}\text{Am} + {}_{-1}^0\text{e}$
96	Curium	Cm	1945	${}_{94}^{239}\text{Pu} + {}_2^4\text{He} \rightarrow {}_{96}^{242}\text{Cm} + {}_0^1\text{n}$
97	Berkelium	Bk	1949	${}_{95}^{241}\text{Am} + {}_2^4\text{He} \rightarrow {}_{97}^{243}\text{Bk} + 2{}_0^1\text{n}$
98	Californium	Cf	1950	${}_{96}^{242}\text{Cm} + {}_2^4\text{He} \rightarrow {}_{98}^{245}\text{Cf} + {}_0^1\text{n}$

# Additional Transuranium Elements

99	Einsteinium	Es	1952	
100	Fermium	Fm	1952	
101	Mendelevium	Md	1955	
102	Nobelium	Nb	1958	
103	Lawrencium	Lr	1961	
104	Rutherfordium	Rf	1964	
105	Dubnium	Db	1970	
106	Seaborgium	Sg	1974	
107	Bohrium	Bh	1981	
108	Hassium	Hs	1984	
109	Meitnerium	Mt	1988	
110	Darmstadtium	Ds	1994	
111	Unununium	Uun	1994	
112	Ununbium	Uub	1996	
114		Uuq	1999	(Russia)
116			2002	(Russia)
118			2006	

# CHAPTER 22

---

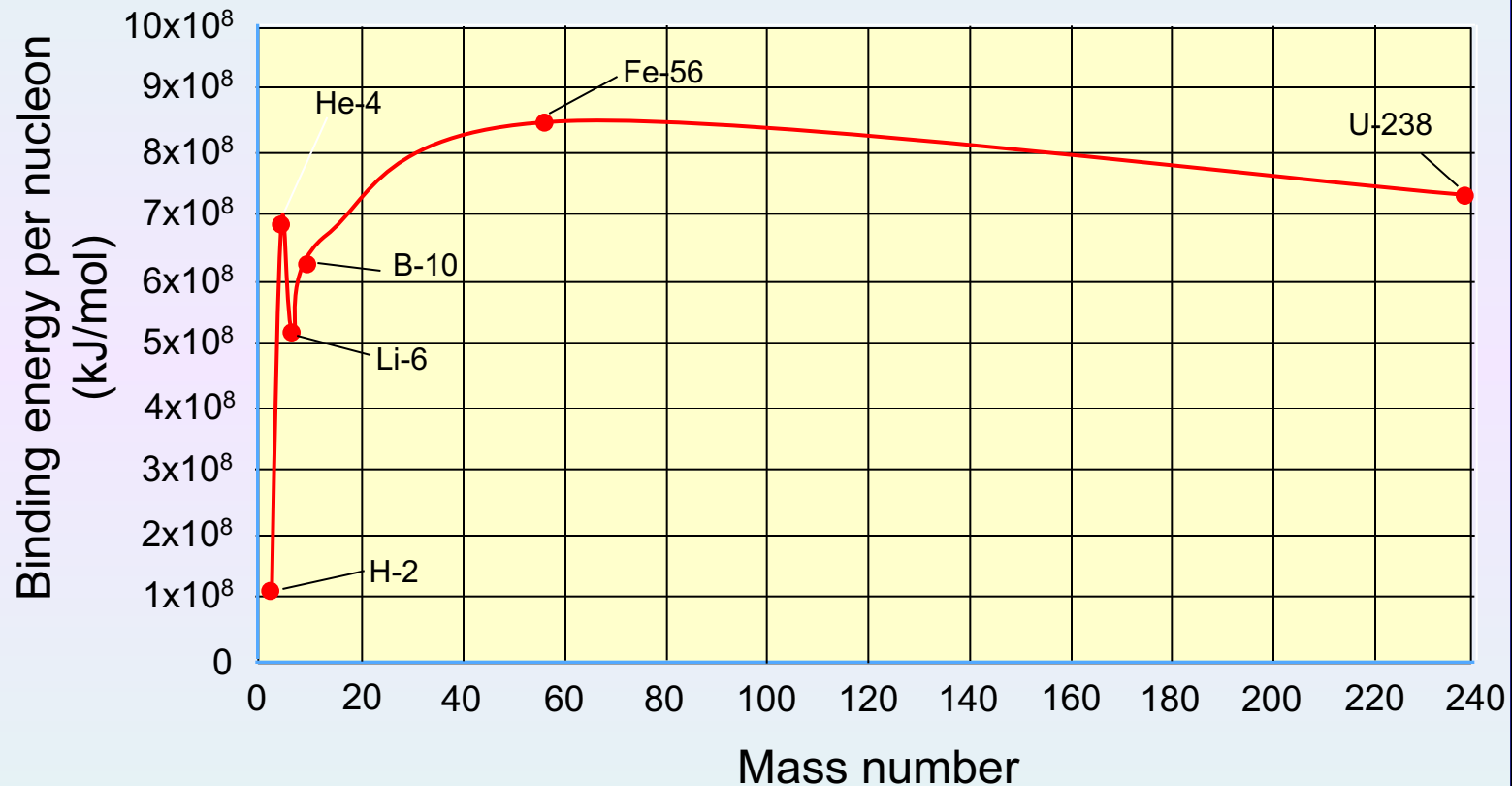
# Nuclear Chemistry



## **I. The Nucleus**

(p. 701 - 704)

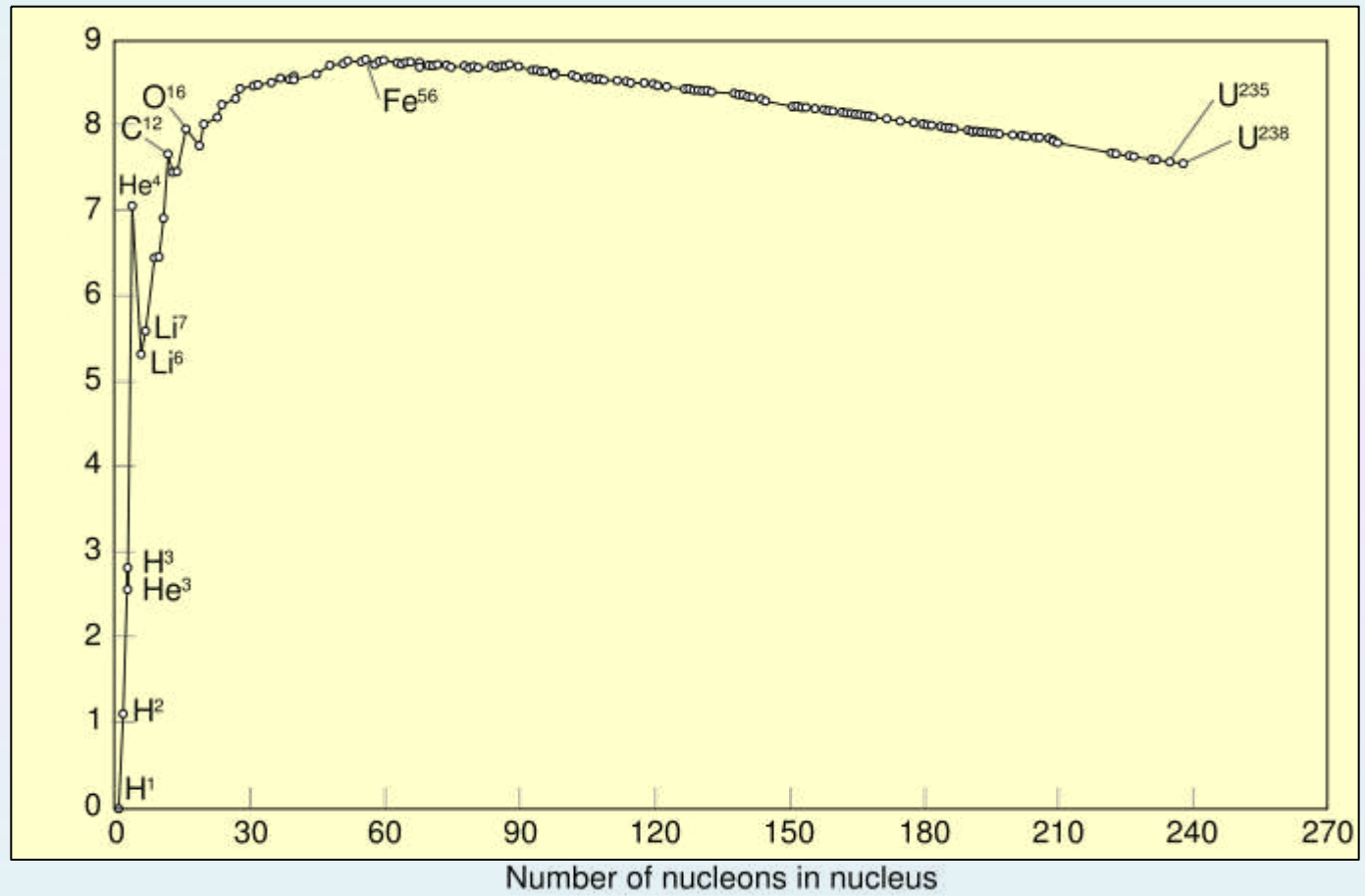
# Nuclear Binding Energy



Unstable nuclides are radioactive and undergo radioactive decay.

# Nuclear Binding Energy

Average binding energy per nucleon (MeV)



Unstable nuclides are radioactive and undergo radioactive decay.

# CHAPTER 22

---

# Nuclear Chemistry



## **II. Radioactive Decay**

(p. 705 - 712)

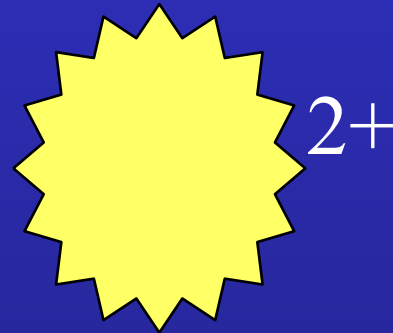




# Types of Radiation

★ Alpha particle ( $\alpha$ )

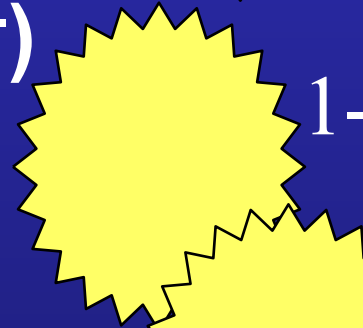
◆ helium nucleus



paper

★ Beta particle ( $\beta^-$ )

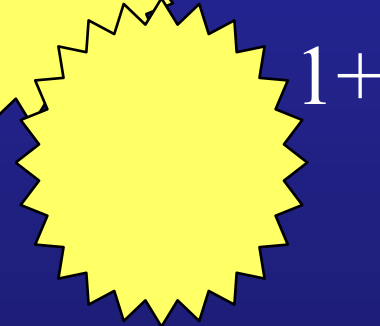
◆ electron



lead

★ Positron ( $\beta^+$ )

◆ positron



★ Gamma ( $\gamma$ )

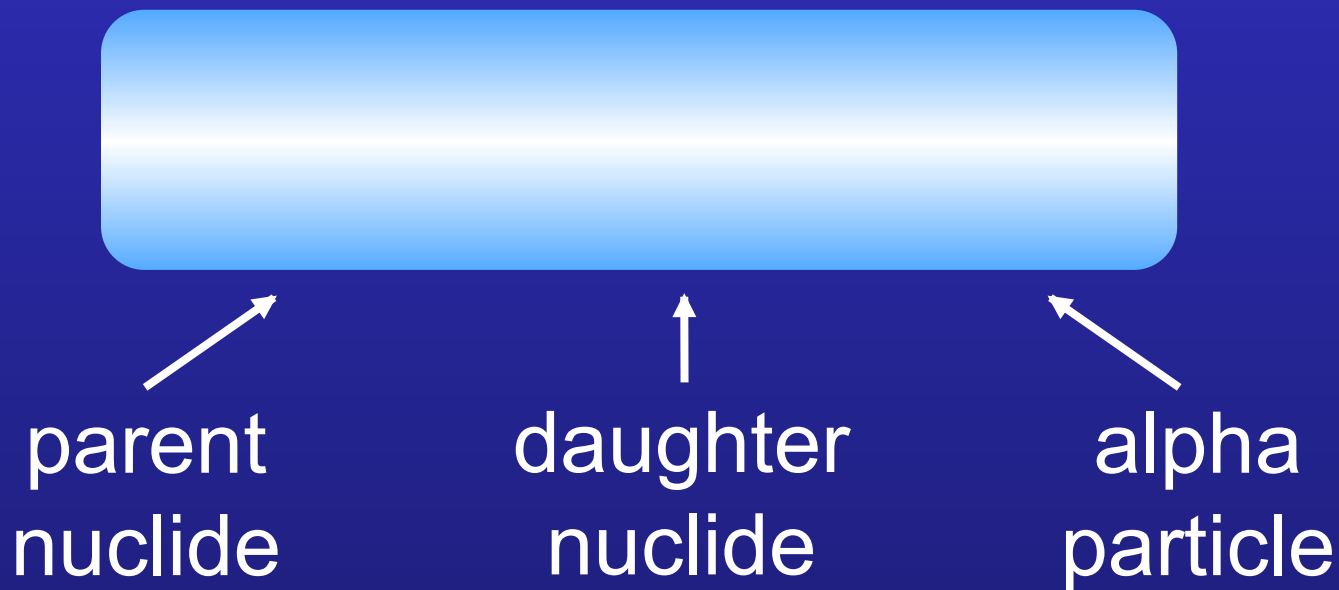
◆ high-energy photon



concrete

# Nuclear Decay

## ★ Alpha Emission



**Numbers must balance!!**

# Nuclear Decay

## ★ Beta Emission



electron

## ★ Positron Emission

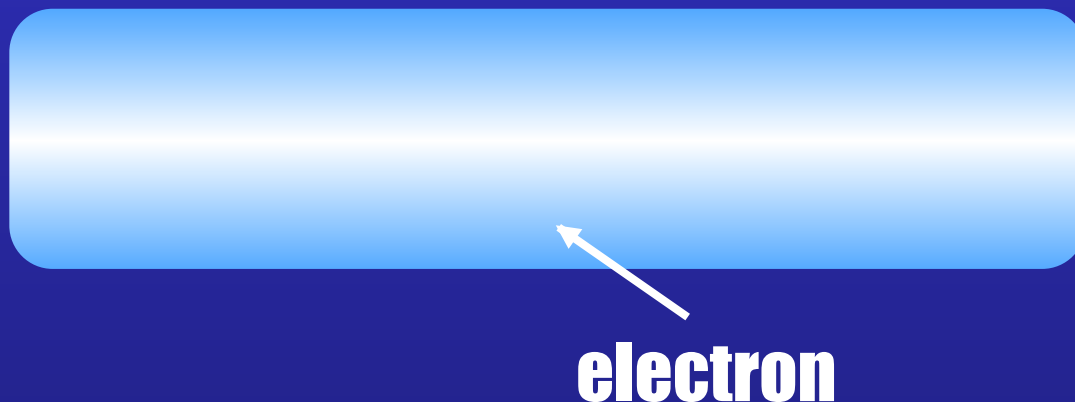


positron



# Nuclear Decay

## ★ Electron Capture



## ★ Gamma Emission

- ◆ Usually follows other types of decay.

## ★ Transmutation

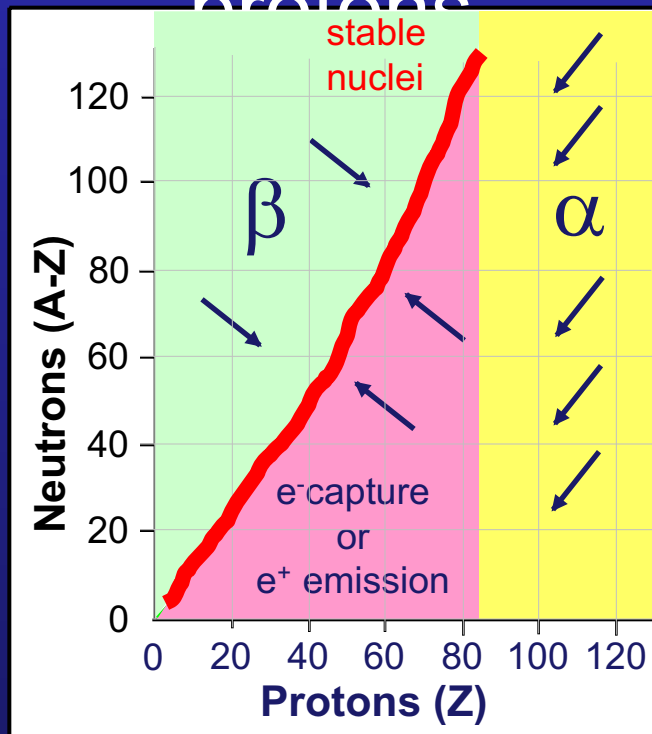
- ◆ One element becomes another.



# Nuclear Decay

## ★ Why nuclides decay...

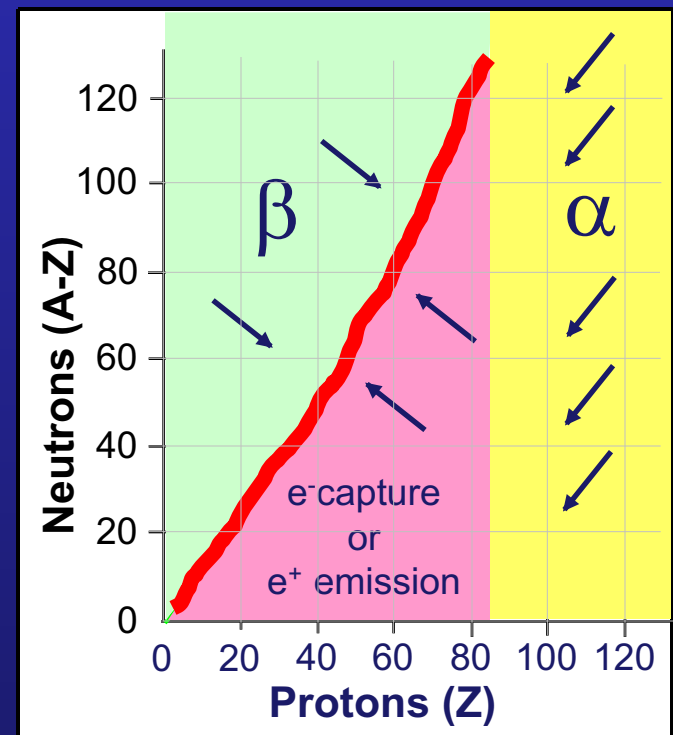
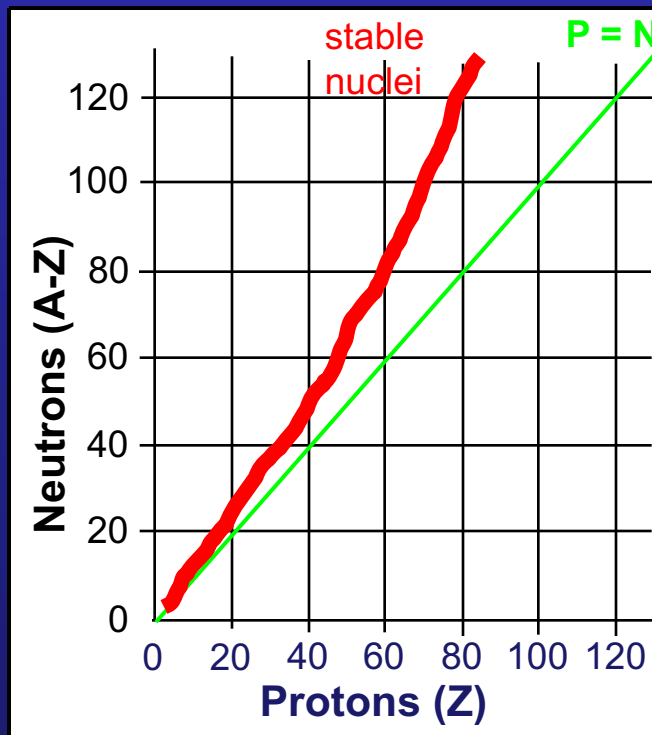
- ◆ need stable ratio of neutrons to protons



# Nuclear Decay

## ★ Why nuclides decay...

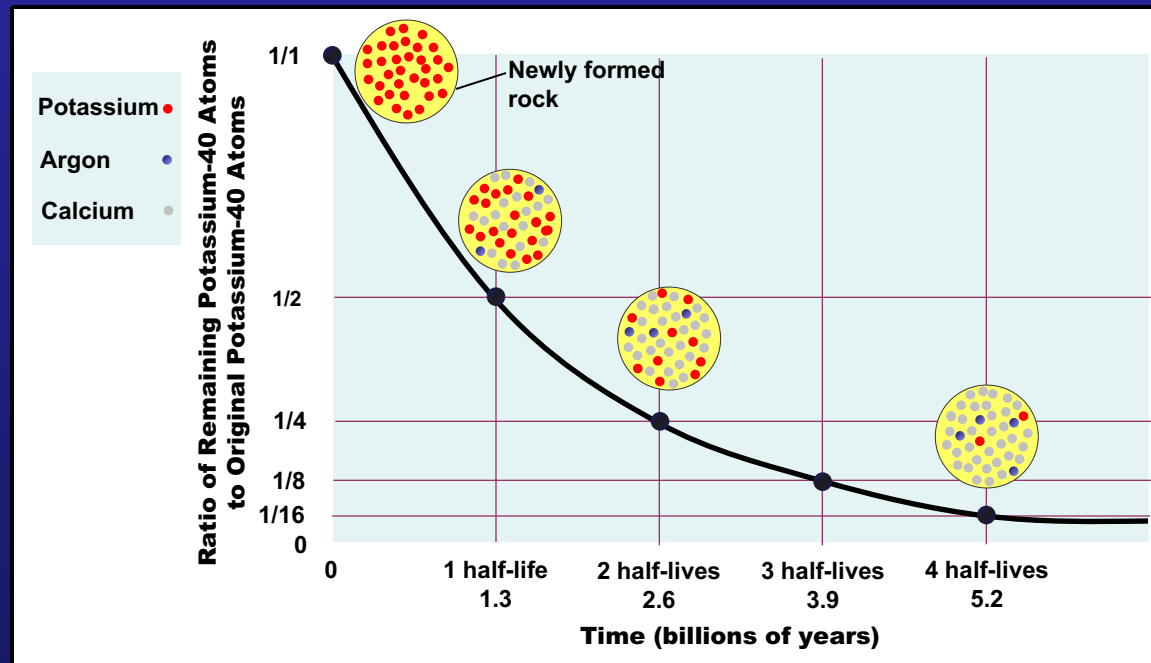
- ◆ need stable ratio of neutrons to protons



# Half-life

## ★ Half-life ( $t_{1/2}$ )

- ◆ Time required for half the atoms of a radioactive nuclide to decay.
- ◆ Shorter half-life = less stable.



# Half-life

---



$m_f$ : final mass

$m_i$ : initial mass

$n$ : # of half-lives





# Half-life

- ★ Fluorine-21 has a half-life of 5.0 seconds. If you start with 25 g of fluorine-21, how many grams would remain after 60.0 s?

GIVEN:

$$t_{1/2} = 5.0 \text{ s}$$

$$m_i = 25 \text{ g}$$

$$m_f = ?$$

$$\text{total time} = 60.0 \text{ s}$$

$$n = 60.0\text{s} \div 5.0\text{s} = 12$$

WORK:

$$m_f = m_i \left(\frac{1}{2}\right)^n$$

$$m_f = (25 \text{ g})(0.5)^{12}$$

$$\mathbf{m_f = 0.0061 \text{ g}}$$

# CHAPTER 22

---

# Nuclear Chemistry



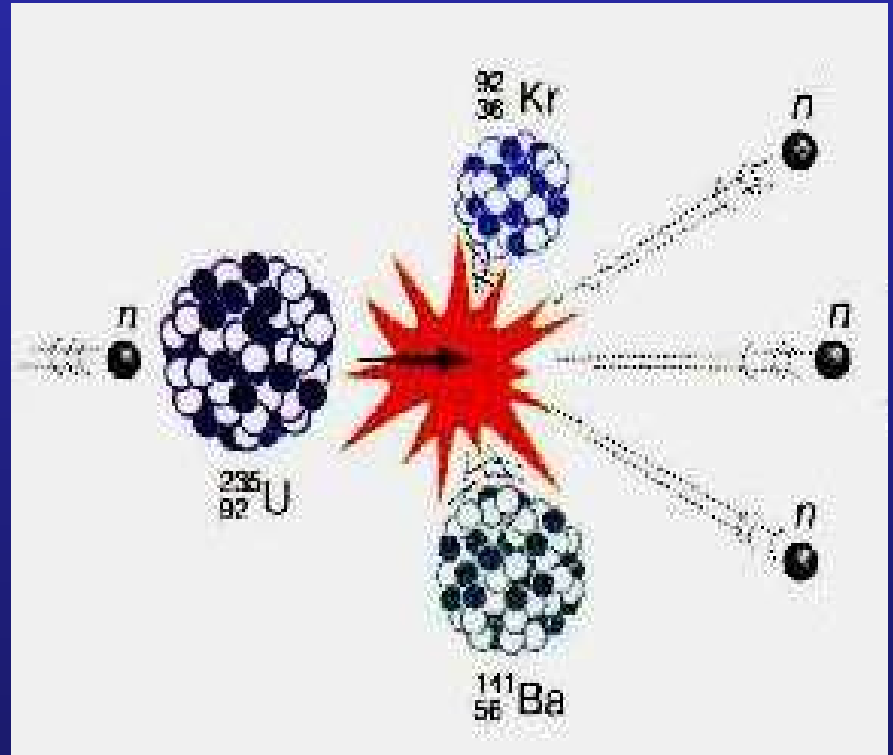
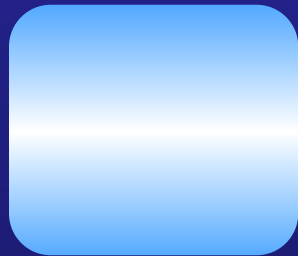
## **III. Fission & Fusion**

(p. 717 – 719)



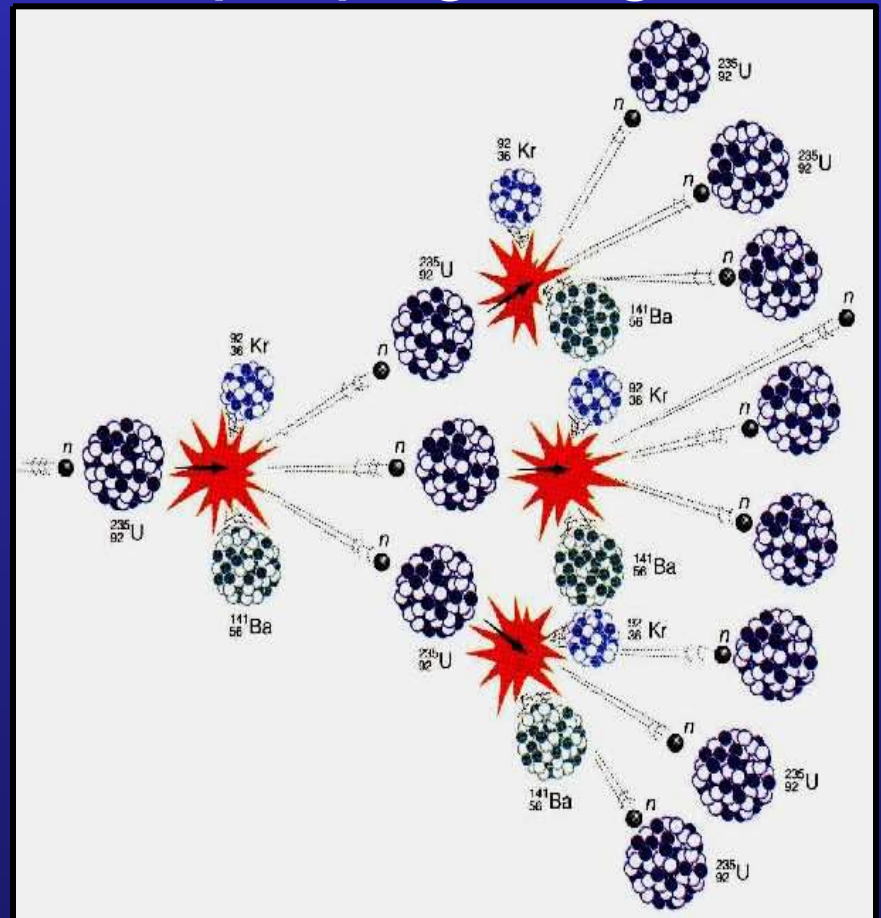
# Fission

- ✦ splitting a nucleus into two or more smaller nuclei
- ✦ 1 g of  $^{235}\text{U}$  = 3 tons of coal



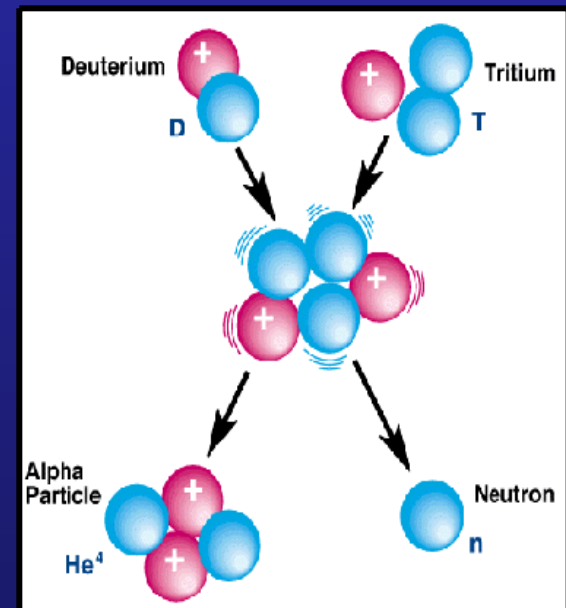
# Fission

- ★ chain reaction - self-propagating reaction
- ★ critical mass - mass required to sustain a chain reaction



# Fusion

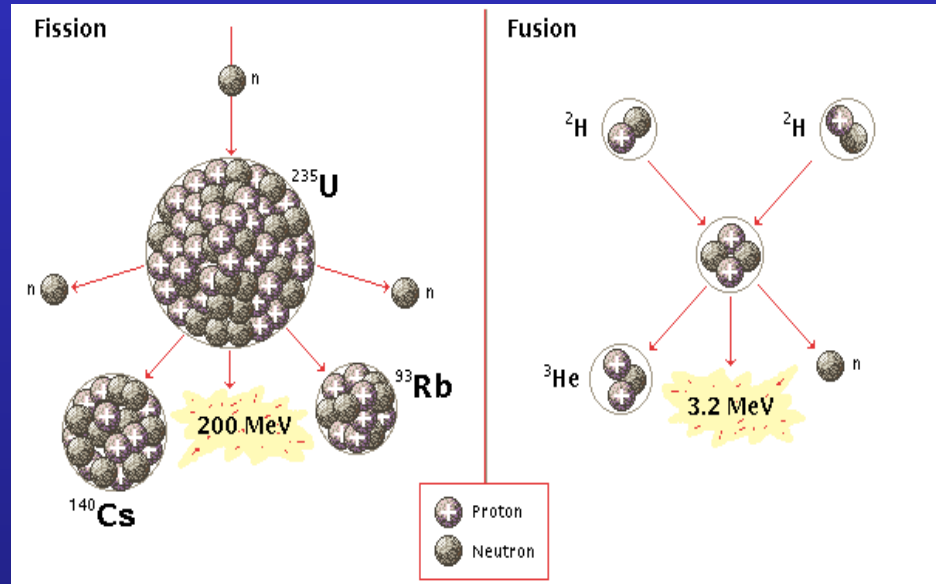
- ✦ combining of two nuclei to form one nucleus of larger mass
- ✦ thermonuclear reaction – requires temp of 40,000,000 K to sustain
- ✦ 1 g of fusion fuel = 20 tons of coal
- ✦ occurs naturally in stars



# Fission vs. Fusion

Courtesy Christy Johannesson [www.nisd.net/communicationsarts/pages/chem](http://www.nisd.net/communicationsarts/pages/chem)

**F  
I  
S  
S  
I  
O  
N**



**F  
U  
S  
I  
O  
N**

- ✦  $^{235}\text{U}$  is limited
- ✦ danger of meltdown
- ✦ toxic waste
- ✦ thermal pollution
- ✦ fuel is abundant
- ✦ no danger of meltdown
- ✦ no toxic waste
- ✦ not yet sustainable



# CHAPTER 22

---

# Nuclear Chemistry

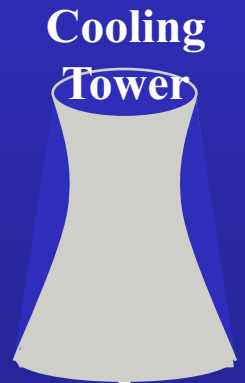
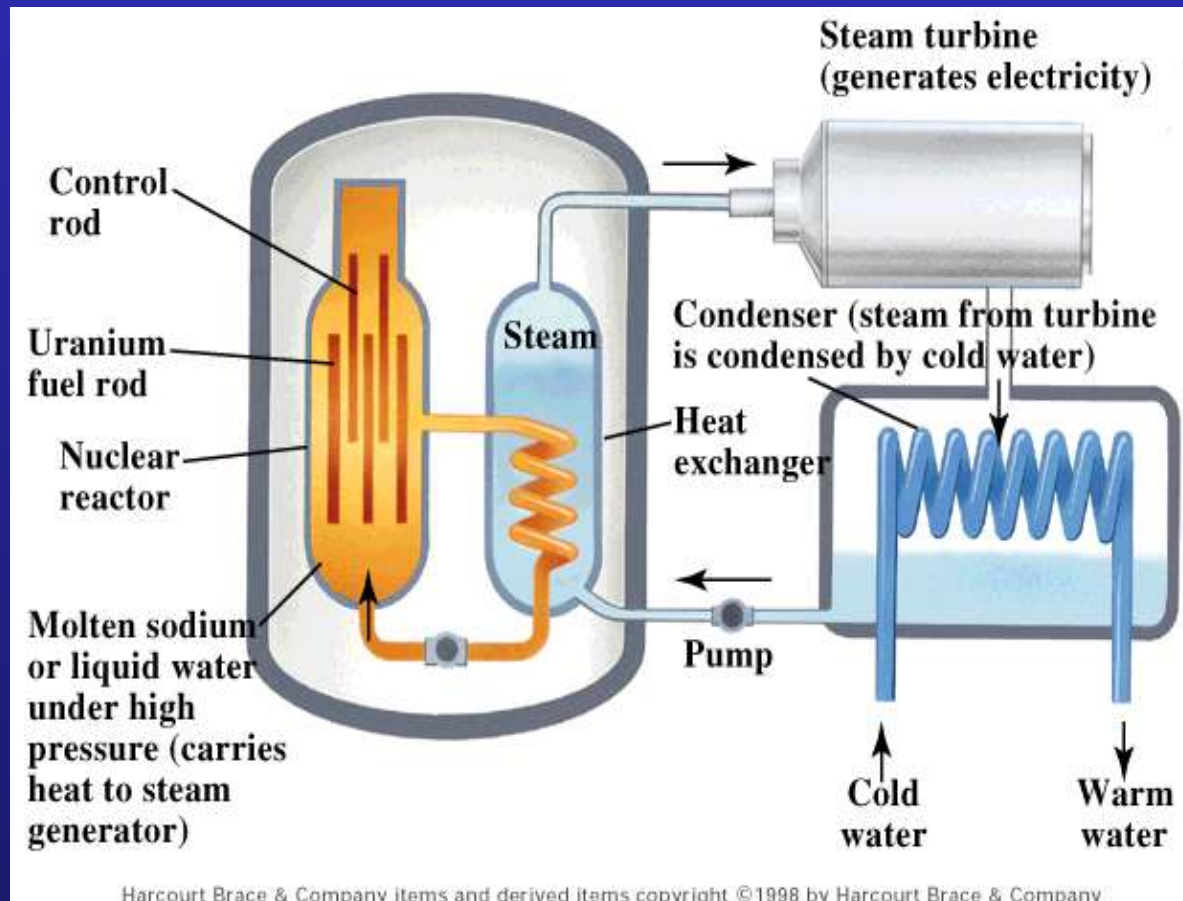


## **IV. Applications**

(p. 713 - 716)

# Nuclear Power

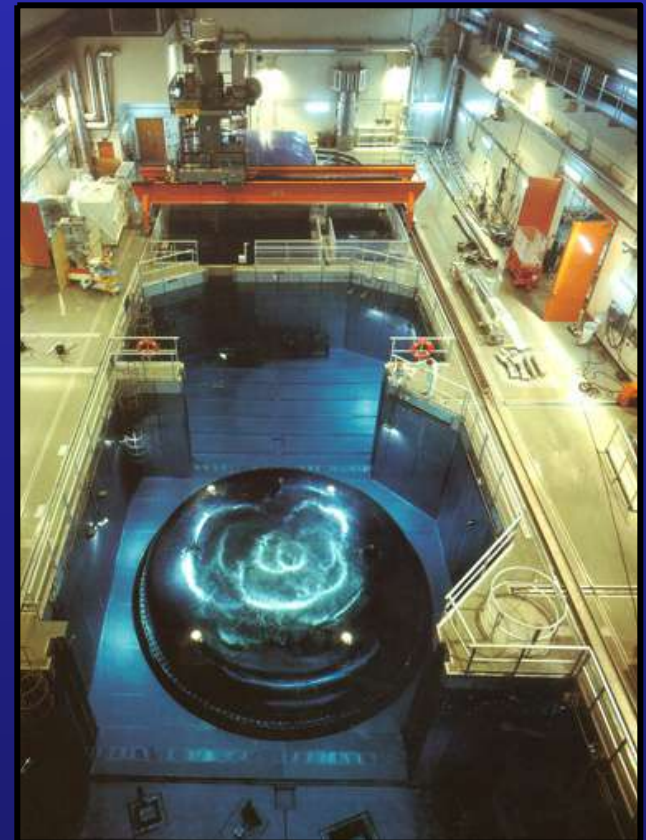
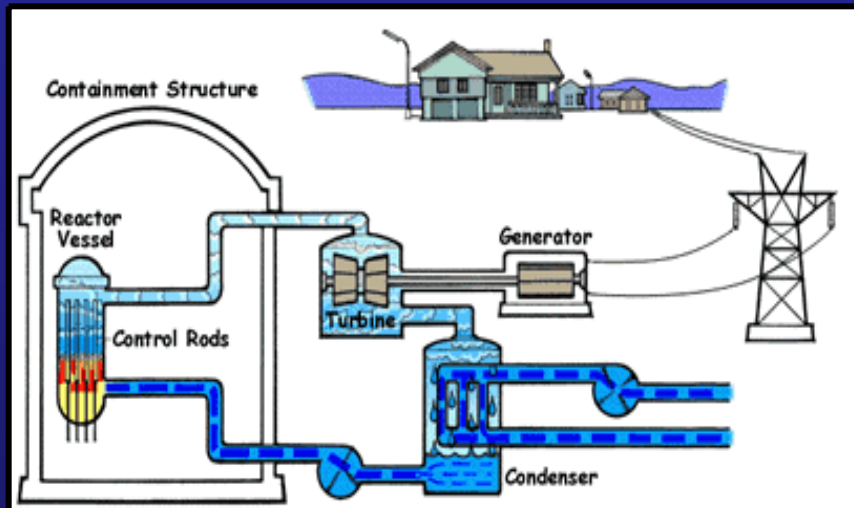
## ★ Fission Reactors





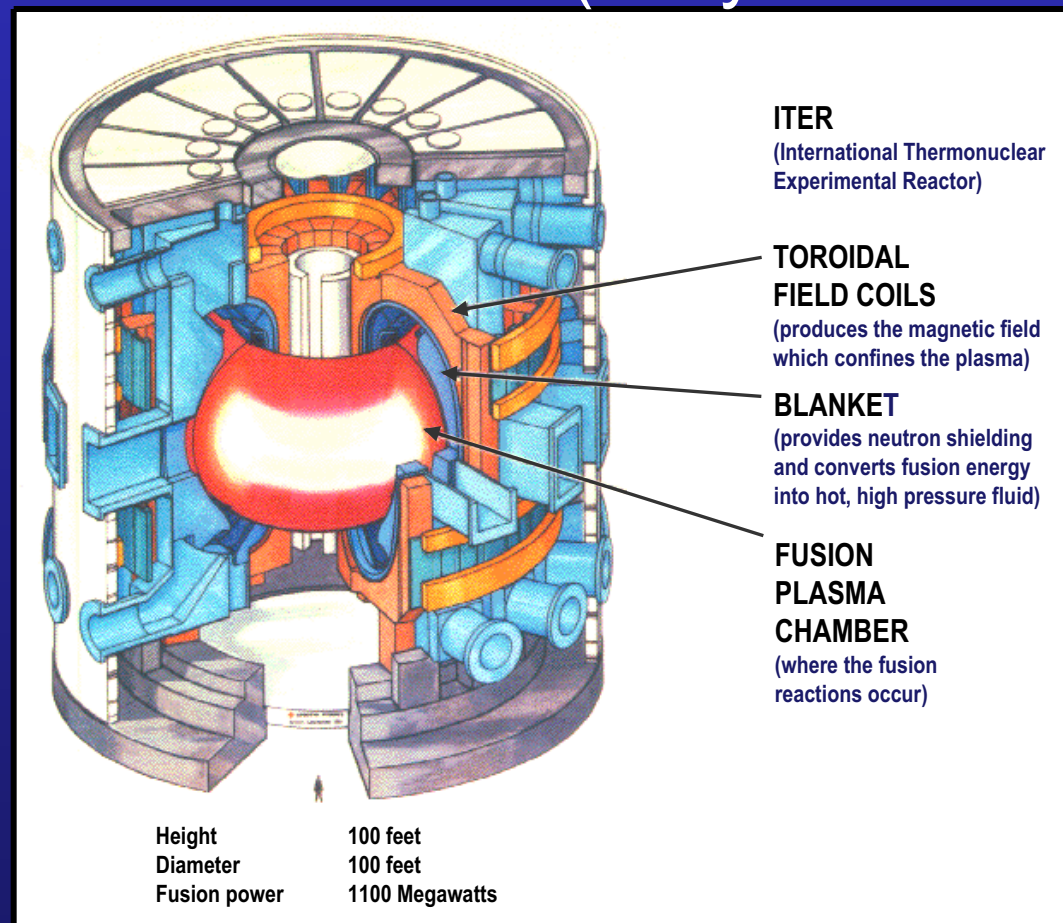
# Nuclear Power

## ★ Fission Reactors



# Nuclear Power

## ★ Fusion Reactors (not yet sustainable)



# Nuclear Power

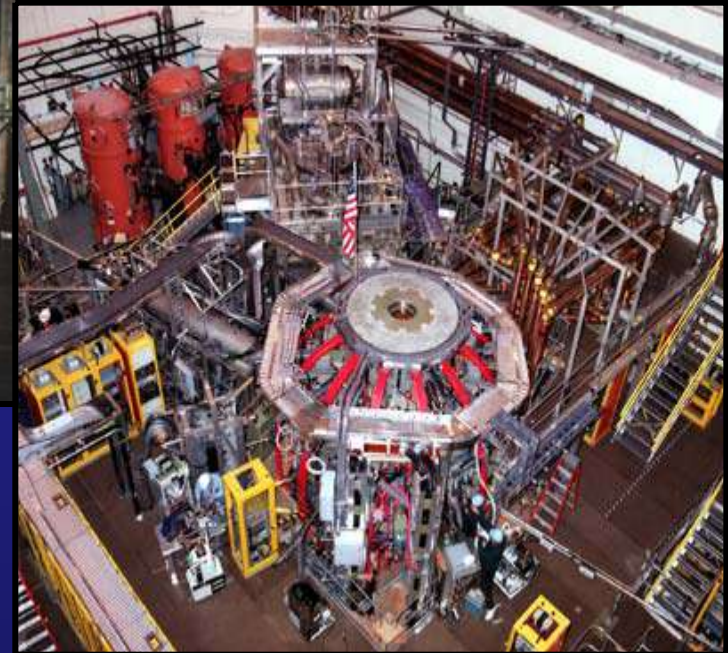
## ★ Fusion Reactors (not yet sustainable)



Tokamak Fusion Test Reactor

Princeton University

National Spherical  
Torus Experiment



# Synthetic Elements

---

## ★ Transuranium Elements

- ◆ elements with atomic #s above 92
- ◆ synthetically produced in nuclear reactors and accelerators
- ◆ most decay very rapidly





# Natural and artificial radioactivity

## Natural radioactivity

Isotopes that have been here since the earth formed.

Example - Uranium

Produced by cosmic rays from the sun.

Example – carbon-14

## Man-made Radioisotopes

Made in nuclear reactors when we split atoms (fission).

Produced using cyclotrons, linear accelerators,...

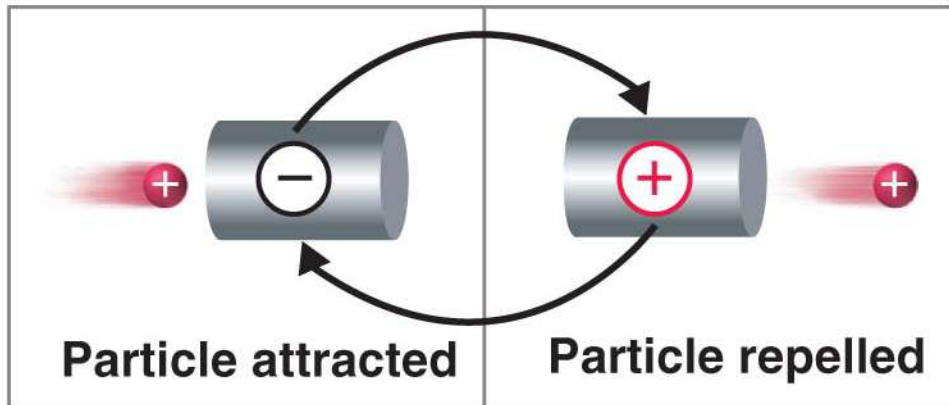
**Positive  
particle  
source**

**Alternating  
voltage**

**Particle  
beam**

**Vacuum**

**Target**



# Radioactive Dating

- ★ half-life measurements of radioactive elements are used to determine the age of an object
- ★ decay rate indicates amount of radioactive material
- ★ EX:  $^{14}\text{C}$  - up to 40,000 years  
 $^{238}\text{U}$  and  $^{40}\text{K}$  - over 300,000 years



# Nuclear Medicine

## ★ Radioisotope Tracers

- ◆ absorbed by specific organs and used to diagnose diseases

## ★ Radiation Treatment

- ◆ larger doses are used to kill cancerous cells in targeted organs
- ◆ internal or external radiation source



Radiation treatment  
using  
 $\gamma$ -rays from cobalt-60.



# Nuclear Weapons

## ★ Atomic Bomb

- ◆ chemical explosion is used to form a critical mass of  $^{235}\text{U}$  or  $^{239}\text{Pu}$
- ◆ fission develops into an uncontrolled chain reaction

## ★ Hydrogen Bomb

- ◆ chemical explosion → fission → fusion
- ◆ fusion increases the fission rate
- ◆ more powerful than the atomic



# Others

---

## ★ Food Irradiation

- ◆  $\gamma$  radiation is used to kill bacteria

## ★ Radioactive Tracers

- ◆ explore chemical pathways
- ◆ trace water flow
- ◆ study plant growth, photosynthesis

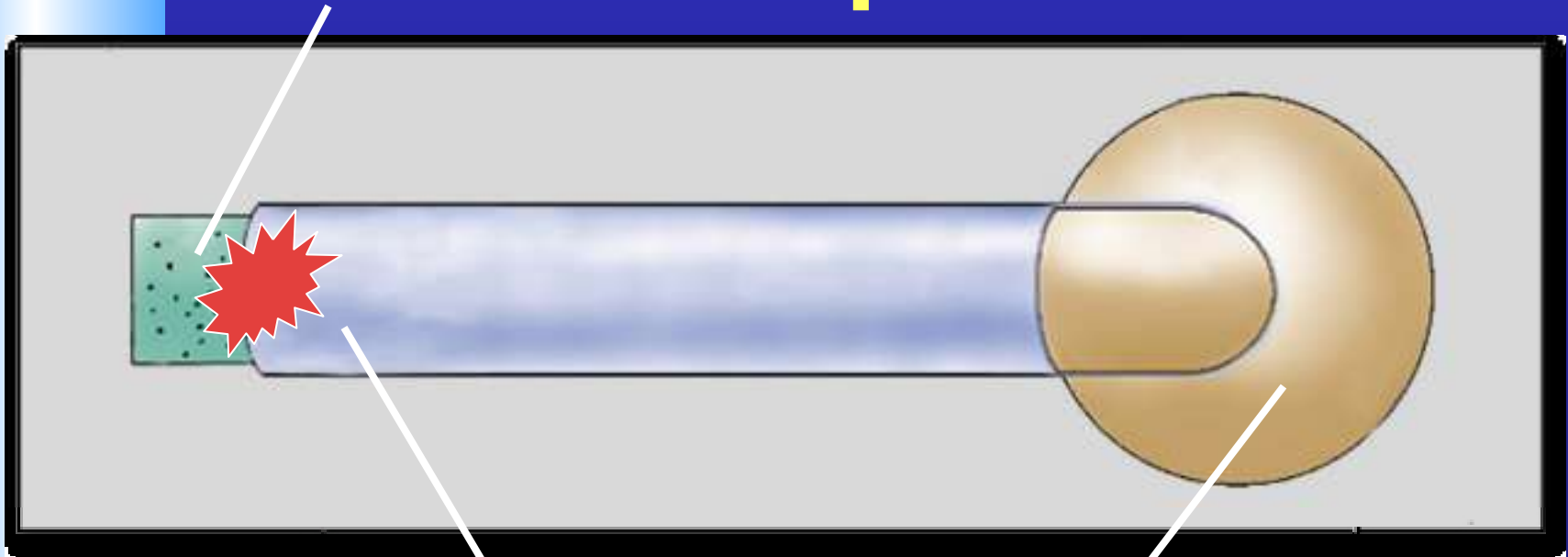
## ★ Consumer Products

- ◆ ionizing smoke detectors -  $^{241}\text{Am}$



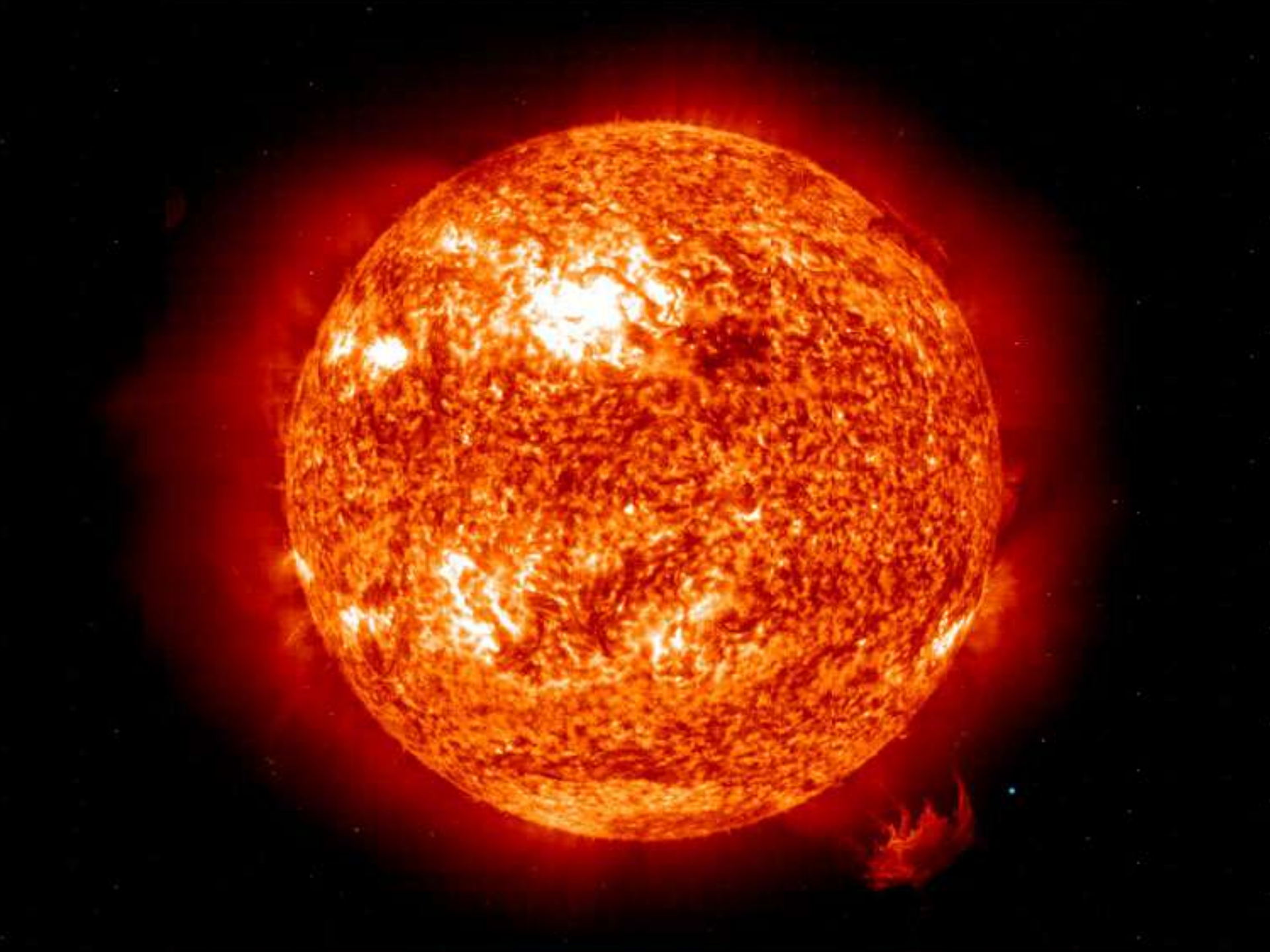
# Simplified diagram of fission bomb

**Chemical Explosive**

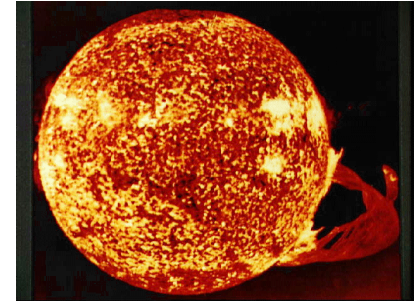


**Subcritical  
masses**

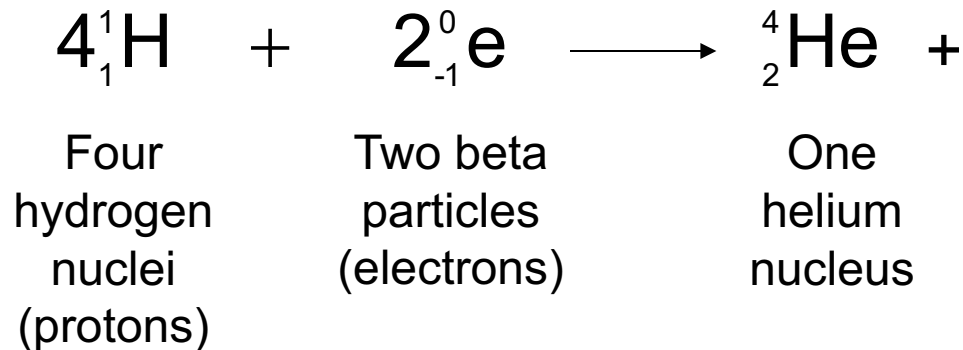
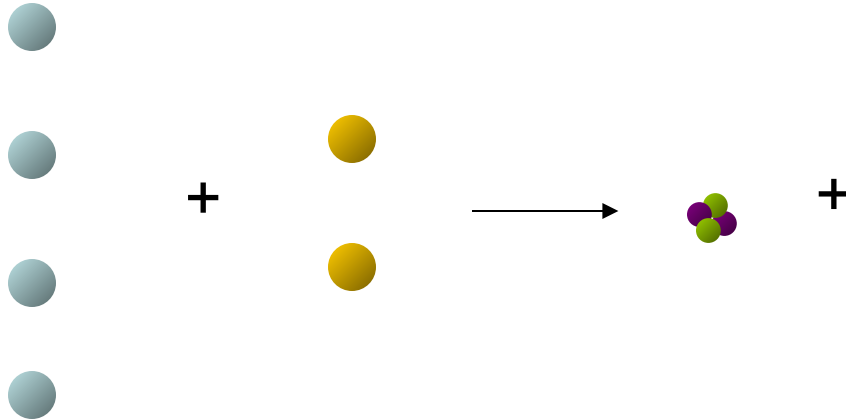
**Critical  
mass**



# Nuclear Fusion



Sun



Energy



# Conservation of Mass

*...mass is converted into energy*

Hydrogen (H<sub>2</sub>)

H = 1.008 amu

Helium (He)

He = 4.004 amu

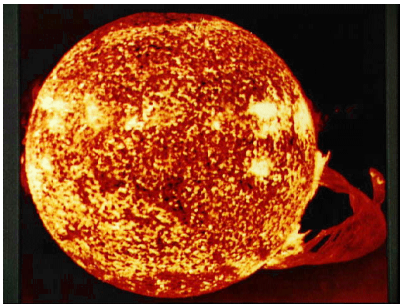
## ***FUSION***



1.008 amu

x 4

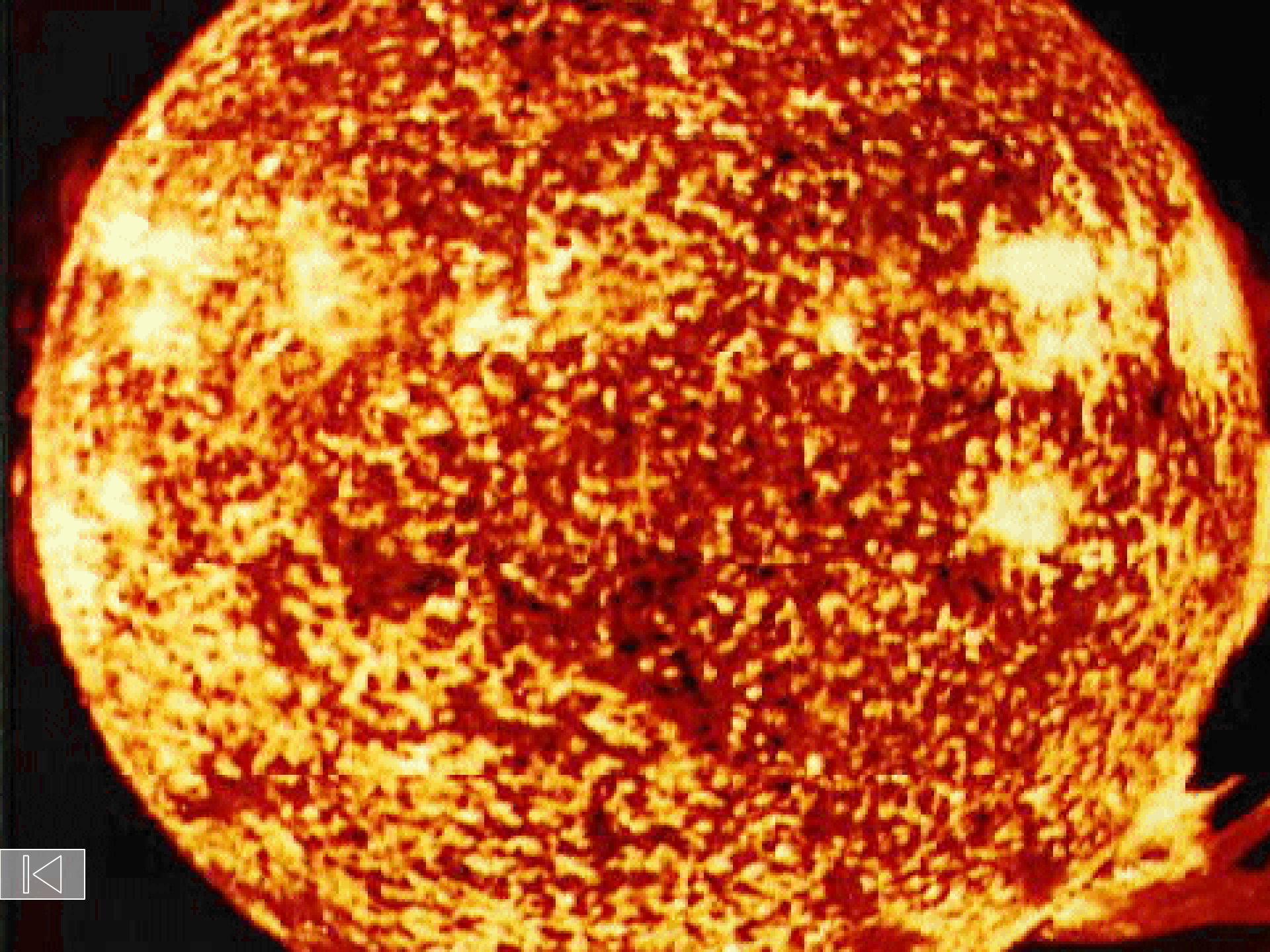
4.0032 amu = 4.004 amu + 0.028 amu



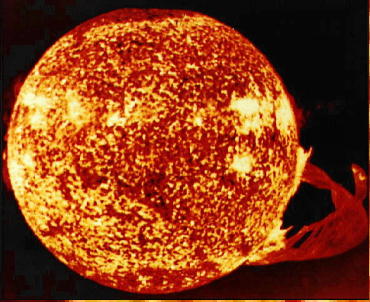
This relationship was discovered by Albert Einstein

$$E = mc^2$$

Energy = (mass) (speed of light)<sup>2</sup>

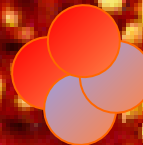
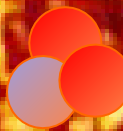






# Nuclear Fusion

(Positron)

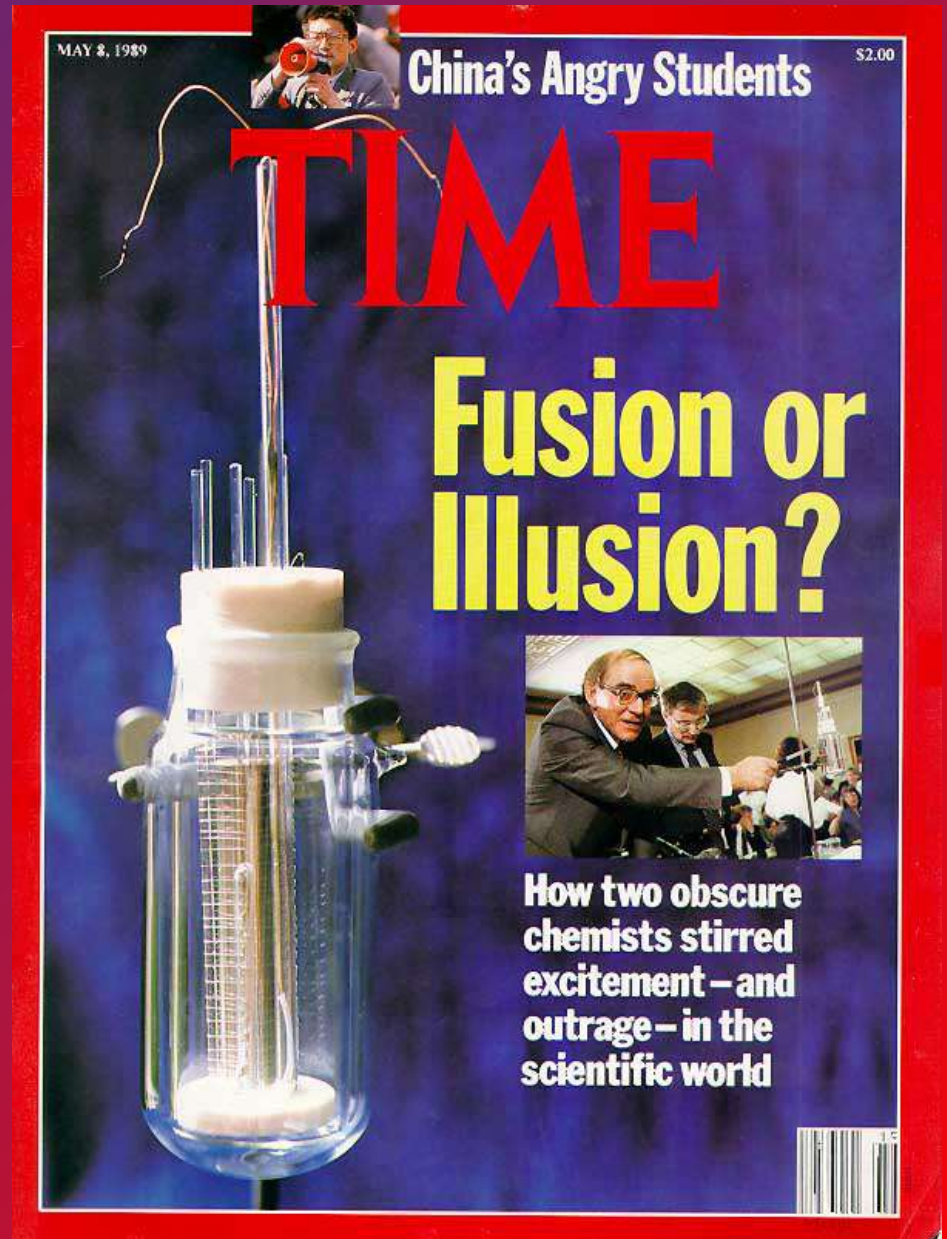




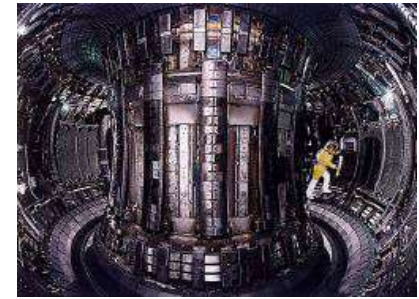
# Cold Fusion

- Fraud?
- Experiments must be repeatable to be valid

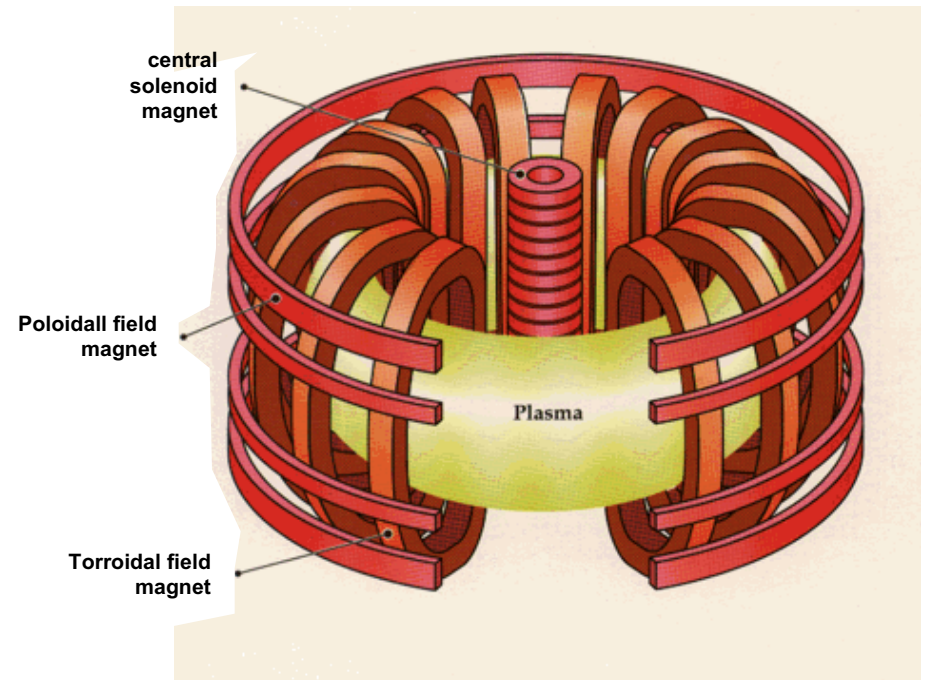
Stanley Pons and  
Martin Fleischman



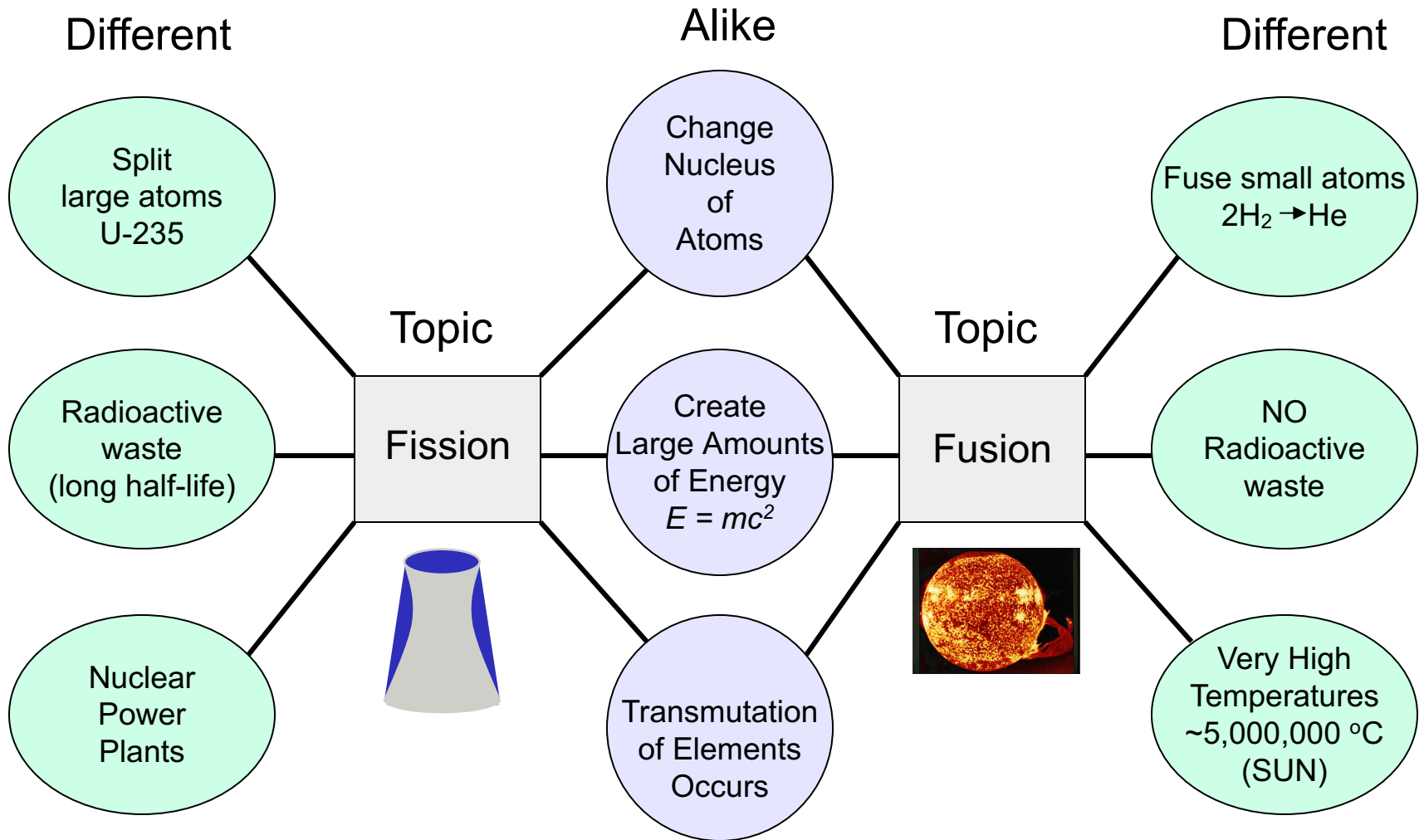
# Tokamak Reactor



- Fusion reactor
- 10,000,000 ° Celsius
- Russian for torroidial (doughnut shaped) ring
- Magnetic field contains *plasma*



# Fission vs. Fusion



# Atomic Structure

- **ATOMS**

- Differ by number of *protons*

carbon vs. oxygen

6 protons

8 protons

- **IONS**

- Differ by number of *electrons*

C

6 e<sup>-</sup>  
6 p<sup>+</sup>

C<sup>4+</sup>

2 e<sup>-</sup>  
6 p<sup>+</sup>

C<sup>4-</sup>

10 e<sup>-</sup>  
6 p<sup>+</sup>

- **ISOTOPES**

- Differ by number of *neutrons*

C-12

vs.

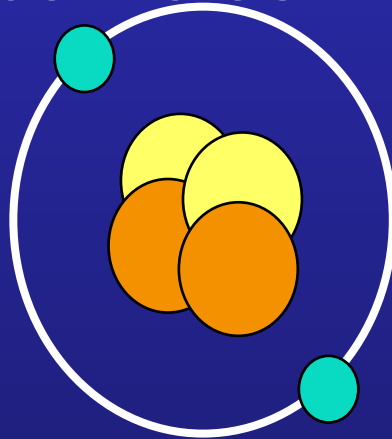
C-14

6 e<sup>-</sup>  
6 p<sup>+</sup>  
6 n<sup>0</sup>

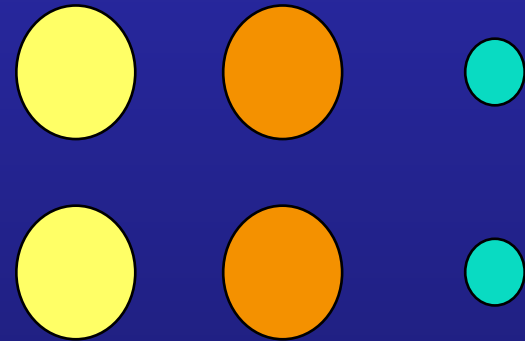
6 e<sup>-</sup>  
6 p<sup>+</sup>  
8 n<sup>0</sup>

# Mass Defect

- ★ Difference between the mass of an atom and the mass of its individual particles.



4.00260 amu



4.03298 amu

# Nuclear Binding Energy

- ★ Energy released when a nucleus is formed from nucleons.
- ★ High binding energy = stable nucleus.

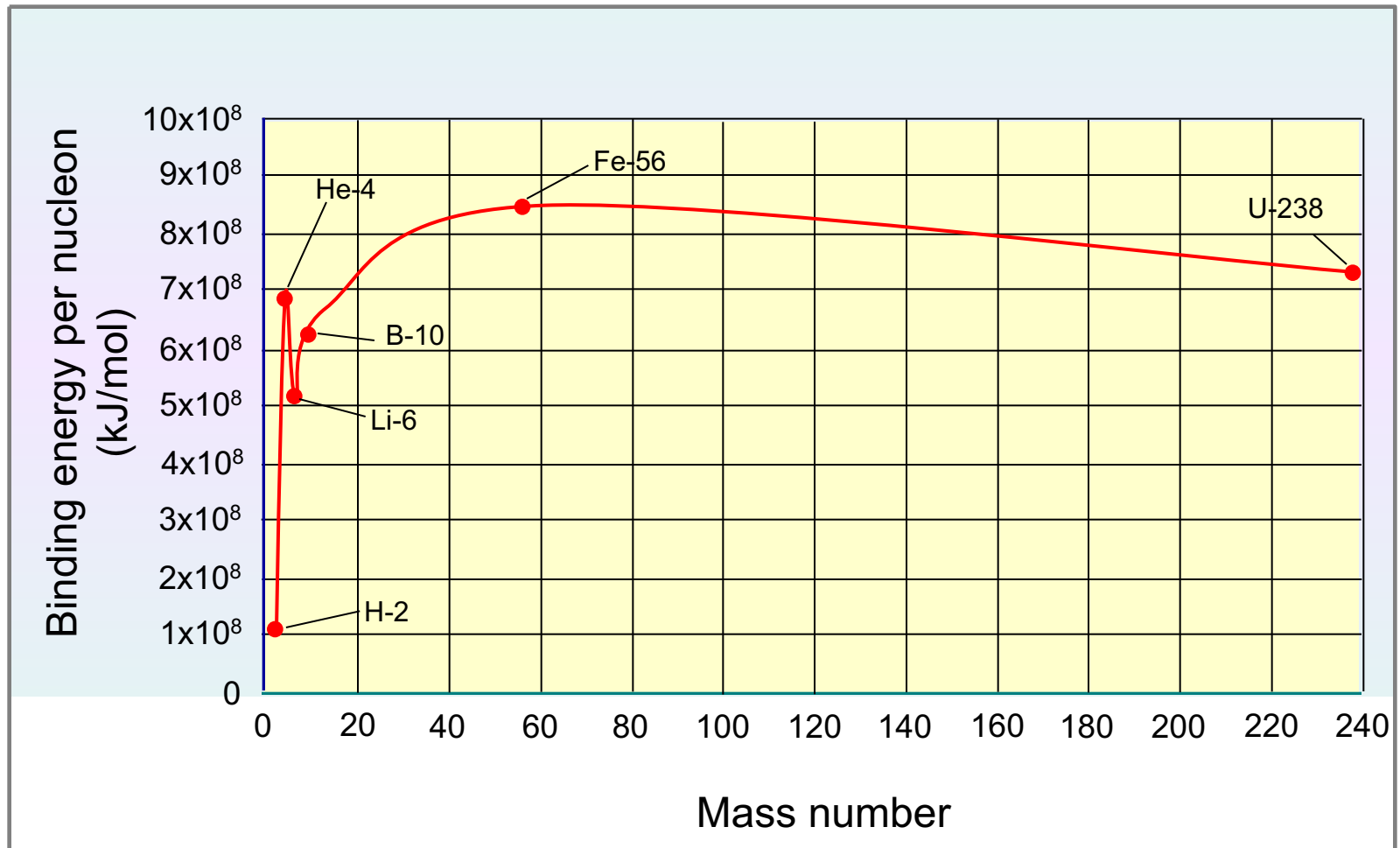

$$E = mc^2$$

E: energy (J)

m: mass defect (kg)

c: speed of light  
( $3.00 \times 10^8$  m/s)

# Nuclear Binding Energy



Unstable nuclides are radioactive and undergo radioactive decay.

# Mass Defect and Nuclear Stability

2 protons:  $(2 \times 1.007276 \text{ amu}) = 2.014552 \text{ amu}$

2 neutrons:  $(2 \times 1.008665 \text{ amu}) = 2.017330 \text{ amu}$

2 electrons:  $(2 \times 0.0005486 \text{ amu}) = 0.001097 \text{ amu}$

---

Total combined mass:  $4.032979 \text{ amu} \neq 4.002602 \text{ amu}$

The atomic mass of He atom is  $4.002602 \text{ amu}$ .

This is  $0.030368 \text{ amu}$  less than the combined mass.

This difference between the mass of an atom and the sum of the masses of its protons, neutrons, and electrons is called the mass defect.



# Nuclear Binding Energy

What causes the loss in mass?

According to Einstein's equation  $E = mc^2$



Convert mass defect to energy units

$$0.030368 \text{ amu} \left( \frac{1.6605 \times 10^{-27} \text{ kg}}{1 \text{ amu}} \right) = 5.0426 \times 10^{-29} \text{ kg}$$

The energy equivalent can now be calculated

$$E = m c^2$$

$$E = (5.0426 \times 10^{-29} \text{ kg}) (3.00 \times 10^8 \text{ m/s})^2$$

$$E = (4.54 \times 10^{-12} \text{ kg m}^2/\text{s}^2) = 4.54 \times 10^{-12} \text{ J}$$

This is the NUCLEAR BINDING ENERGY, the energy released when a nucleus is formed from nucleons.

# Binding Energy per Nucleon

## 1) Calculate mass defect

protons: 1.007276 amu

neutrons: 1.008665 amu

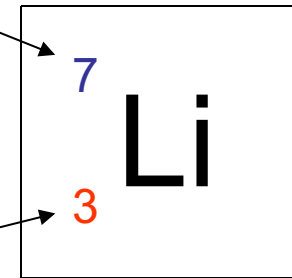
electrons: 0.0005486 amu

mass number

(# of protons  
+ neutrons)

atomic number

(# of protons)



Li - 7

## 2) Convert amu $\longrightarrow$ kg

$$\text{_____ amu} \left( \frac{1.6605 \times 10^{-27} \text{ kg}}{1 \text{ amu}} \right) = \text{_____ kg}$$

## 3) $E = mc^2$

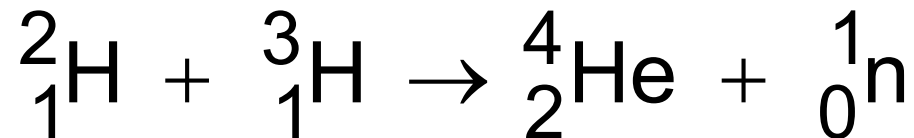
speed of light (c)  $3.00 \times 10^8 \text{ m/s}$

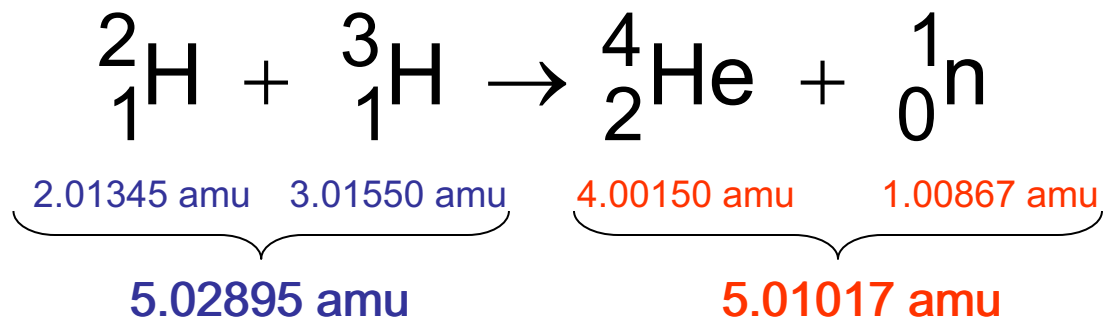
## 4) Divide binding energy by number of nucleons

# The Energy of Fusion

The fusion reaction releases an enormous amount of energy relative to the mass of the nuclei that are joined in the reaction. Such an enormous amount of energy is released because some of the mass of the original nuclei is converted to energy. The amount of energy that is released by this conversion can be calculated using Einstein's relativity equation  $E = mc^2$ .

Suppose that, at some point in the future, controlled nuclear fusion becomes possible. You are a scientist experimenting with fusion and you want to determine the energy yield in joules produced by the fusion of one mole of deuterium (H-2) with one mole of tritium (H-3), as shown in the following equation:

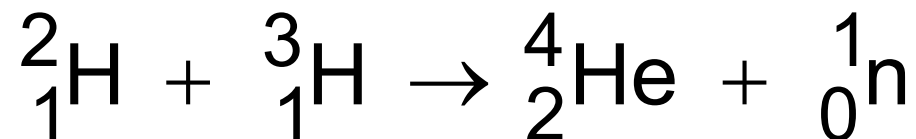




First, you must calculate the mass that is "lost" in the fusion reaction. The atomic masses of the reactants and products are as follows: deuterium (2.01345 amu), tritium (3.01550 amu), helium-4 (4.00150 amu), and a neutron (1.00867 amu).

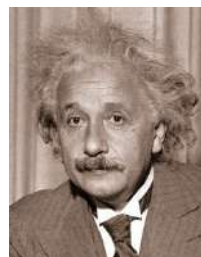
Mass defect:

$$\begin{array}{r}
 - \\
 \hline
 0.01878 \text{ amu}
 \end{array}$$



Mass defect = 0.01878 amu

According to Einstein's equation  $E = mc^2$



Convert mass defect to energy units

$$0.01878 \text{ amu} \left( \frac{1.6605 \times 10^{-27} \text{ kg}}{1 \text{ amu}} \right) = 3.1184 \times 10^{-29} \text{ kg}$$

The energy equivalent can now be calculated

$$E = m c^2$$

$$E = (3.1184 \times 10^{-29} \text{ kg}) (3.00 \times 10^8 \text{ m/s})^2$$

$$E = (2.81 \times 10^{-12} \text{ kg m}^2/\text{s}^2) = 2.81 \times 10^{-12} \text{ J}$$

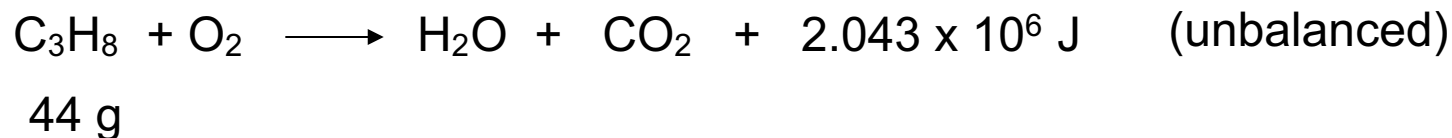
This is the NUCLEAR BINDING ENERGY, for the formation of a single Helium atom from a deuterium and tritium atom.

Therefore, one mole of helium formed by the fusion of one mole of deuterium and one mole of hydrogen would be  $6.02 \times 10^{23}$  times greater energy.

$$\begin{array}{r} 2.81 \times 10^{-12} \text{ J} \\ \times 6.02 \times 10^{23} \\ \hline 1.69 \times 10^{12} \text{ J of energy released per mole of helium formed} \end{array}$$

1,690,000,000,000 J

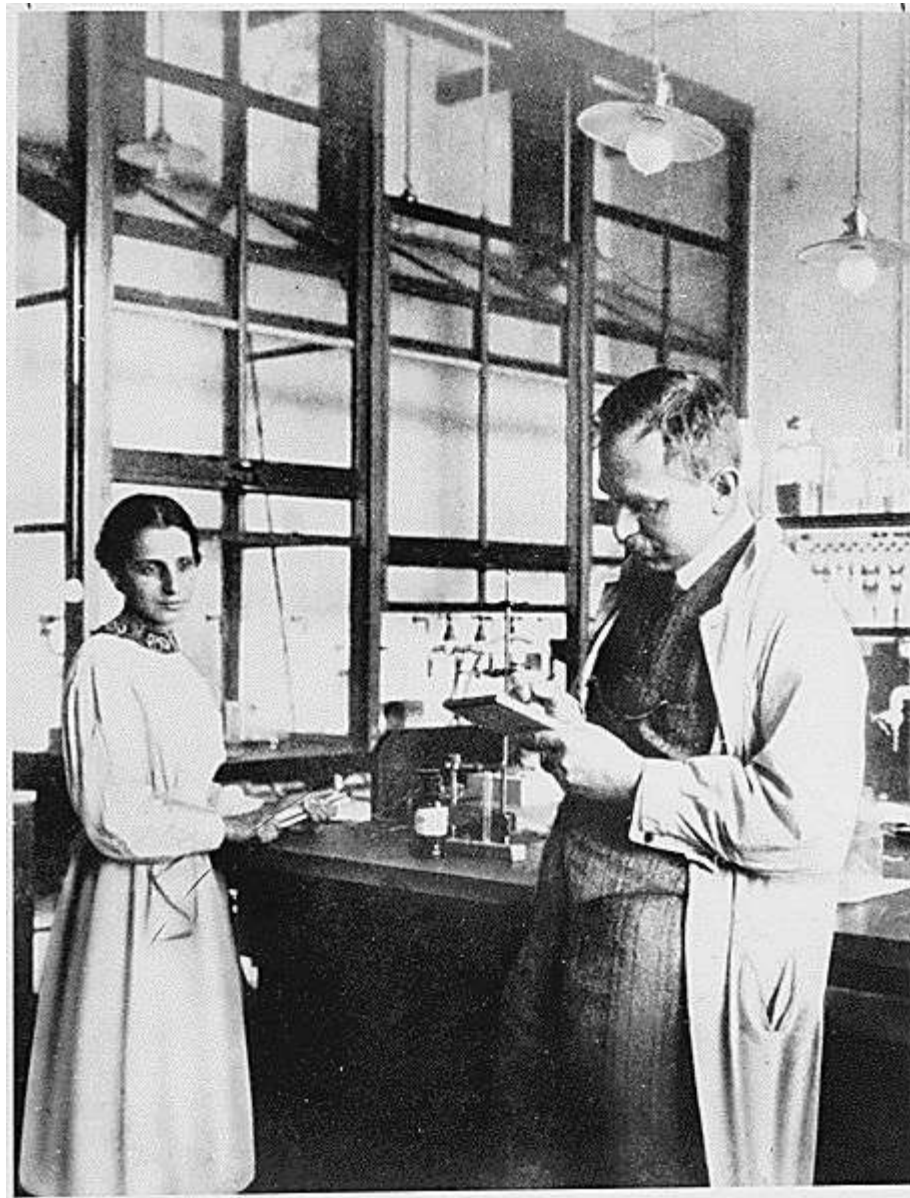
The combustion of one mole of propane ( $\text{C}_3\text{H}_8$ ), which has a mass of 44 g, releases  $2.043 \times 10^6$  J. How does this compare to the energy released by the fusion of deuterium and tritium, which you calculated?



4 g He	1,690,000,000,000 J
44 g $\text{C}_3\text{H}_8$	2,043,000 J

Fusion produces ~1,000,000 x  
more energy/mole

# Lise Meitner and Otto Hahn



# Atoms for Peace



Bombing of Japan in WW II

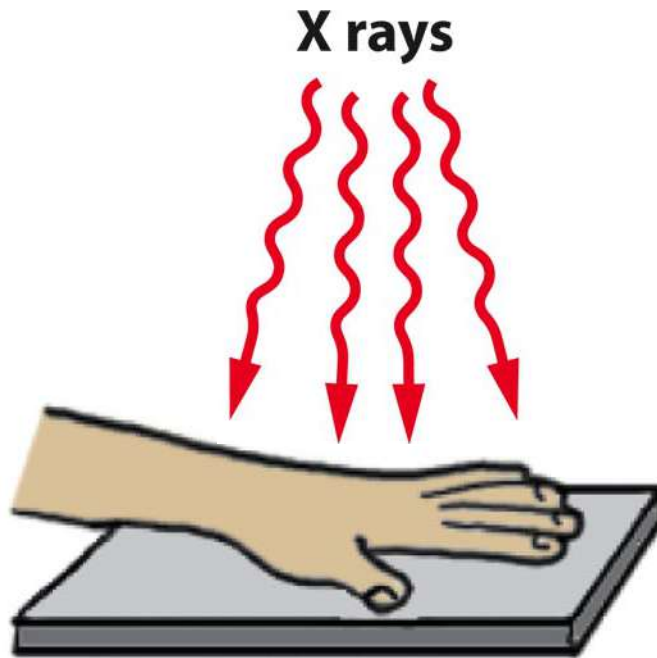
- Eisenhower
  - Show nuclear science is not evil
  - Has good uses, too.
- Food irradiation
- Cancer treatment
- PET & CAT scan
- Destroy ANTHRAX bacteria



Copyright © 2006 Pearson Benjamin Cummings. All rights reserved.



# Radiology



Photographic film enclosed  
in lightproof holder



Exposed and developed  
photographic film

# X-rays

Chest X-ray showing  
scoliosis corrected  
with steel rod



# Radioisotopes

- Radioactive isotopes
- Many uses
  - Medical diagnostics
  - Optimal composition of fertilizers
  - Abrasion studies in engines and tires



Radioisotope is injected into the bloodstream to observe circulation.

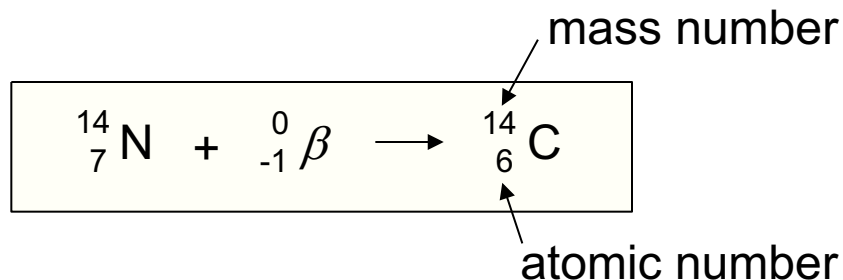
# Isotopes of Three Common Elements

Element	Symbol	Mass Number	Mass (amu)	Fractional Abundance	Average Atomic Mass
Carbon	$^{12}_6\text{C}$	12	12 (exactly)	99.89%	12.01
	$^{13}_6\text{C}$	13	13.003	1.11%	
Chlorine	$^{35}_{17}\text{Cl}$	35	34.969	75.53%	35.45
	$^{37}_{17}\text{Cl}$	37	36.966	24.47%	
Silicon	$^{28}_{14}\text{Si}$	28	27.977	92.21%	28.09
	$^{29}_{14}\text{Si}$	29	28.976	4.70%	
	$^{30}_{14}\text{Si}$	30	29.974	3.09%	

# Radioactivity and Nuclear Energy

## Practice Quiz

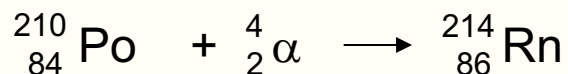
1. Which of the following is *not* an example of spontaneous radioactive process?
  - a. alpha-decay
  - b. beta-decay
  - c. positron production
  - ★ d. autoionization
  - e. electron capture
2. If a nucleus captures an electron, describe how the atomic number will change.
  - a. It will increase by one
  - ★ b. It will decrease by one
  - c. It will not change because the electron has such a small mass
  - d. It will increase by two
  - e. It will decrease by two



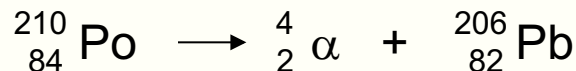
# Radioactivity and Nuclear Energy

3. Polonium is a naturally radioactive element decaying with the loss of an alpha particle.  ${}^{210}_{84}\text{Po} \rightarrow {}^4_2\text{He} + \underline{\text{?}}$ . What is the second product of this decay?

- a. Rn-214
- ★ b. Pb-206
- c. At-206
- d. Hg-208
- e. none of these



alpha *absorption*



alpha *emission*

4. Thorium-234 undergoes beta particle production. What is the other product?

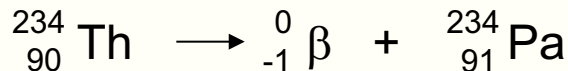
- ★ a.  ${}^{234}_{91}\text{Pa}$

- b.  ${}^{234}_{89}\text{Ac}$

- c.  ${}^{233}_{90}\text{Th}$

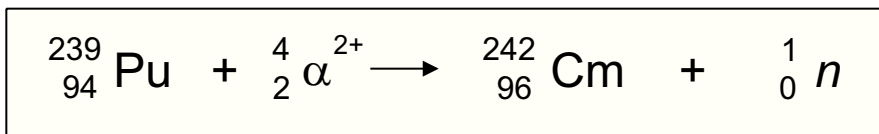
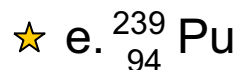
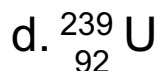
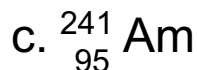
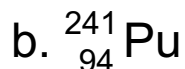
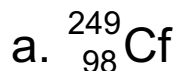
- d.  ${}^{233}_{91}\text{Th}$

- e. none of these

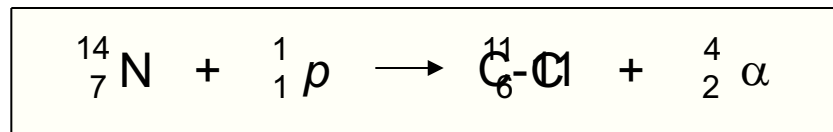


# Radioactivity and Nuclear Energy

5. The element curium ( $Z = 96$ ,  $A = 242$ ) can be produced by positive-ion bombardment when an alpha particle collides with which of the following nuclei? Recall that a neutron is also a product of this bombardment.



6. When  ${}_{7}^{14}\text{N}$  is bombarded by (and absorbs) a proton, a new nuclide is produced plus an alpha particle. The nuclide produced is \_\_\_\_\_?

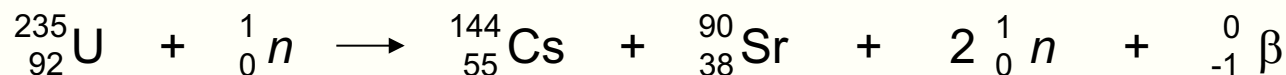


# Radioactivity and Nuclear Energy

7. When the uranium-235 nucleus is struck with a neutron, the cesium-144 and strontium-90 nuclei are produced with some neutrons and electrons.  
a) How many neutrons are produced? b) How many electrons are produced?

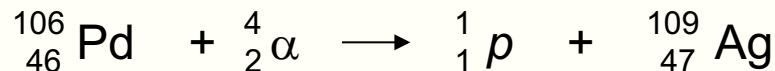
- ★ a. 2  
b. 3  
c. 4  
d. 5  
e. 6

- ★ a. 1  
b. 2  
c. 3  
d. 4  
e. 5



8. When the palladium-106 nucleus is struck with an alpha particle, a proton is produced along with a new element. What is the new element?

- a. cadmium-112  
b. cadmium-109  
c. silver-108  
★ d. silver-109  
e. none of these





# Radioactivity and Nuclear Energy

9. Strontium-90 from radioactive fallout is a health threat because, like \_\_\_\_\_, it is incorporated into bone.

- a. iodine
- b. cesium
- c. iron
- ★ d. calcium
- e. uranium

*Strontium (Sr) and calcium (Ca) are alkaline earth metals. Strontium is chemically more reactive than calcium.*

10. Nuclear fusion uses heavy nuclides such as  ${}_{92}^{235}\text{U}$  as fuel. **True / False**

*FALSE,*

*Nuclear fission splits heavy nuclides such as U-235 for fuel in nuclear reactors. Nuclear fusion joins light nuclides such as H-1 into He-4 (on the Sun).*

# Textbook Problems

## Modern Chemistry

### Chapter 22

Pg 704 #1-4 Section Review

Pg 712 #1-5 Section Review

Pg 715 #1-4

Pg 719 #1-4

End of Chapter #25-47 (pg 723-724)

25. The mass of a Ne-20 atom is 19.99244 amu.  
Calculate its mass defect.

26. The mass of Li-7 is 7.01600 amu.  
Calculate its mass defect.

27. Calculate the nuclear binding energy of one lithium-6 atom.  
The measured atomic mass of lithium-6 is 6.015 amu.